

# Direct methods for the inversion of limb scattering measurements by machine learning techniques.

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# Outline

- Context and objectives
  - Inverse problems
    - Limb geometry
- Test bench with OMPS
  - Conclusions

# Context and objectives

- **ALTIUS level 2 processor** is well advanced for ozone retrievals.
  - it is based on standard methods
  - heavy hardware constraints : f.i. no GPU allowed -> **no “direct chain” possible with our Radiative Transfer Model (RTM)**
  - proxies and look-up tables and “L-M” algos
  - many instrumental side effects are possible and difficult to manage. straylight, convolutions,..etc
- Objective: **to explore “direct methods”** = combining two powerful numerical weapons
  - use of orthogonal function bases given by Principal Component Analysis (PCA)
  - nonlinear regression by Machine Learning (ML)



# Inverse problems

- Probably the most frequent problems in experimental physics: the retrieved quantity results from one or several integrations of an unknown distribution
- Huge amount of references and methods: Bayesian optimal estimation, Philips-Twomey-Tikhonov regularization, constrained non-linear LS (L-M) for L2, L1,..norms, linear and log (Chahine) relaxation methods, Backus-Gilbert, Maximum Entropy Methods, ..etc

Inverse problem:  
measure “ $\mathbf{y}$ ”, then compute ‘ $\mathbf{x}$ ’

$$\mathbf{y} = \mathbf{F}(\mathbf{x}) + \epsilon$$

- In the Bayesian approach, there is a prior knowledge of  $\mathbf{x} = \mathbf{x}_a$  characterized by an associated covariance matrix  $\mathbf{S}_a$  that will combine with the measurement error covariance matrix  $\mathbf{S}_\epsilon$  to define the iterative update of the solution state vector as:

$$\mathbf{x}_{i+1} = \mathbf{x}_i + (\mathbf{S}_a^{-1} + \mathbf{K}_i^T \mathbf{S}_\epsilon^{-1} \mathbf{K}_i)^{-1} [\mathbf{K}_i^T \mathbf{S}_\epsilon^{-1} (\mathbf{y} - \mathbf{F}(\mathbf{x}_i)) - \mathbf{S}_a^{-1} (\mathbf{x}_i - \mathbf{x}_a)]$$

- In the least-squares Philips-Twomey-Tikhonov method (Twomey, 1977), the stabilization of the ill-posed problem is usually achieved by using a regularizing operator  $\mathbf{H}$  embedded in the Levenberg-Marquardt algorithm (Marquardt, 1963):

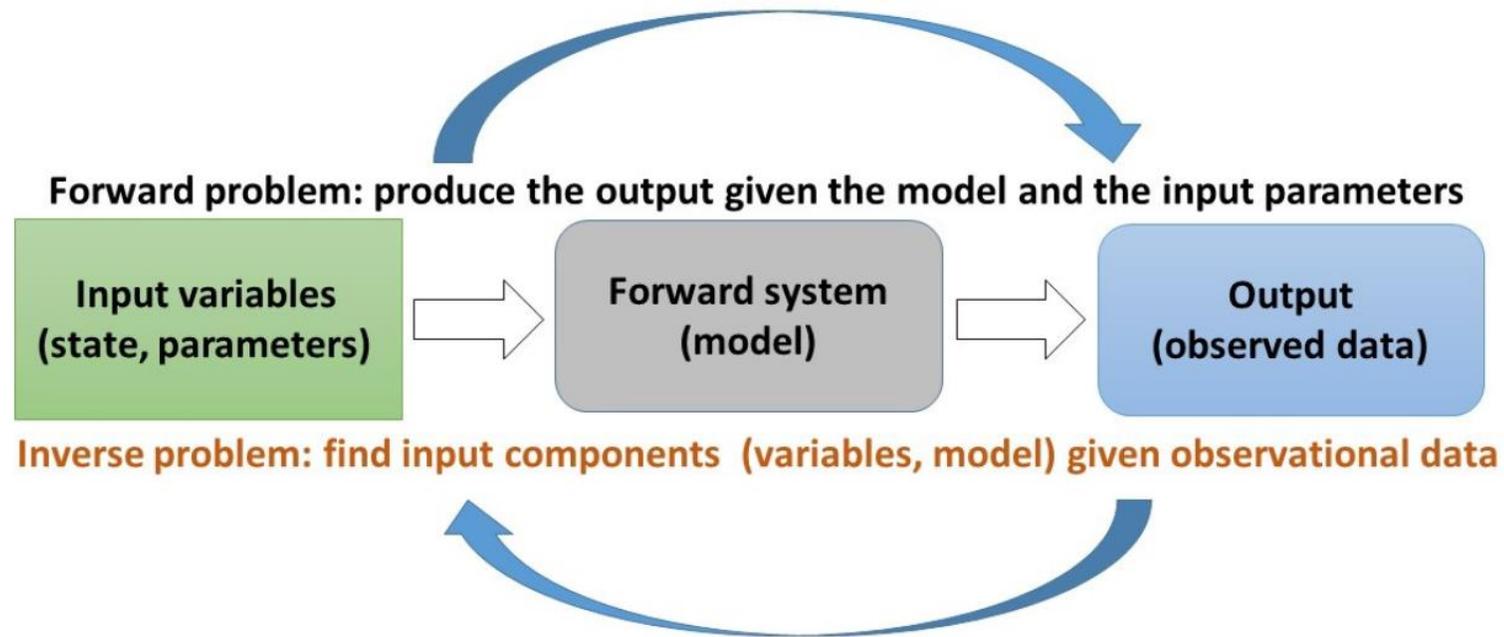
$$\mathbf{x}_{i+1} = \mathbf{x}_i + (\mathbf{K}_i^T \mathbf{K}_i + \gamma_i \mathbf{H})^{-1} \mathbf{K}_i^T (\mathbf{y} - \mathbf{F}(\mathbf{x}_i))$$

- Relaxation methods (Chahine, 1968; Twomey et al., 1977) are slow and need to be stopped when noise amplification starts to dominate but they do not require the computation of the Jacobian at every step:

$$\mathbf{x}_{i+1}^j = \mathbf{x}_i^j \frac{\mathbf{y}^j}{\mathbf{F}(\mathbf{x}_i)^j}$$

# Trying direct inverse methods...

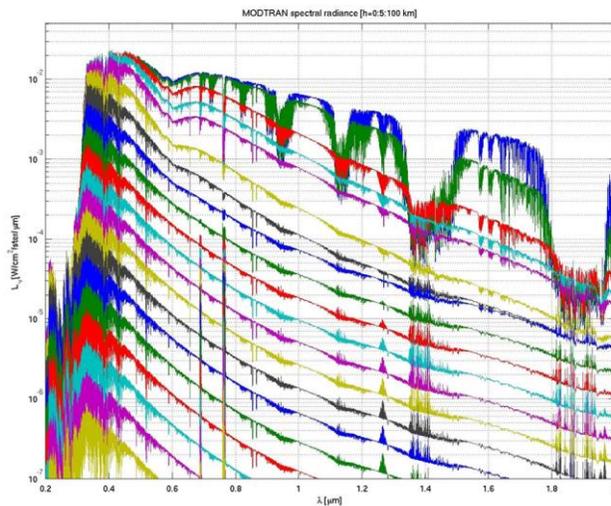
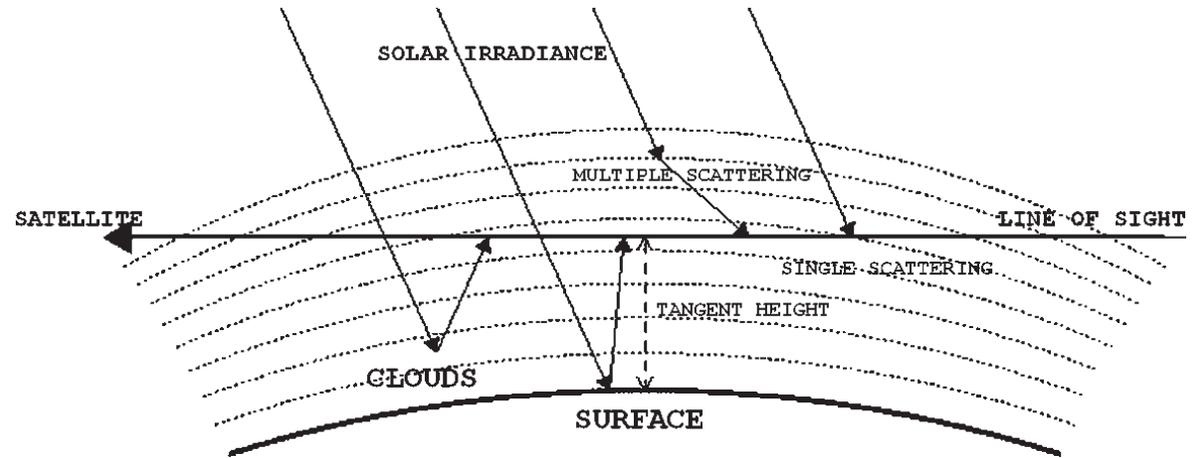
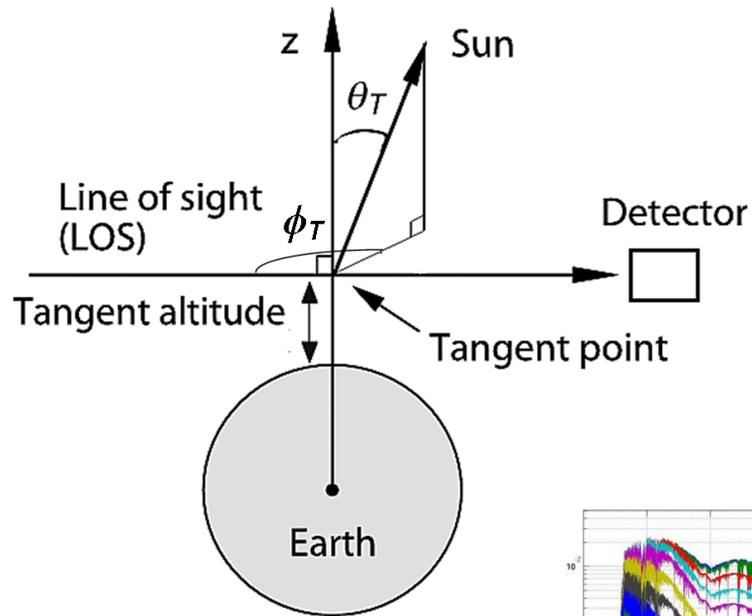
- WHEN a large number of observations has to be processed, the TOTAL computational cost may be considered. **How many calls to the forward model ?**
- Non-linearities in retrieval imply **iterative** schemes.
  - f.i. in ALTIUS, resolution is driving the number of forward model calls by the L-M algorithm.
  - even worse: all intermediate computations along the minimization path, including Jacobians, are lost.
- A **large “training” set (LTS)** is affordable at an equivalent computing load. Hereafter, “large” means 25 600 synthetic simulations by our Monte-Carlo radiative transfer code “SmartG”
  - > 2 weeks / 2 Tera photons shot / 0,3 % precision



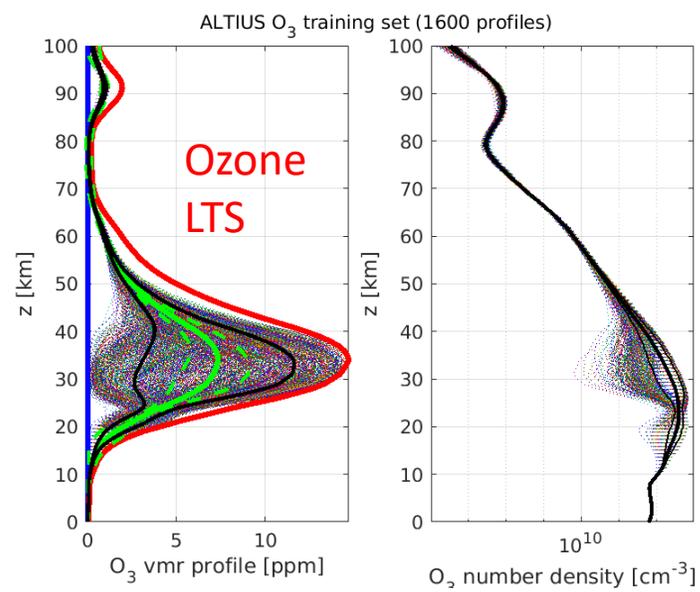
## Direct inversion ?

- Generate a **LTS**
- Just **swap** green and blue boxes.
- Replace the forward model by a **black box**.
- For LTS, force the **black box** to predict the **green box** from the **blue box**.

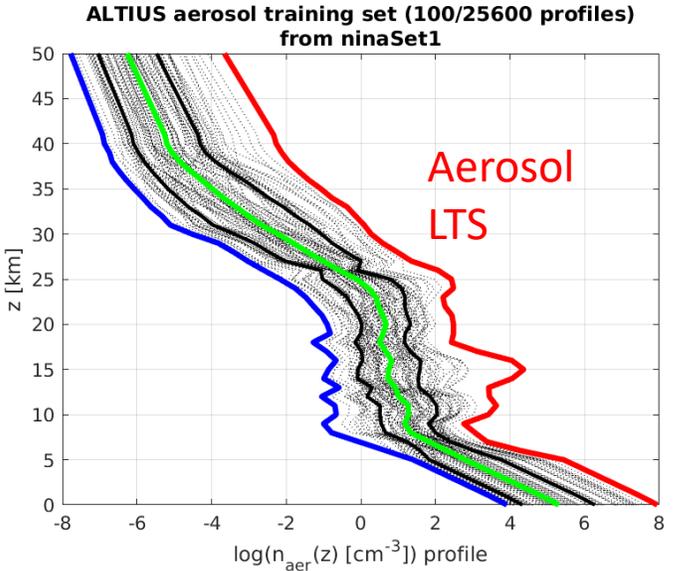
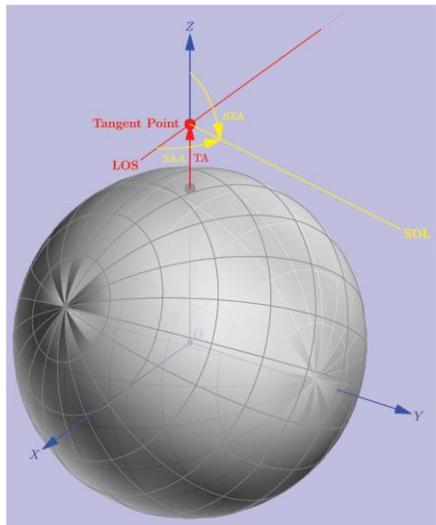
# Focus on solar limb scattering geometry



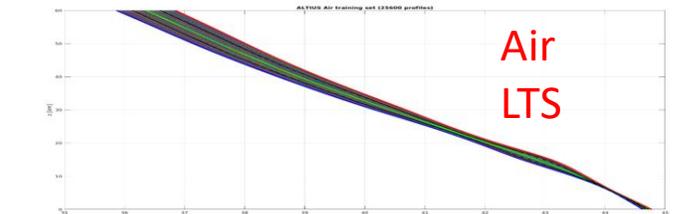
Simplified atmosphere:  
air + **ozone** + stratospheric  
aerosols + effective albedo +  
solar irradiance @  $z=0:1:100$   
km



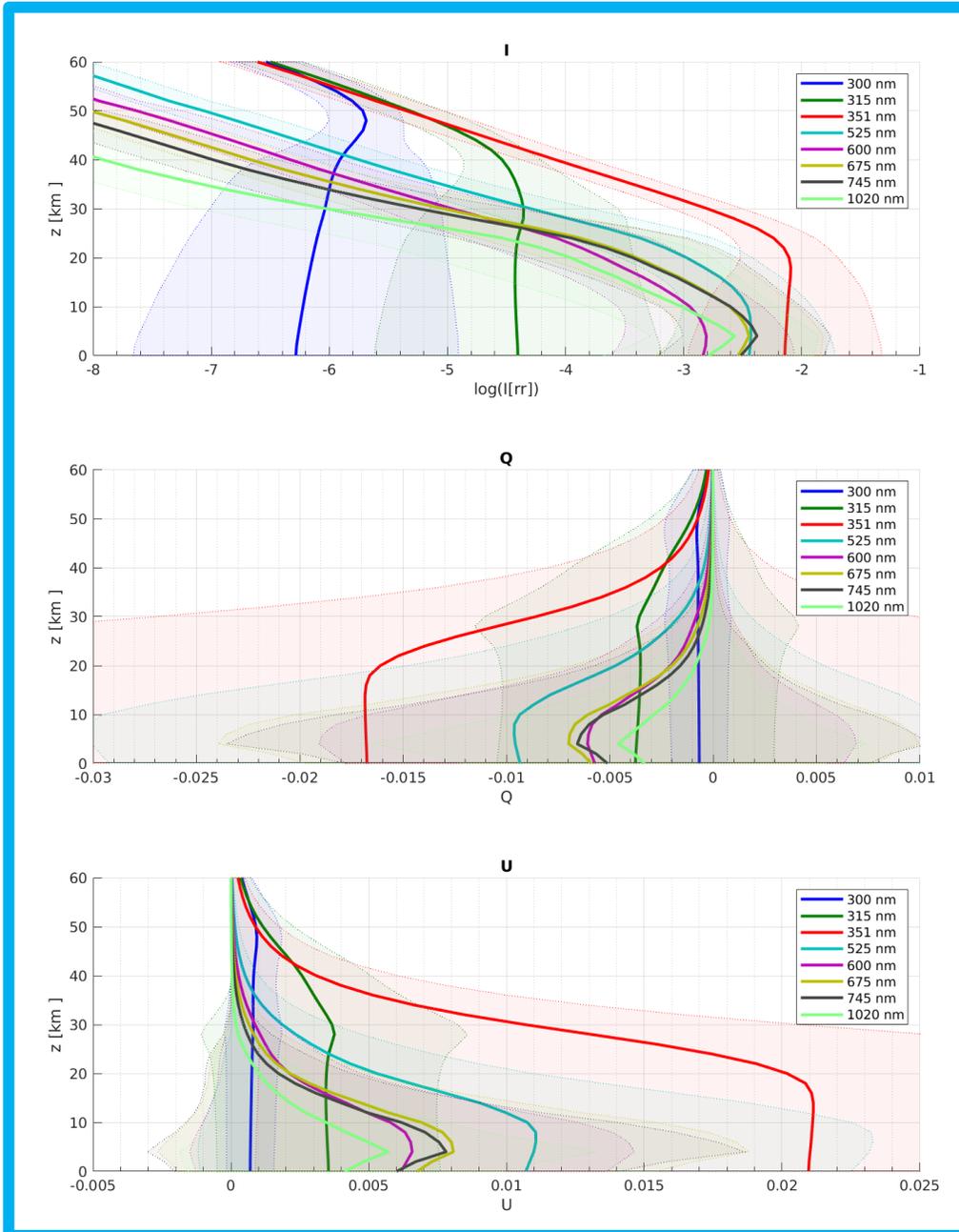
# LTS Factory



+ PSD, albedo  
+ Solar angles  
LTS

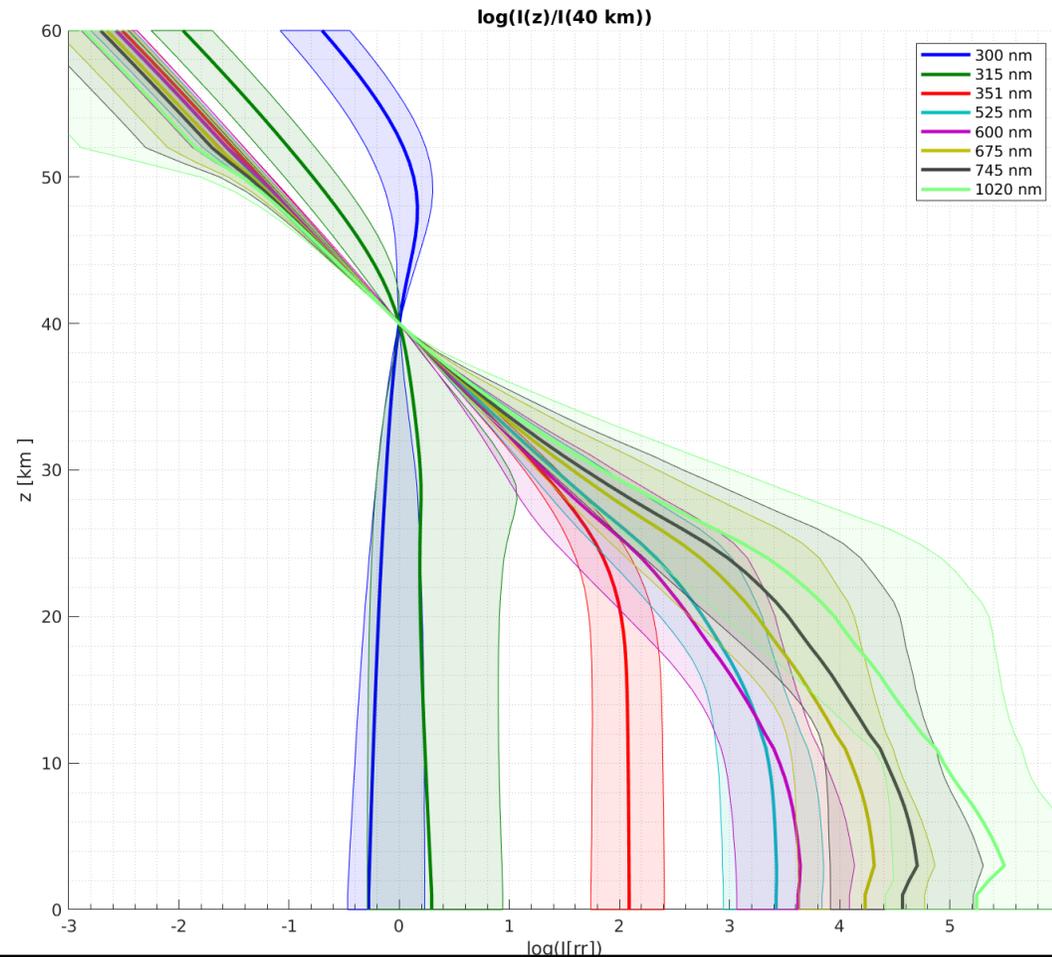


$\lambda = 300, 315, 351, 525, 600, 675, 745, 1020$  nm



Stokes vector profiles

# ONNI: Ozone Neural Network Inversion

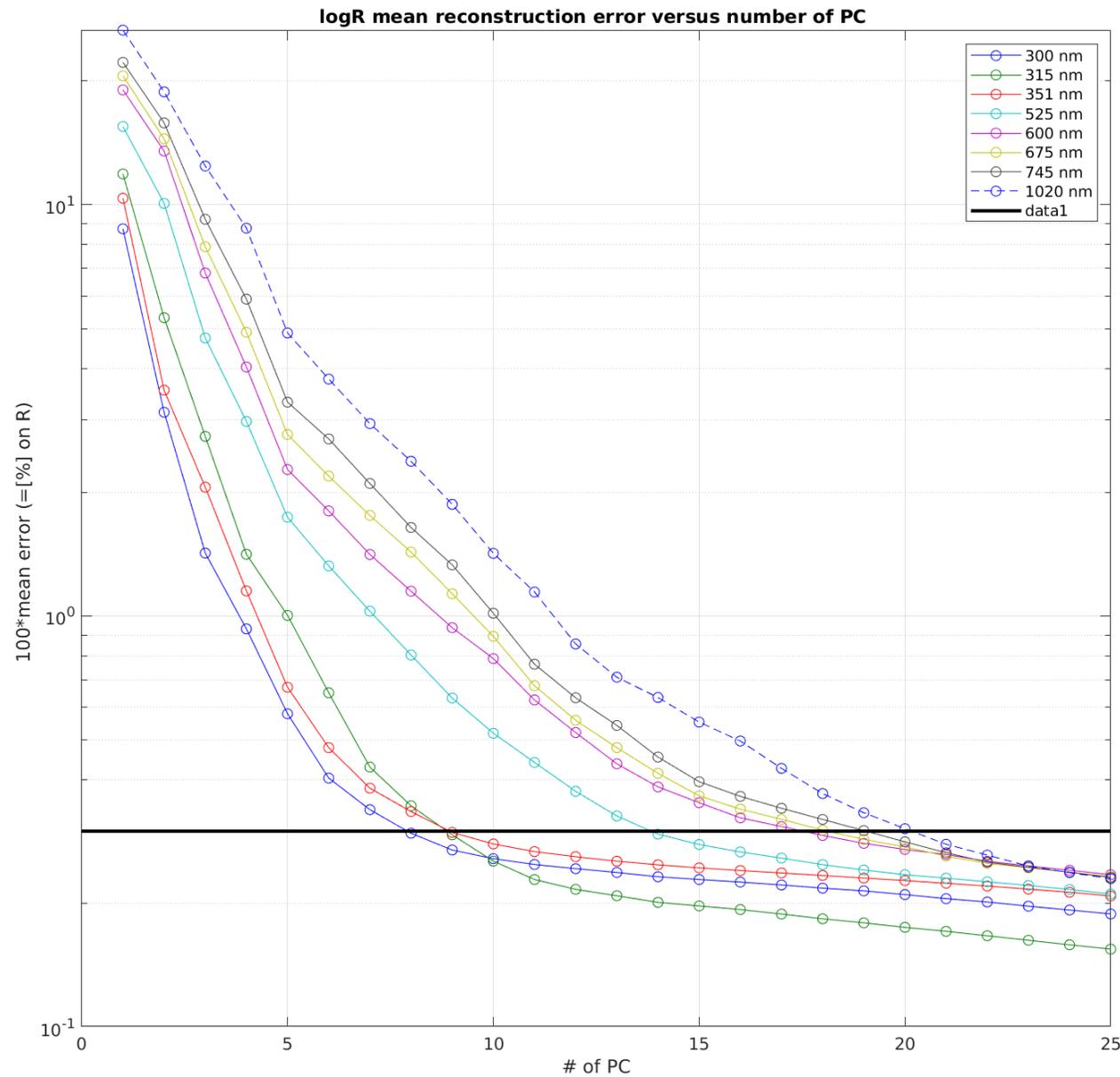


limited to unpolarized instrument for the time being

**Key concept 1: what is the information content of a radiance profile ? -> LTS PCA !!!**

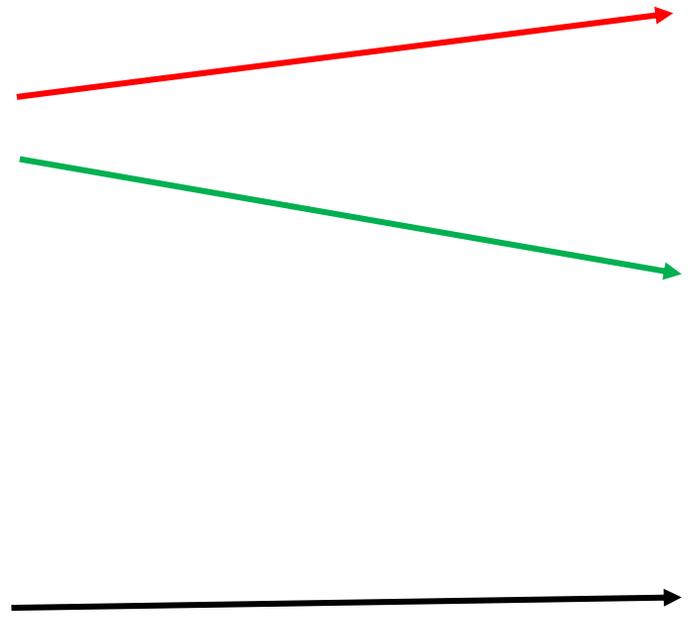
Ch / $\lambda$ [nm]	# PC
1 / 300	8
2 / 315	9
3 / 351	9
4 / 525	14
5 / 600	18
6 / 675	19
7 / 745	20
8 / 1020	21

@ 0.3 % accuracy level



Information content is reduced from  $8 \cdot 61 = 488$  to 118

Ch / $\lambda$ [nm]	# PC
1 / 300	8
2 / 315	9
3 / 351	9
4 / 525	14
5 / 600	18
6 / 675	19
7 / 745	20
8 / 1020	21

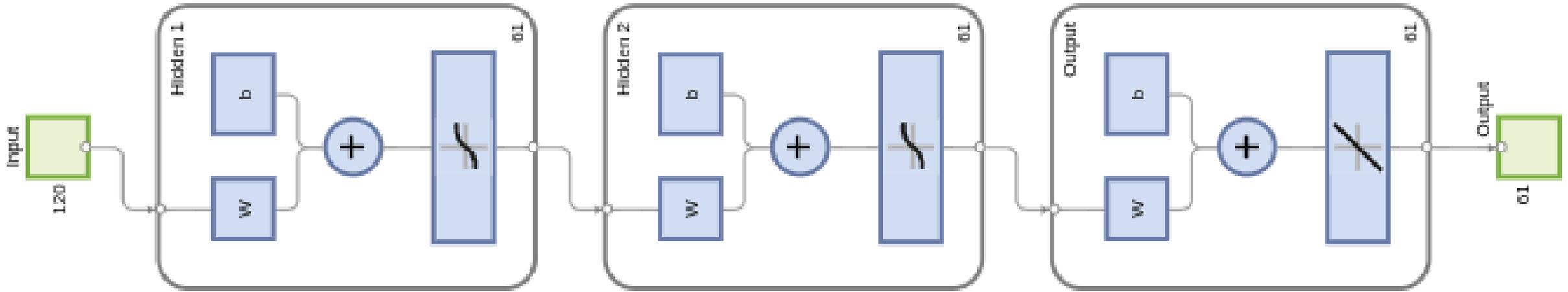


Define <b>NRV</b> (Normalized Radiance Vector)
a1,1
a1,2
⋮
a1,8
a2,1
⋮
a2,9
⋮
a8,1
⋮
a8,21
SZA
SAA

length(RV)=  
91+2=93

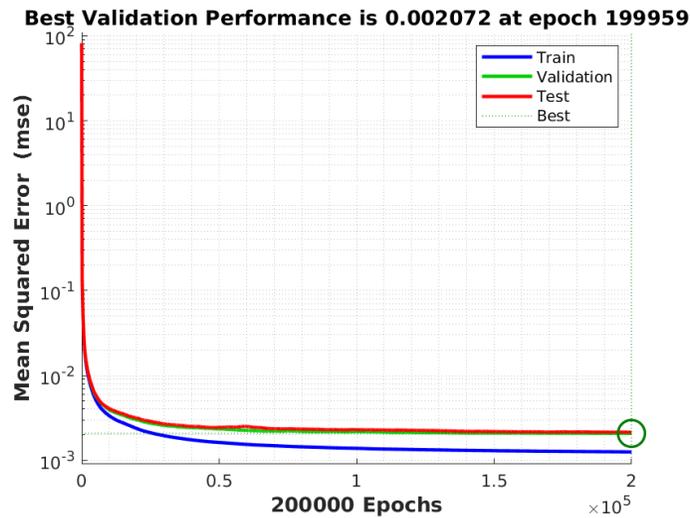
if all (8) channels  
are considered

**Key concept 2 : build an reverse mapping (118+2) NRV to ozone vmr profile (61)**

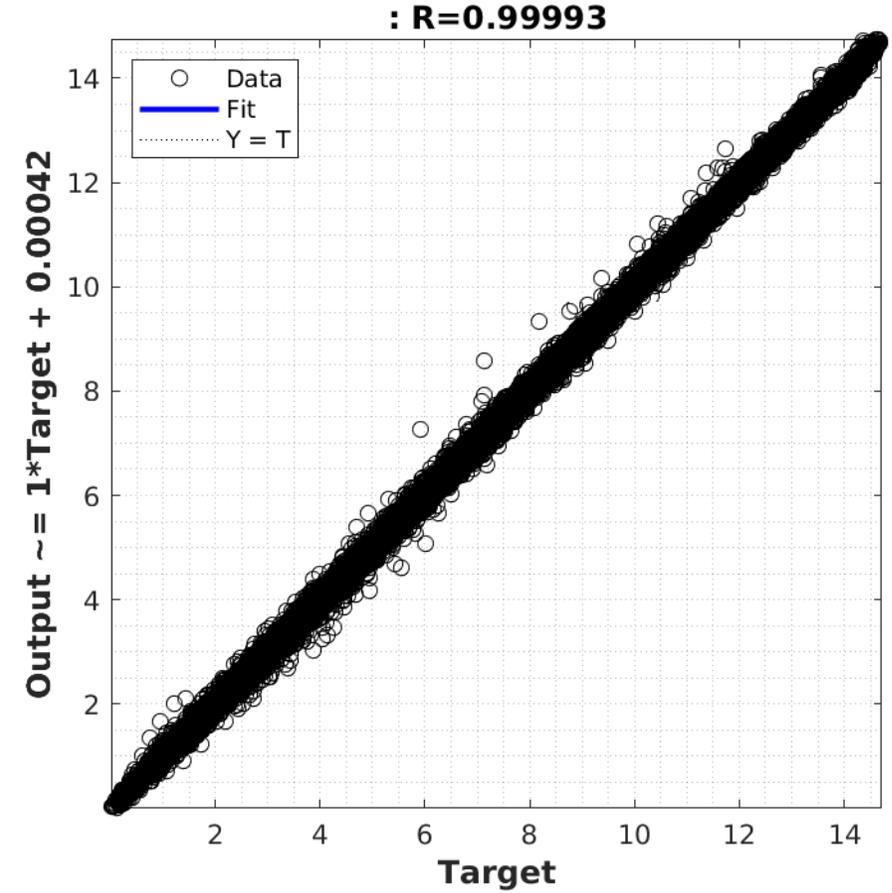


How to map a large measurement vector of onto another large state vector by a nonlinear transformation ?

**No LUT ! → Use an Artificial Neural Network**



ANN topology is: 120 x 61 x 61 x 61  
 case subsets: training= 70 % / convergence= 15% / test=15 %  
 algo= Scaled Conjugated Gradients



**Key concept 3 : reverse mapping by a shallow « deep » neural network**

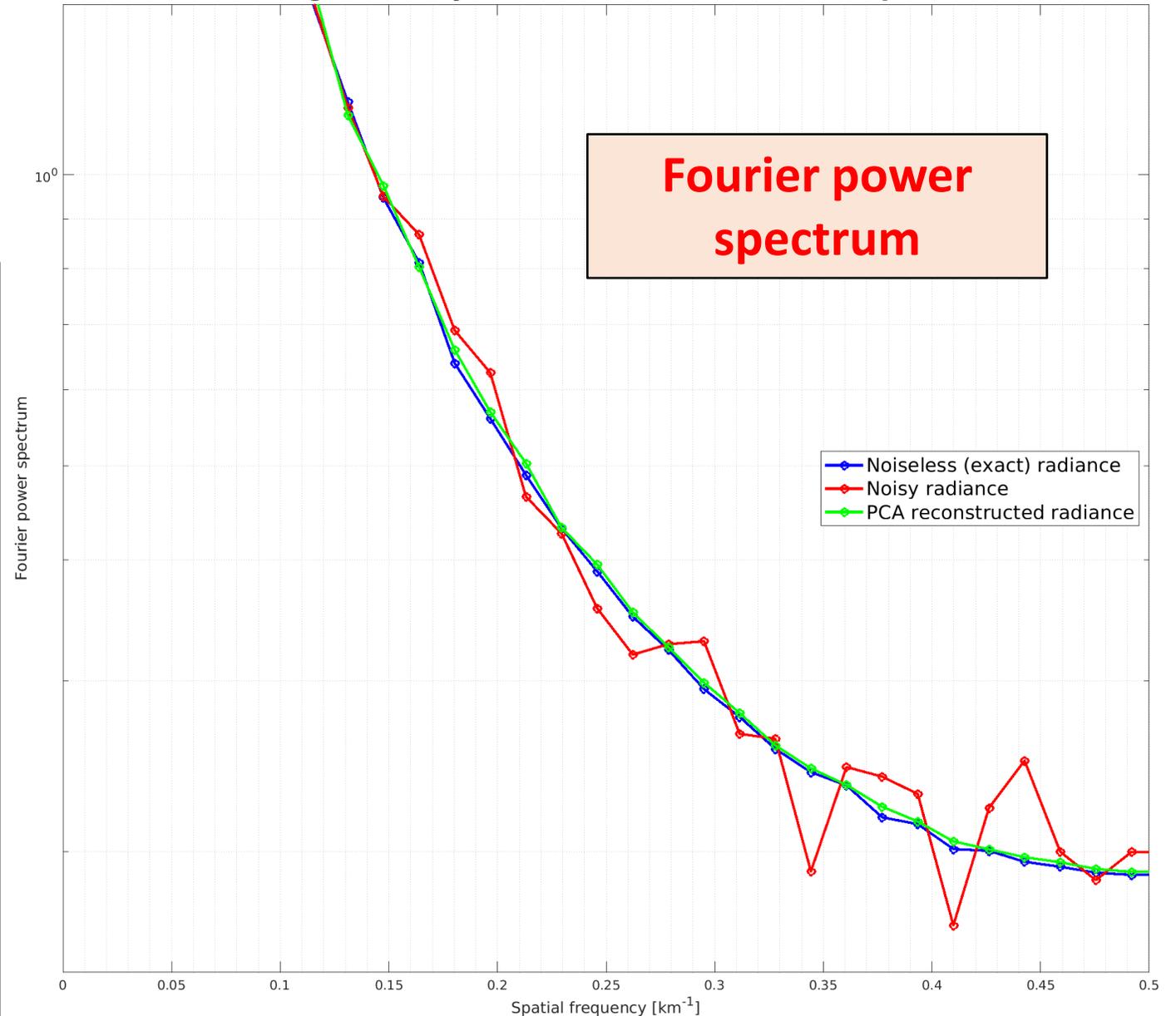




# The Magic of radiance PCA!

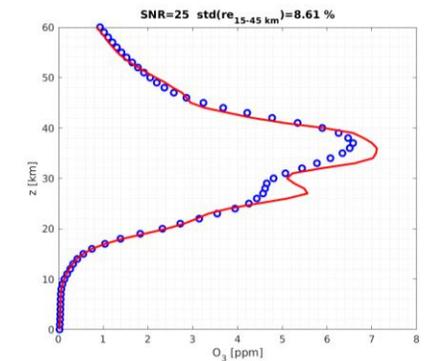
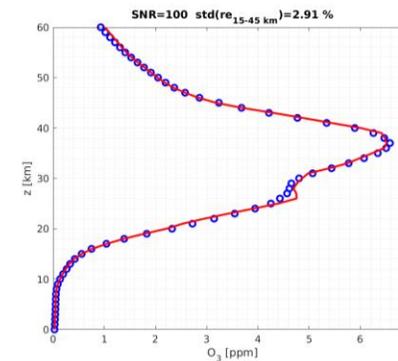
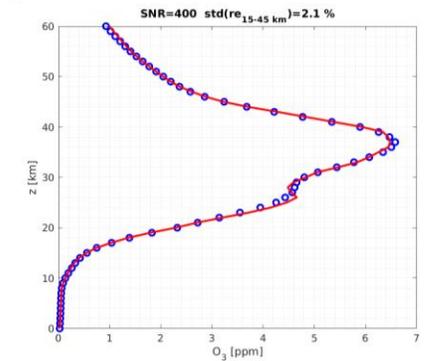
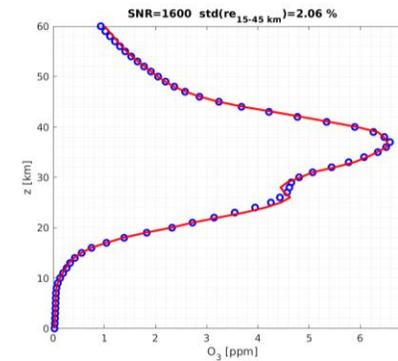
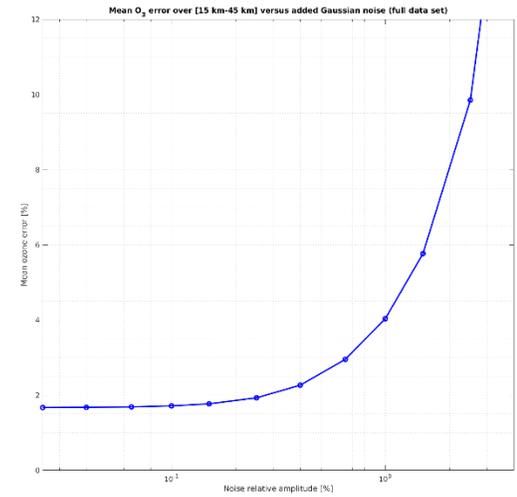
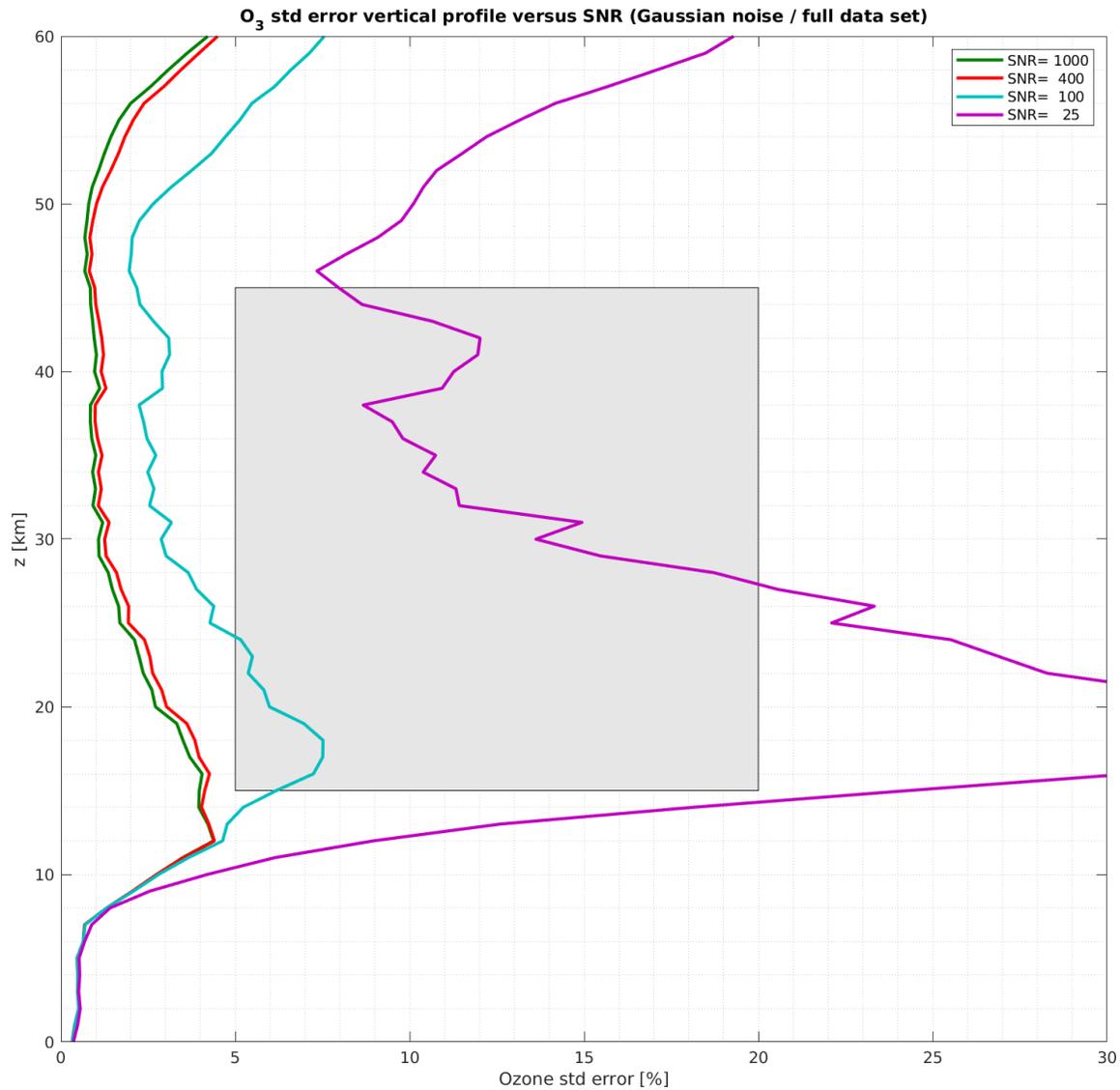
1. it tells you the information content of your observation.
2. it is a powerful (and natural) denoising filter → no need to regularize the inversion process.

Gaussian noise (amp=4 %) added on radiance profile @ 600 nm  
Filtering action by PCA reconstruction (13 components)



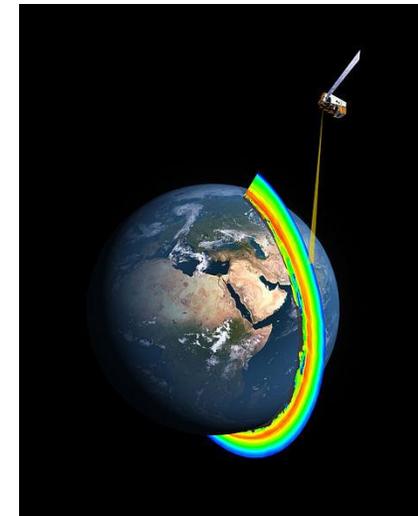
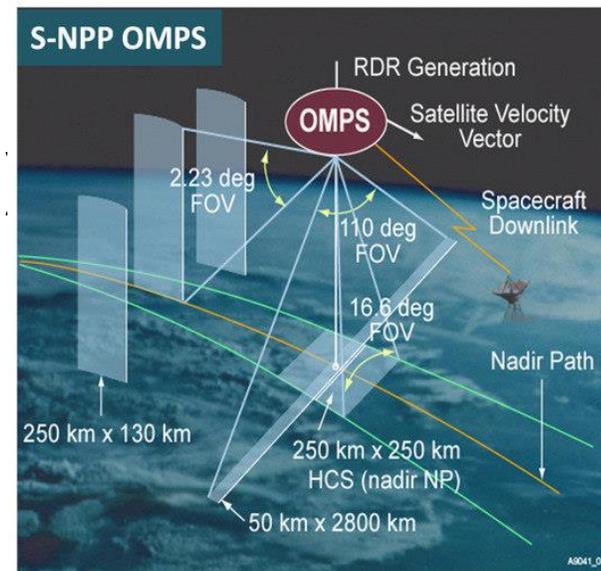
# Check robustness against noise.

e.g. ALTIUS requirements=[5 %, 20 %] in the 15-45 km range

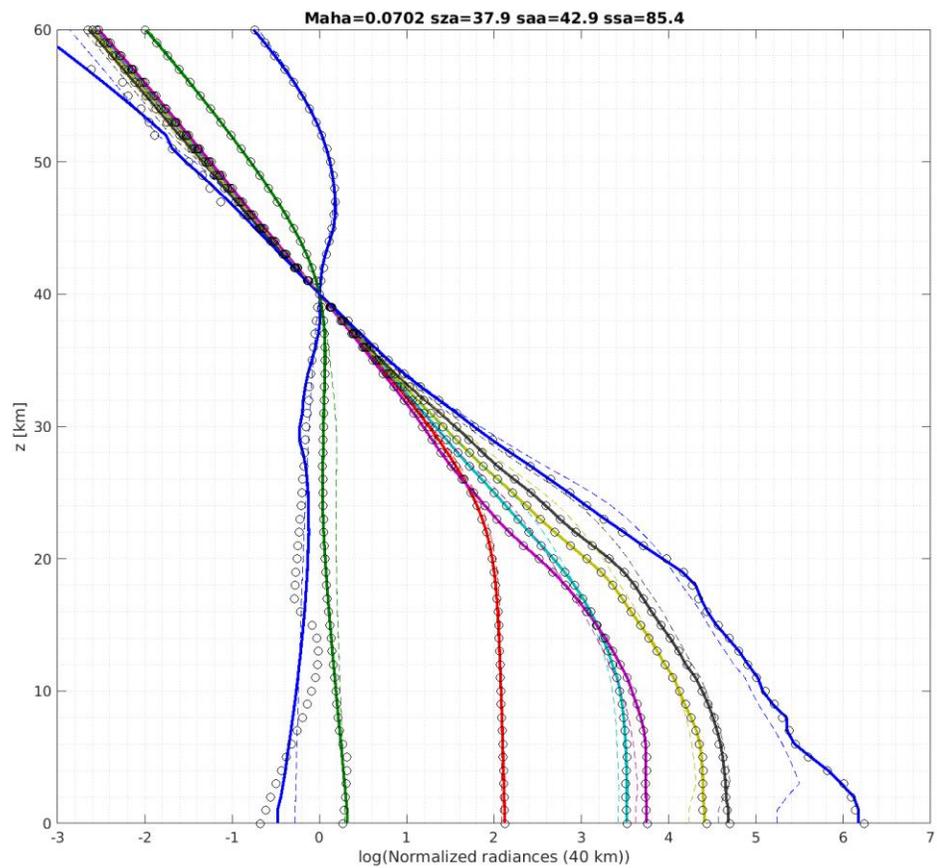


## ONNI applied to OMPS data

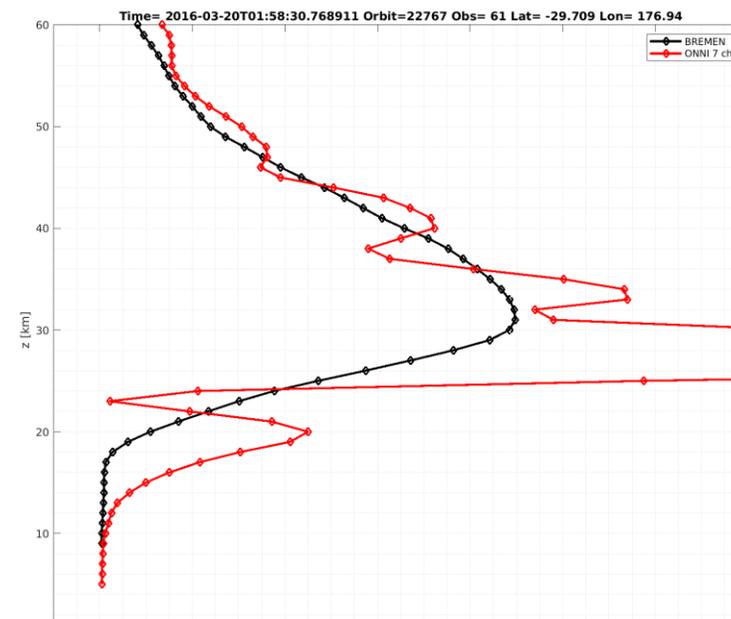
- target is March 2016 (31 x 15 x 140 potential observations)
- processing of **RAW L1 data**. No instrumental functions, no cloud detection, no TGH correction, no straylight, no extensive error budget,..etc
- objective: to show that ONNI gives “reasonable” (or even “good”) results when applied to a real case with respect to other methods. Fine tuning is possible but out of scope.
- Two competitors:
  - NASA v2.6 (Kramarova et al. [2018])
  - BREMEN (Arosio et al. [2018])



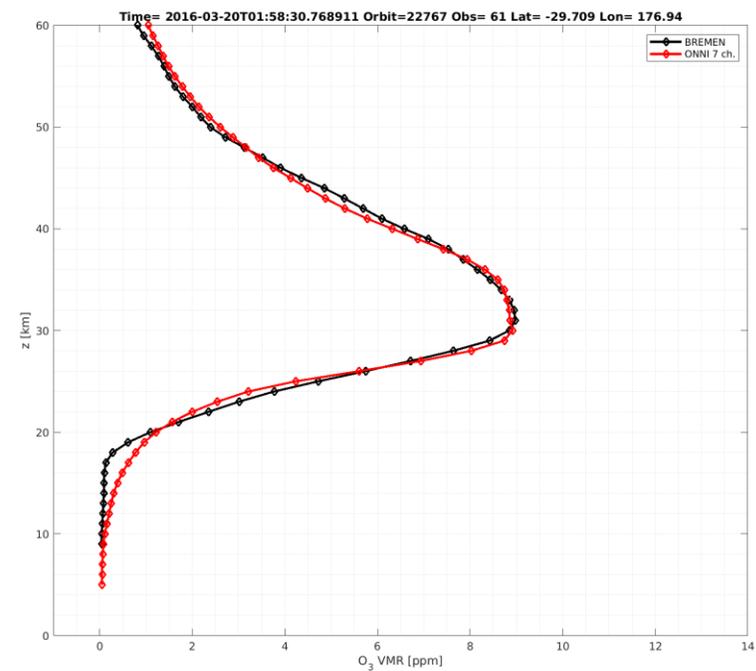
1020 nm OMPS channel is contaminated → ONNI  
was recomputed for 7 channels



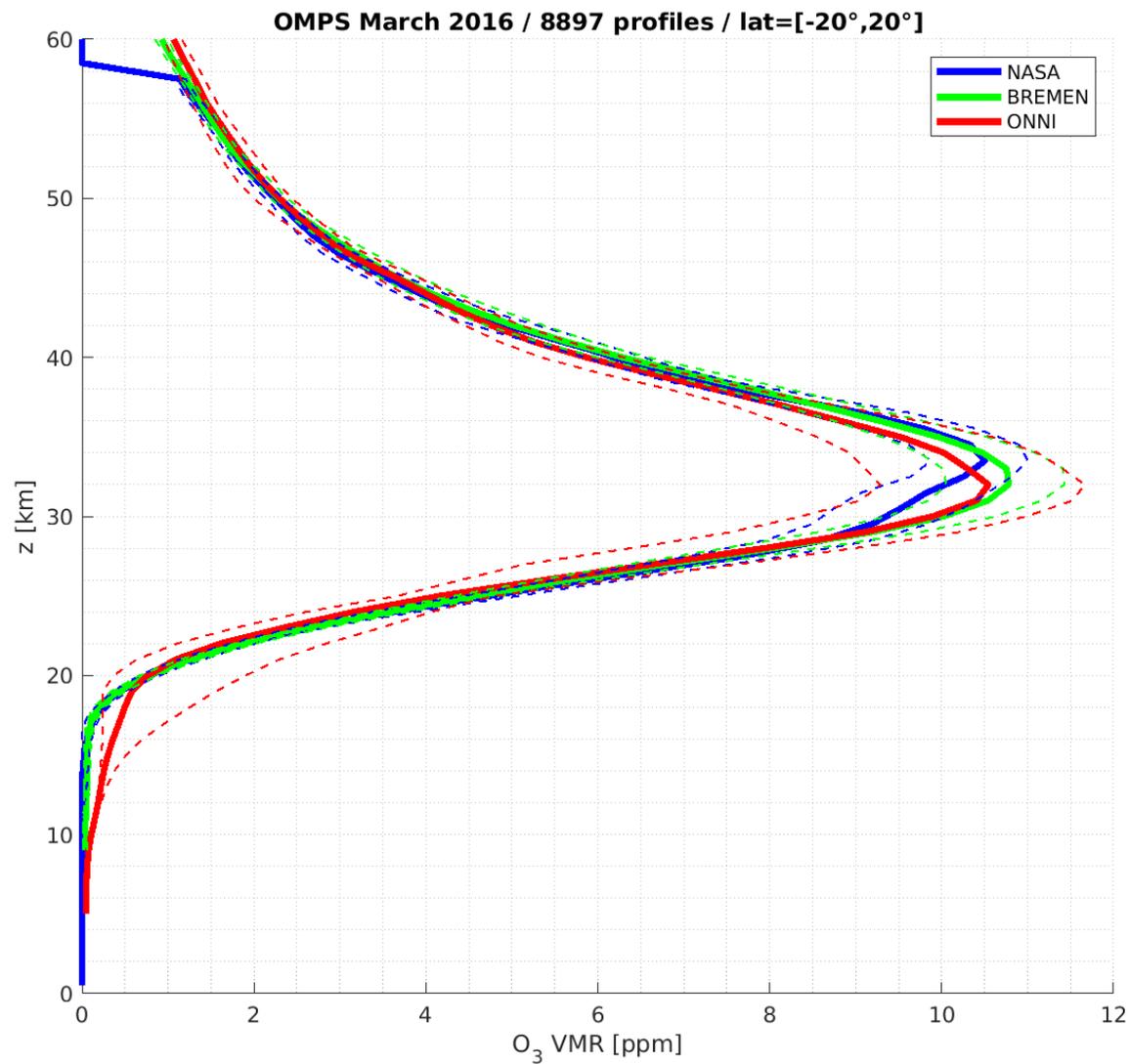
8 ch



7 ch

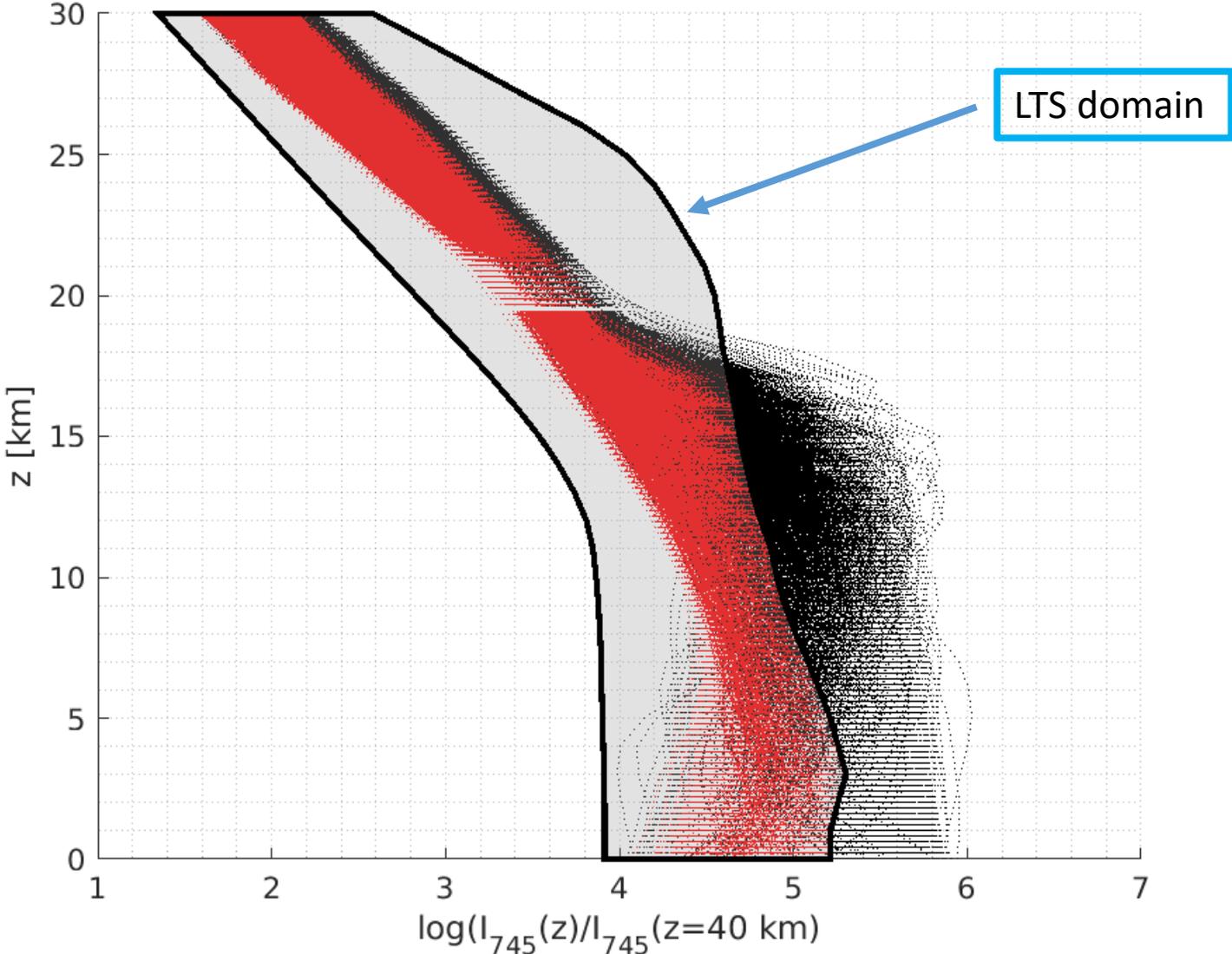


# Intercomparison in tropical band (as Arosio et al)

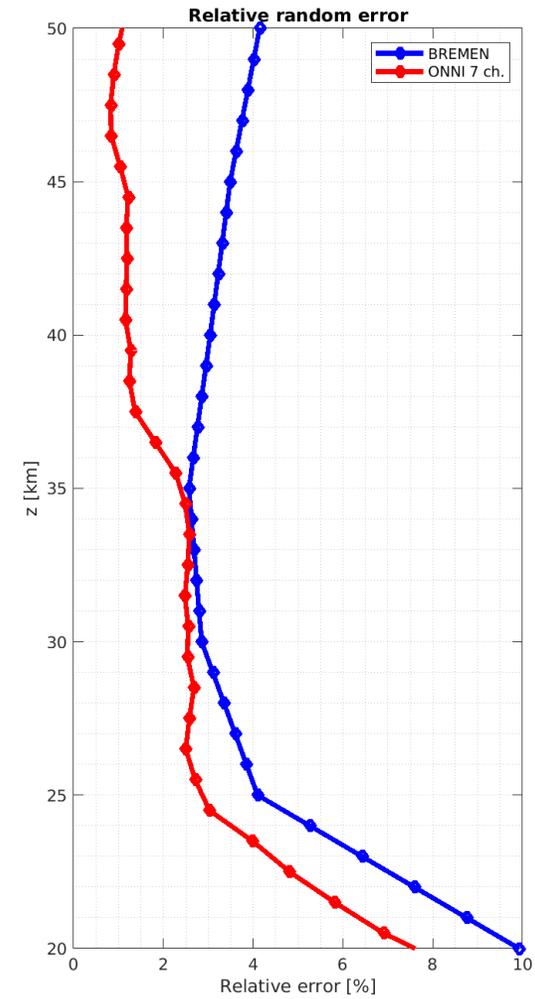
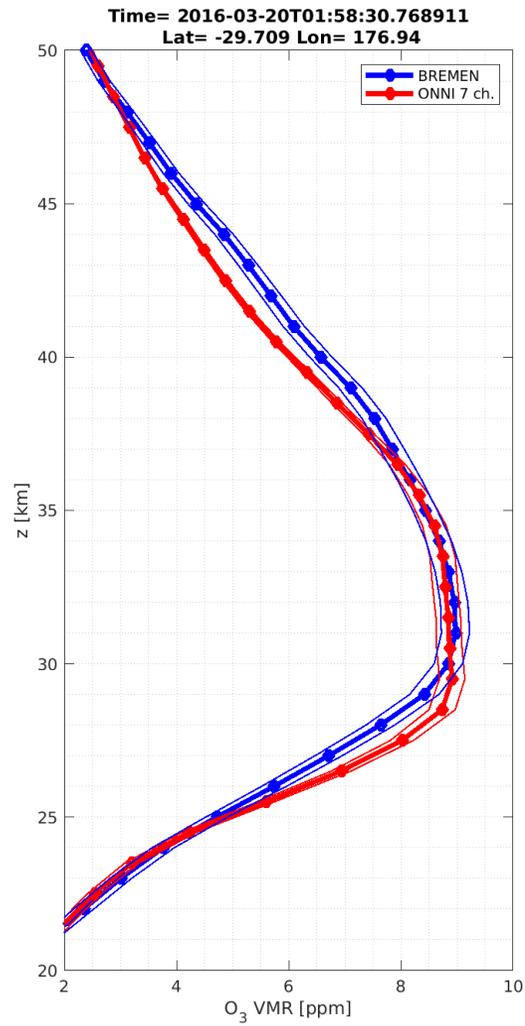
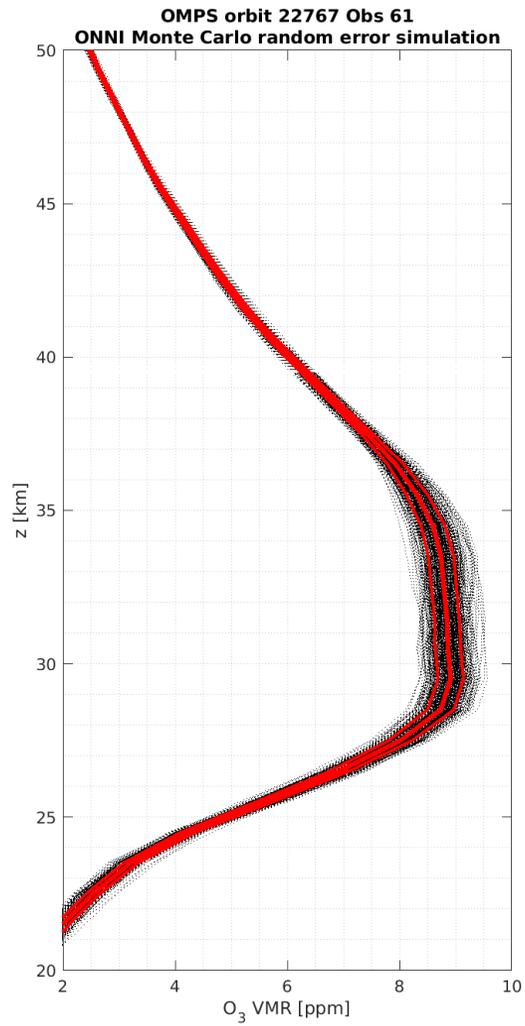


Nice but ONNI has a larger dispersion (more wavy profiles). Why ?

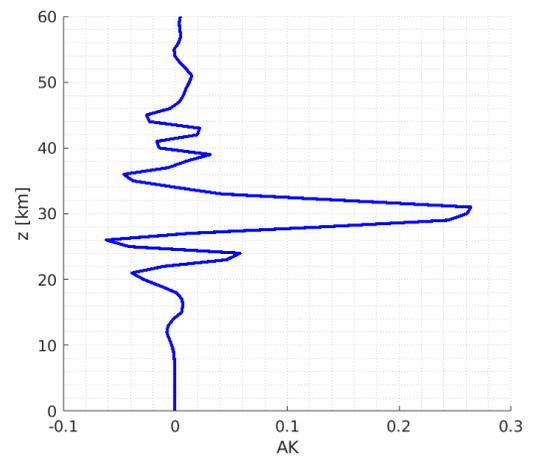
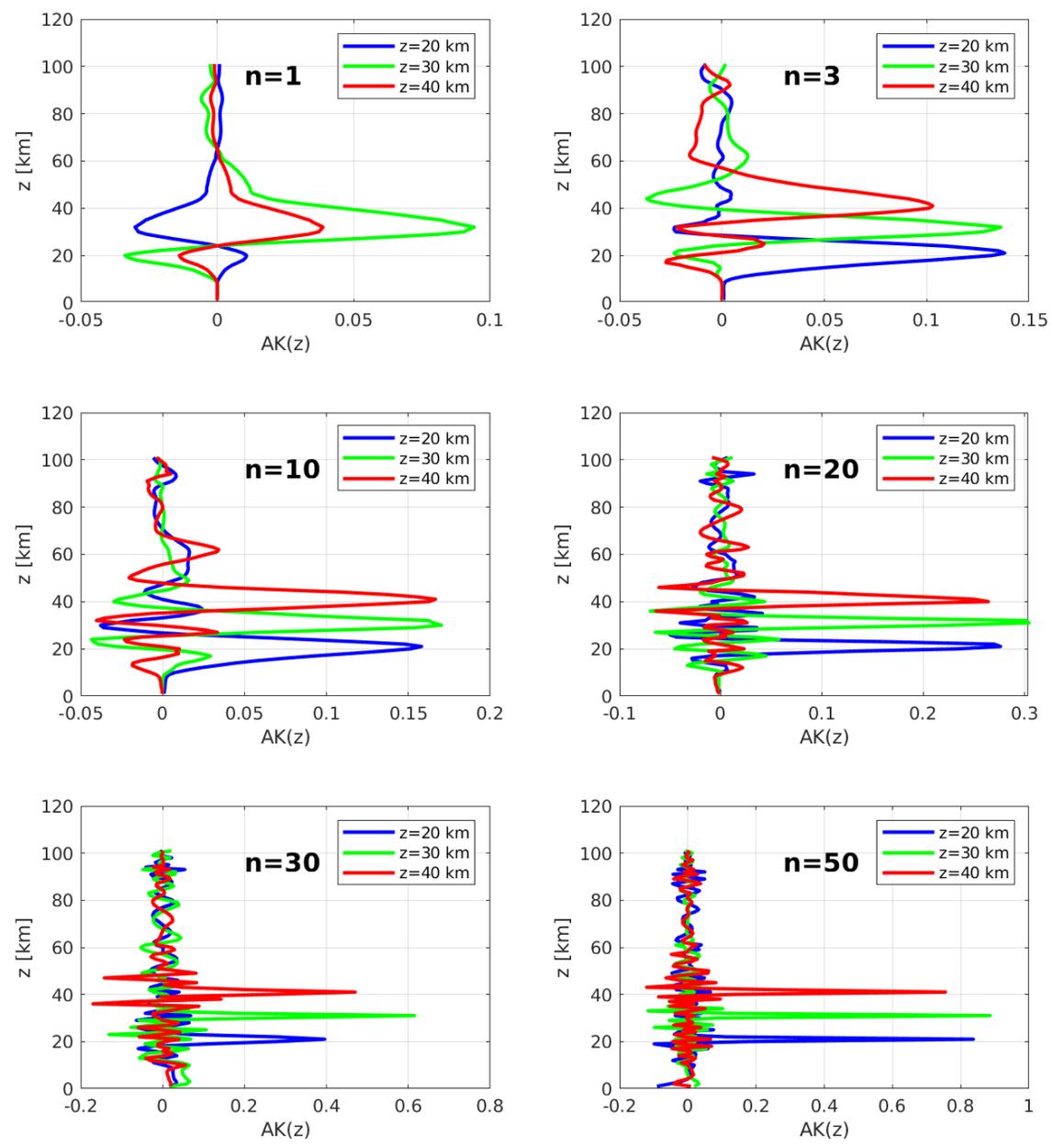
# CLOUDS !



# ONNI error computed with full Monte-Carlo simulations using logRad residuals



Profile representation by PCA defines the user-defined vertical resolution. It should match the experimental SNR level.



## Summary of “Direct inversion methods” (potentially for all applications)

### PRO

- main computing load moved to **synthetic LTS** generation.
- LTS allows for **PCA analysis** in measurement space and information compressing.
- orthogonality of PCA -> increase resolution up to the SNR.
- **reverse mapping** “measurement vector to state vector” is a nonlinear regression where **ANN** are considered to be superior to any method.
- trained **ANN are extremely cheap**: use brute force to derive error budgets and observation kernels
- **self-denoising**: no regularization needed

### CON

- A **change in the forward** model triggers a re-generation of the LTS (except if these changes are parameterized and perturbative)
- ANN topology ? **heuristic**, no clear rules, trial and error. Number of neurons, hidden layers, activation functions. Optimal topology may depend on the solution...
- so far, nobody understands **what happens inside the box ?**

Spare slides

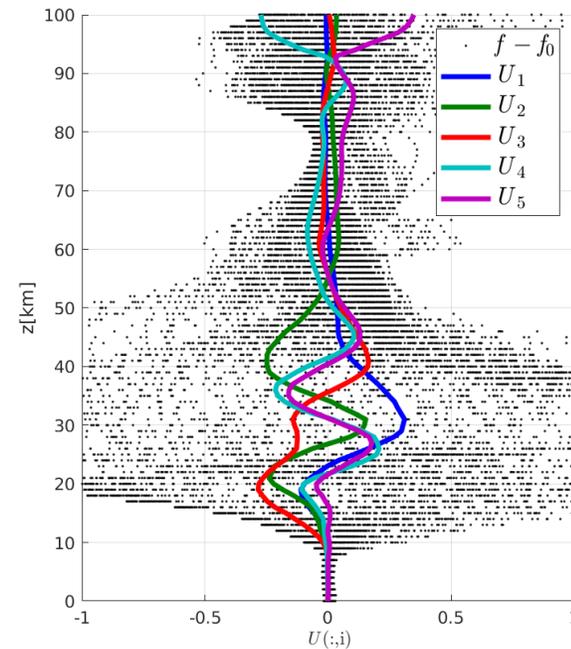
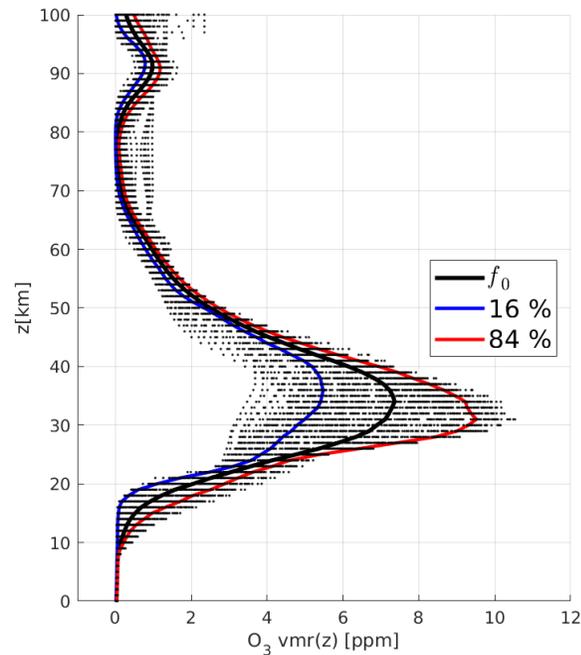
## Some open questions:

- investigate hallucinations ( 2-3 % of cases)
- is full PCA (spatial+spectral) better to avoid spectral non-orthogonality ?
- instrumental parameters
  - direct parameterization to append to the measurement vector
  - LTS update from a subset and training from unperturbed ANN
- parallelization of L-M algorithm (Python ?)
- optimal LTS generation: which minimal training set of state vectors must be generated to represent the measurement space (GAN ? variational autoencoders ? )
- ...

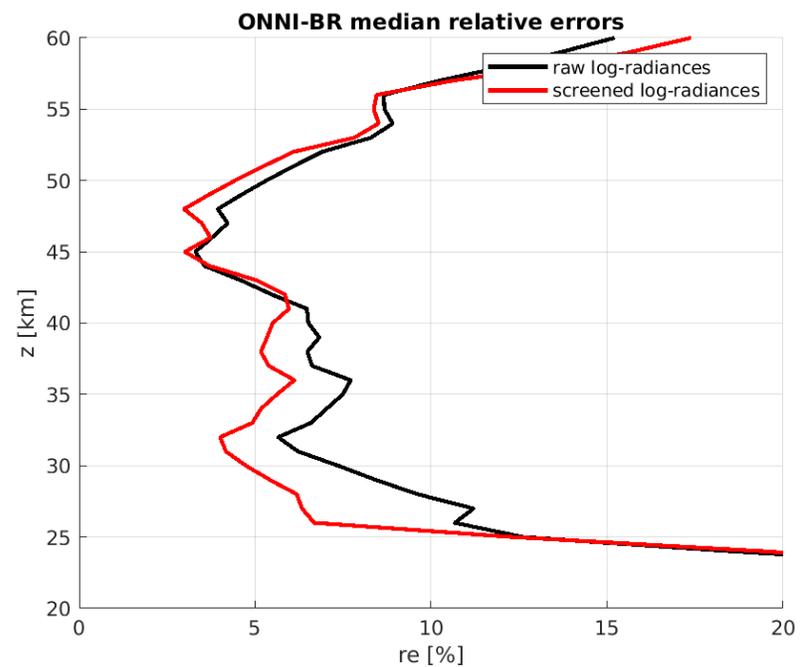
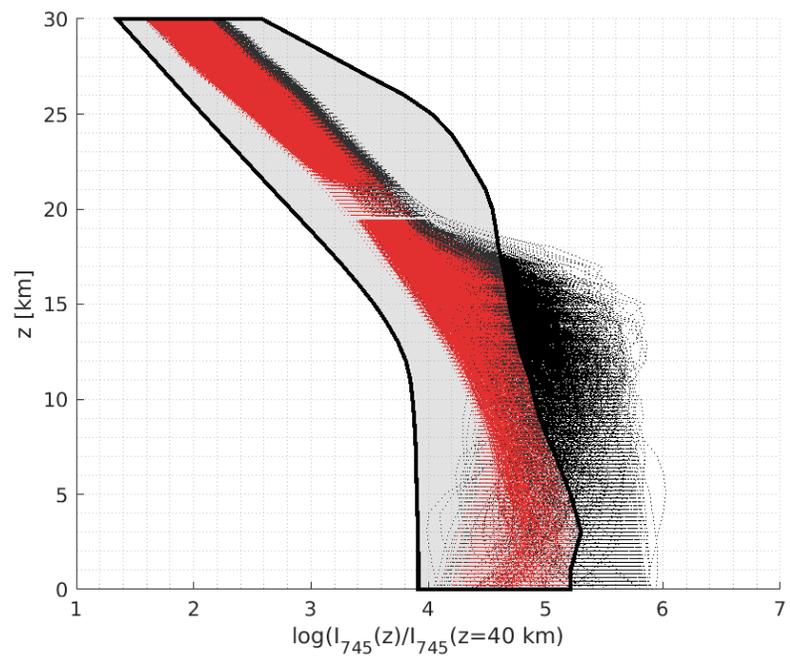
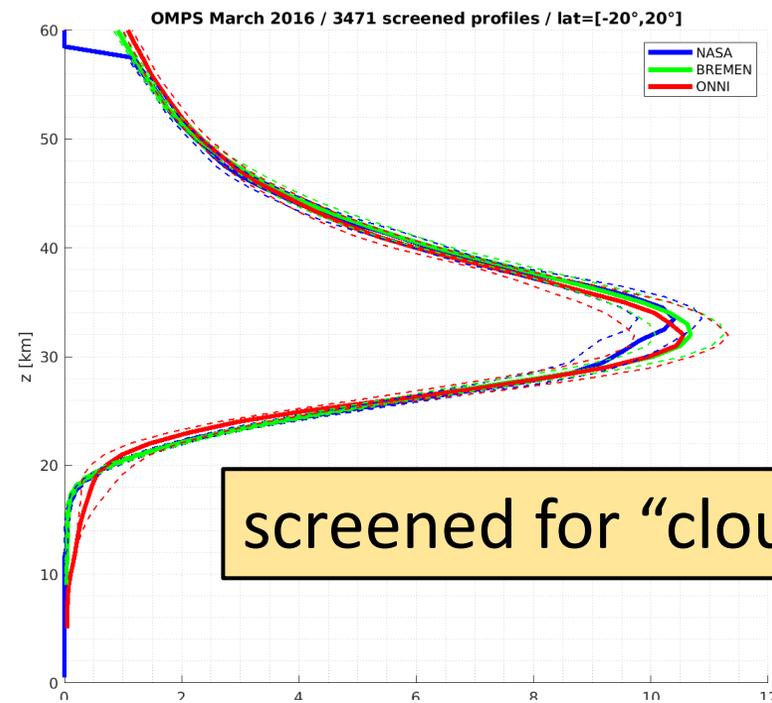
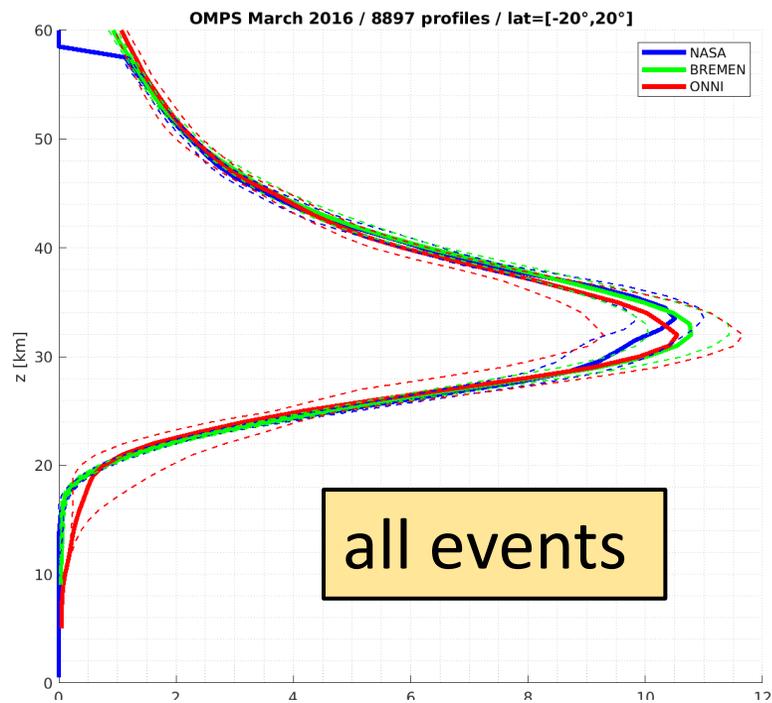
# Reminder about PCA

Start from a climatology  $\rightarrow$  compute the covariance matrix of the vertical profiles  $\rightarrow$  SVD  $\rightarrow$  use the eigenvectors as “principal components”

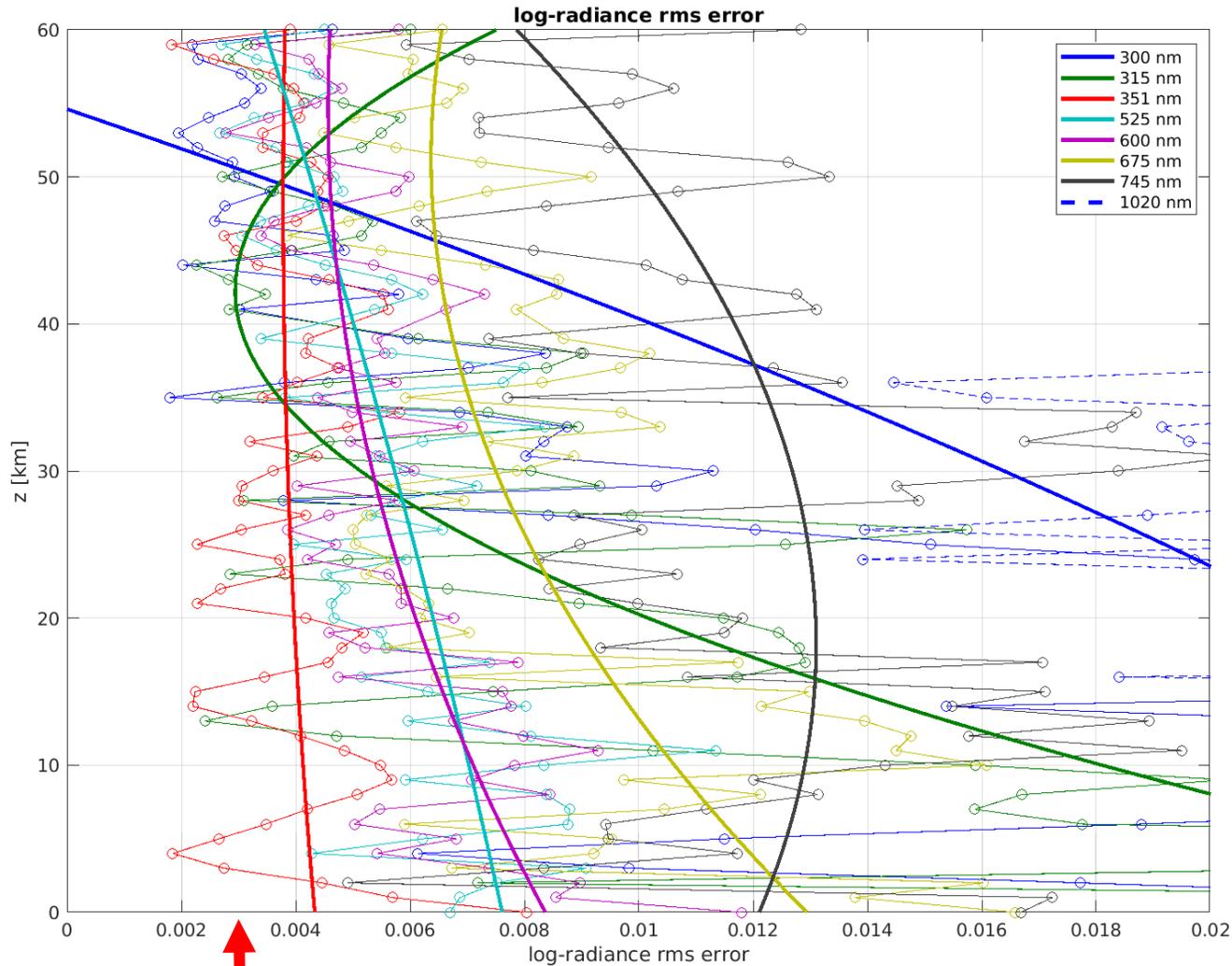
$$f(z) \simeq f_0(z) + \sum_{i=1}^n a_i \phi_i(z) \equiv \mathbf{f}_0 + \mathbf{a}(1:n) * \mathbf{U}(:, 1:n)^T$$



	RTM	Spectral range	normalization	regularization	Albedo, aerosols etc	CS	Algo
<b>NASA</b>	<u>GSLs</u>	1 triplet @ 600 nm [549-633 nm] for 12.5-35.5 km 3 doublets [302,312,322]/355 for 28.5-50.5 km	UV: 55.5 km VIS: 40.5 km	Covariance matrix (McP-Labow)	Albedo retrieved at 675 nm Retrieve ozone 1 km above cloud height (color method) Aerosol: independent retrieval (Loughman, [2017])	Bass &Paur	Opt.Estimation
<b>BREMEN</b>	<u>SCIATRAN</u>	285-302 nm 305-313 322-331 508-660	63.5 km 52.5 47.5 42.5 $\text{Log}(y/y_*)-P_n$	1 <sup>st</sup> order Tikhonov	Albedo: simul retrieval Reject TH's below threshold (color method) Aerosol: extinction at 869 nm + frozen log-norm and Mie	Serdyuchenko	Opt.Estimation
<b>ONNI</b>	<u>SMART-G</u>	300,315,351,525,600,675,745,(1020) nm	z= 40 km	PCA low-pass filter on logRad	NNI integrates the ozone signature wrt NRV	Serdyuchenko	Direct NNI



# Stochastic error budget



theoretical limit=0,3%

Measurement noise is estimated from PCA residuals (same approach as Arosio et al. using inversion residuals)

The noise amplitude is altitude and wavelength dependent.

## Summary of ONNI (specific to limb scattering)

- logRadiance **PCA representation coupled with a NN reverse mapping** is an efficient inversion method for ozone retrieval
- **self-consistency and robustness** have been verified
- **no regularization**. The PCA is a low-pass filter consistent with measurement SNR
- **error budget and AK's** are cheap to compute, even with full Monte-Carlo simulations.
- A **raw** application of ONNI to OMPS data shows **fair inter-comparison with NASA and BREMEN algos**