

Advances in research of equatorial plasma depletions enabled by the Swarm missions

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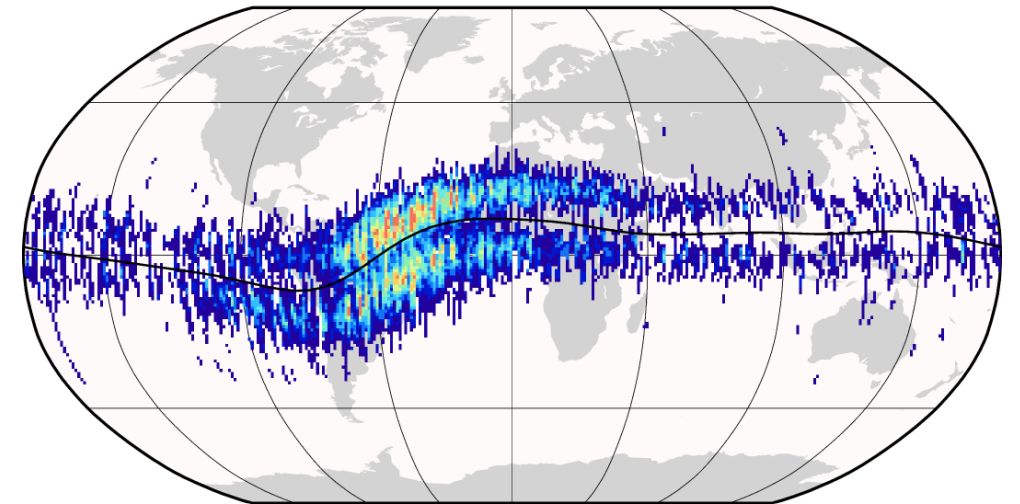
³ Department of Space Physics, Electronic Information School, Wuhan University

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⁵ Geomagnetism, GFZ Potsdam

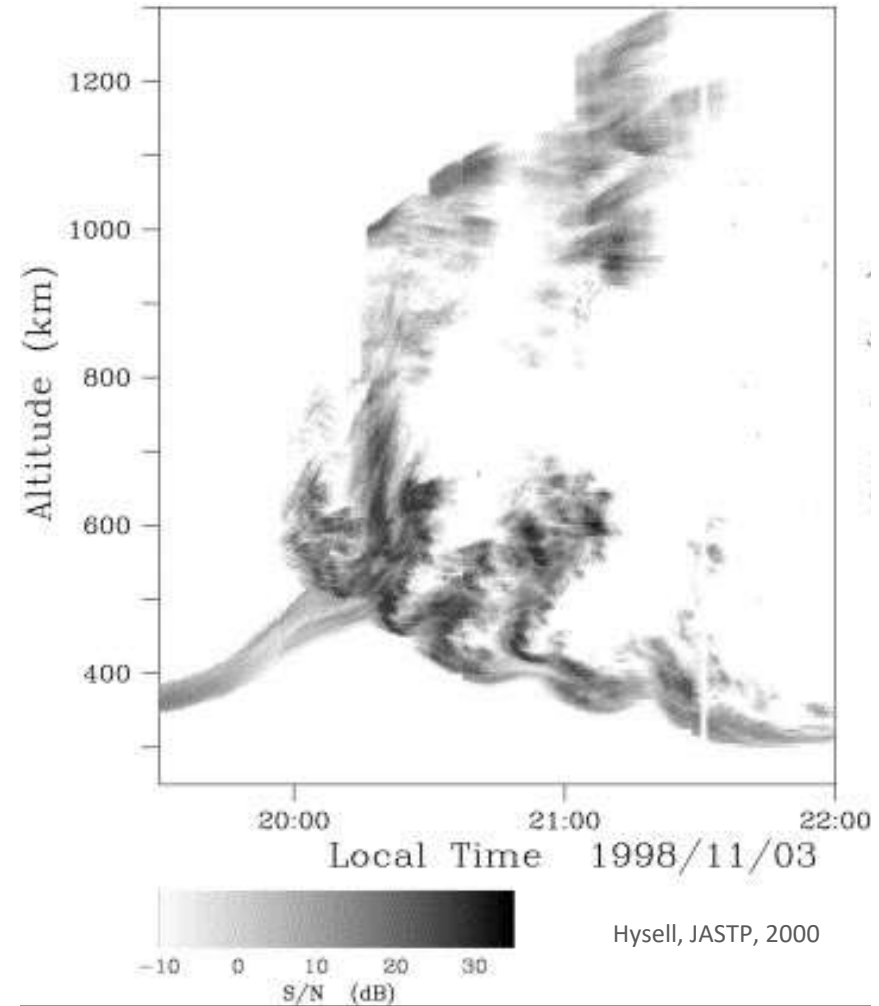
⁶ Korea Astronomy and Space Science Institute

⁷ Planetary Environmental and Astrobiological Research Laboratory, Sun Yat-sen University



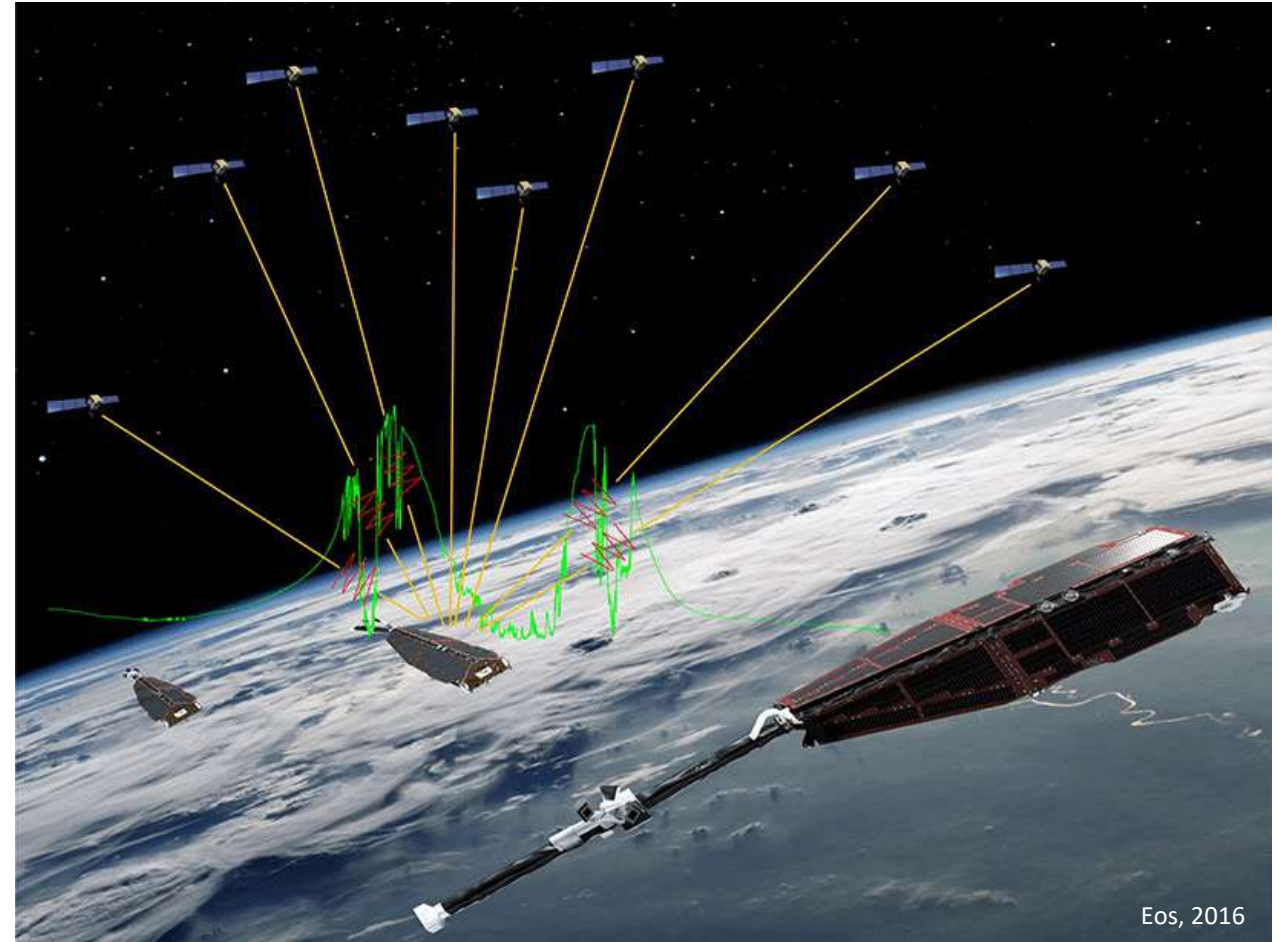
Equatorial plasma irregularities

- Major event (“convective storm”) in space physics
 - Magnetic field/ionosphere/atmosphere interaction
 - During geomagnetic quiet and disturbed times
 - Initiation and growth processes under research



Equatorial plasma irregularities

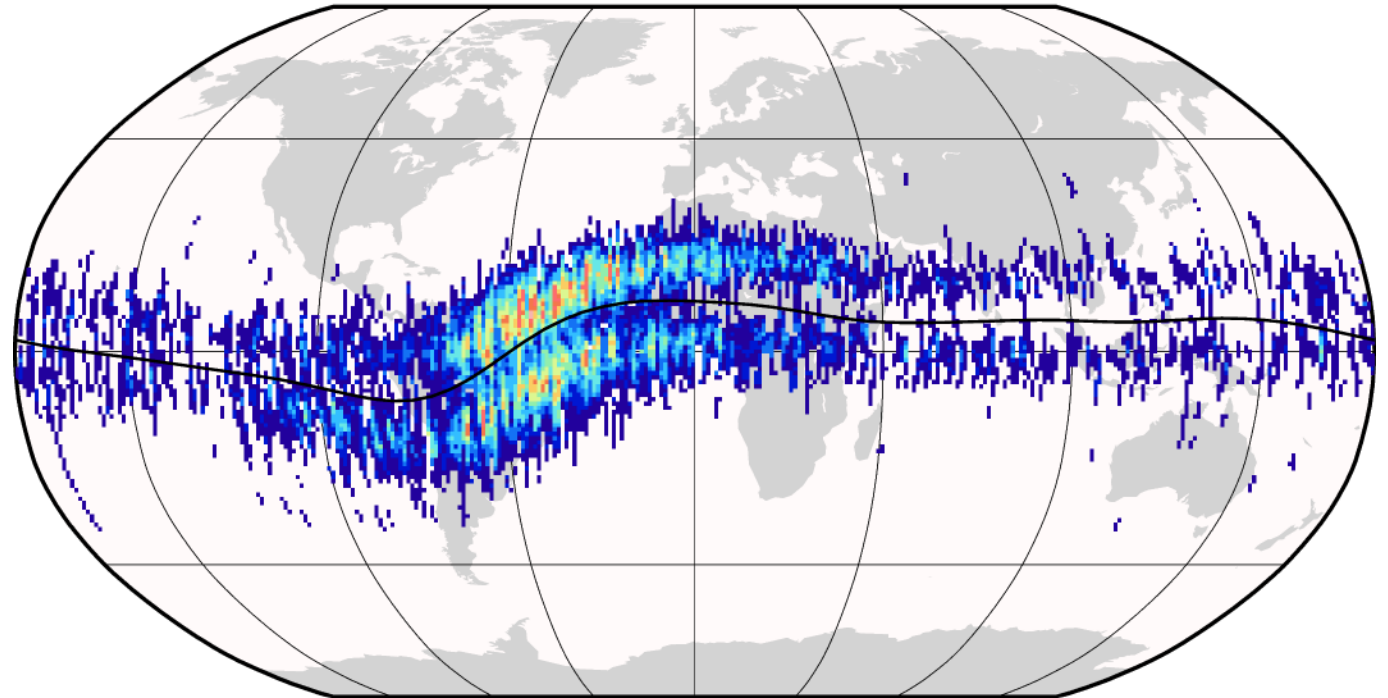
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 - Initiation and growth processes under research
- Affecting radio wave propagation
 - Disturbance / Loss of GNSS signals



Equatorial plasma irregularities

Swarm C 12/2013 - 03/2024

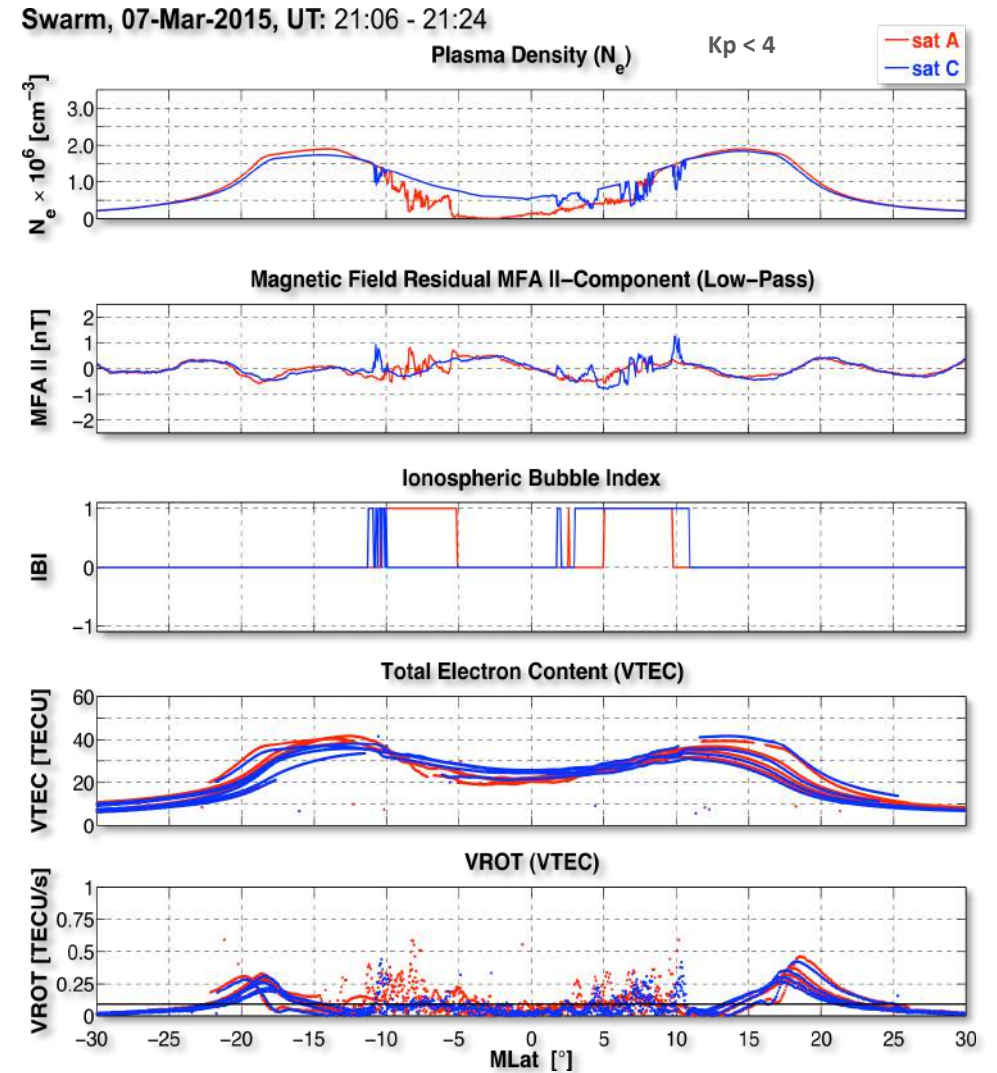
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- Systematic signatures in the magnetic field



Number of detections of irregularities in the magnetic field (IBI)

Swarm observations of plasma irregularities

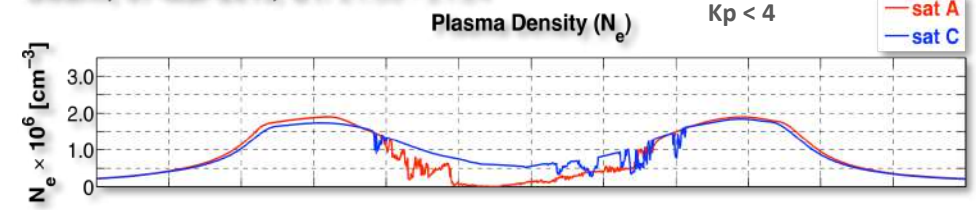
- In situ data – at about 450 km
- Multi parameter, high-precision



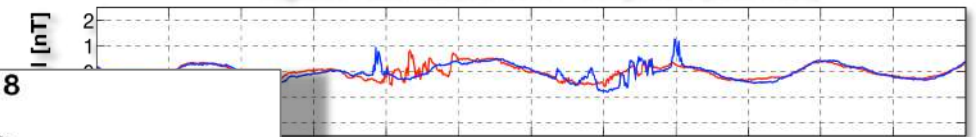
Swarm observations of plasma irregularities

- In situ data – at about 450 km
- Multi parameter, high-precision
- Complements remote sensing methods

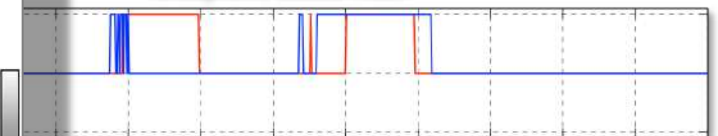
Swarm, 07-Mar-2015, UT: 21:06 - 21:24



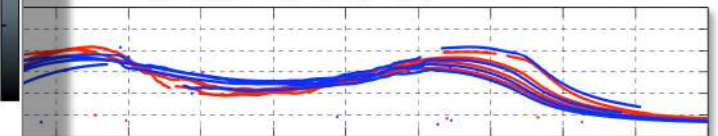
Magnetic Field Residual MFA II-Component (Low-Pass)



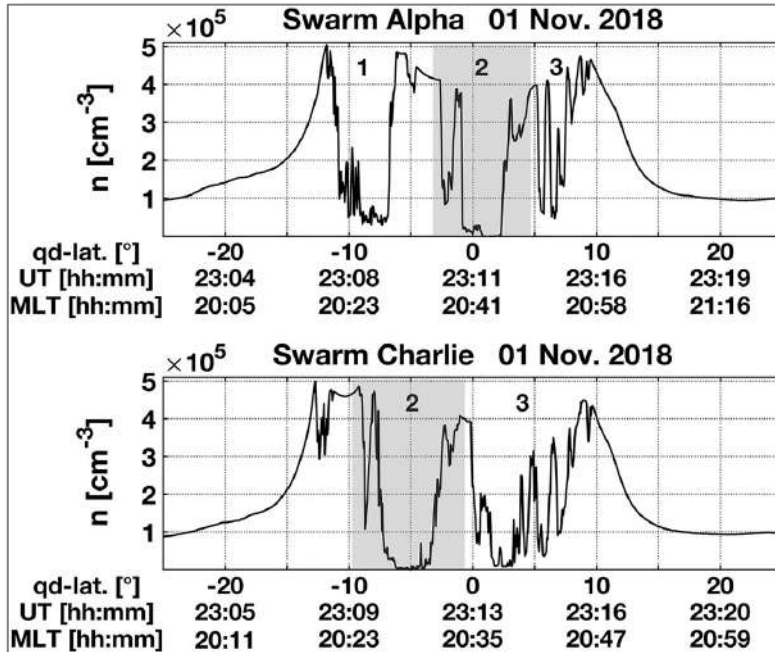
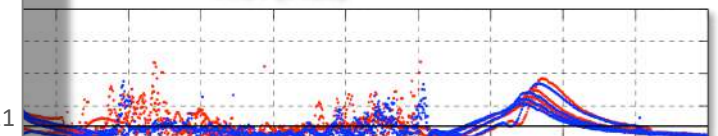
Ionospheric Bubble Index



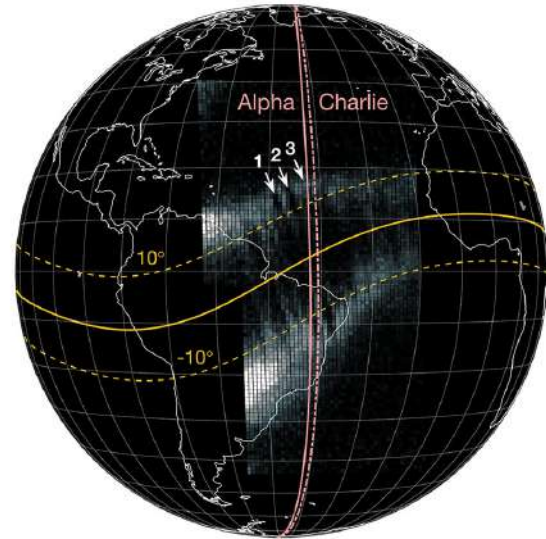
Total Electron Content (VTEC)



VROT (VTEC)



GOLD 01 Nov. 2018



Scans start at 23:10 UT

Rodríguez-Zuluaga et al., ESS, 2021

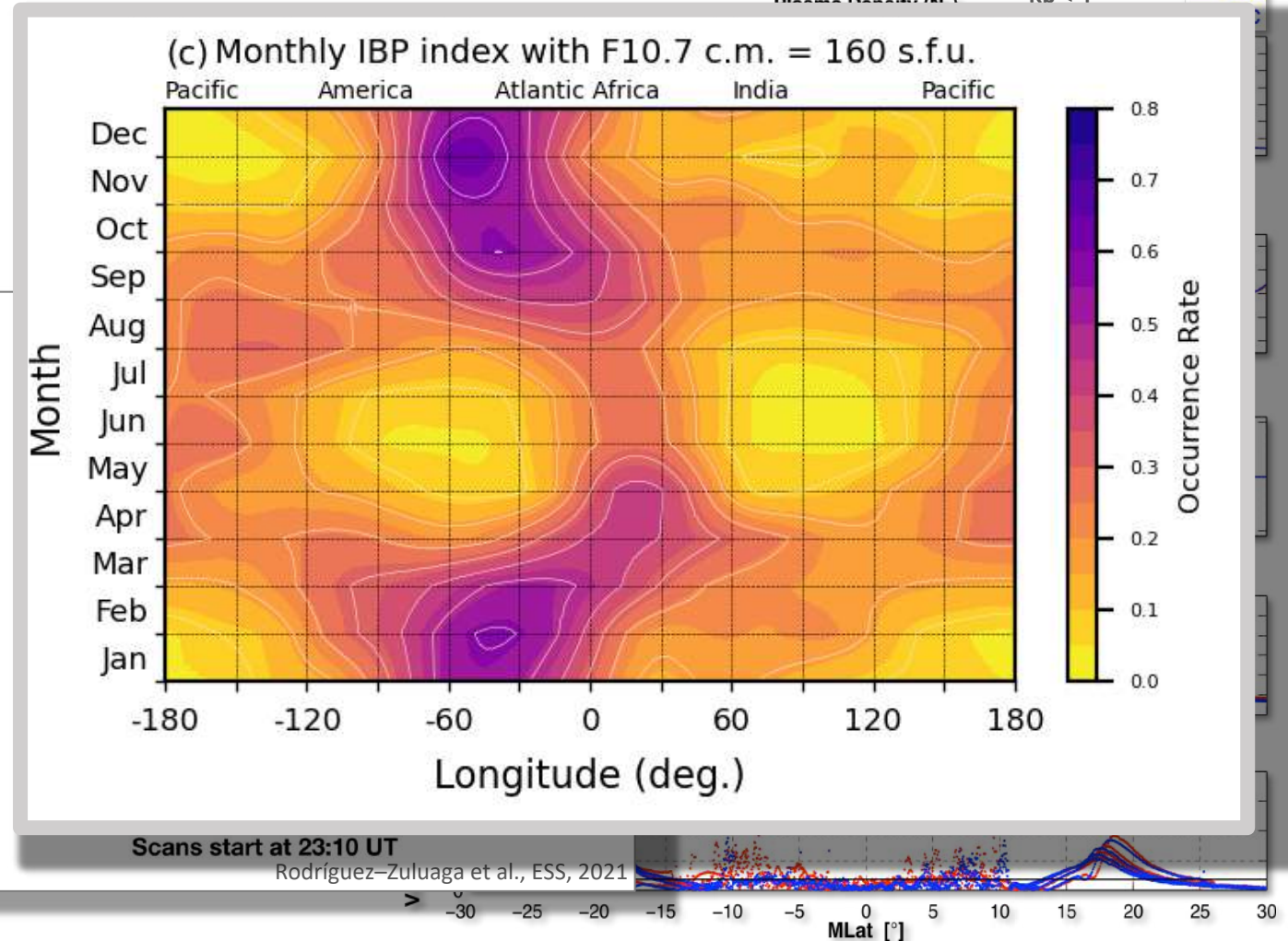
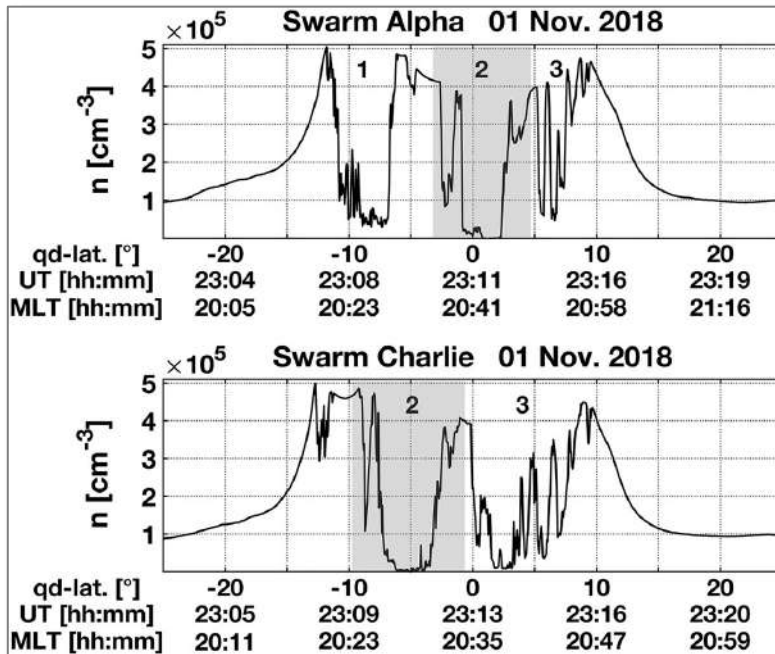
MLat [°]

Swarm observations of plasma irregularities

- In situ data – at about 450 km
- Multi parameter, high-precision
- Complements remote sensing methods
- stable and long-term (> 10 years)

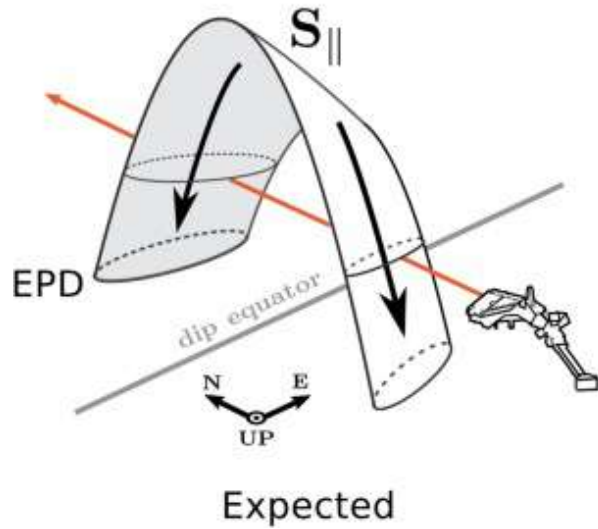
Swarm, 07-Mar-2015, UT: 21:06 - 21:24

Plasma Density (M) Kp < 4 — sat A



Determining the location of the dynamo source

Field-aligned Poynting flux (S_{\parallel})

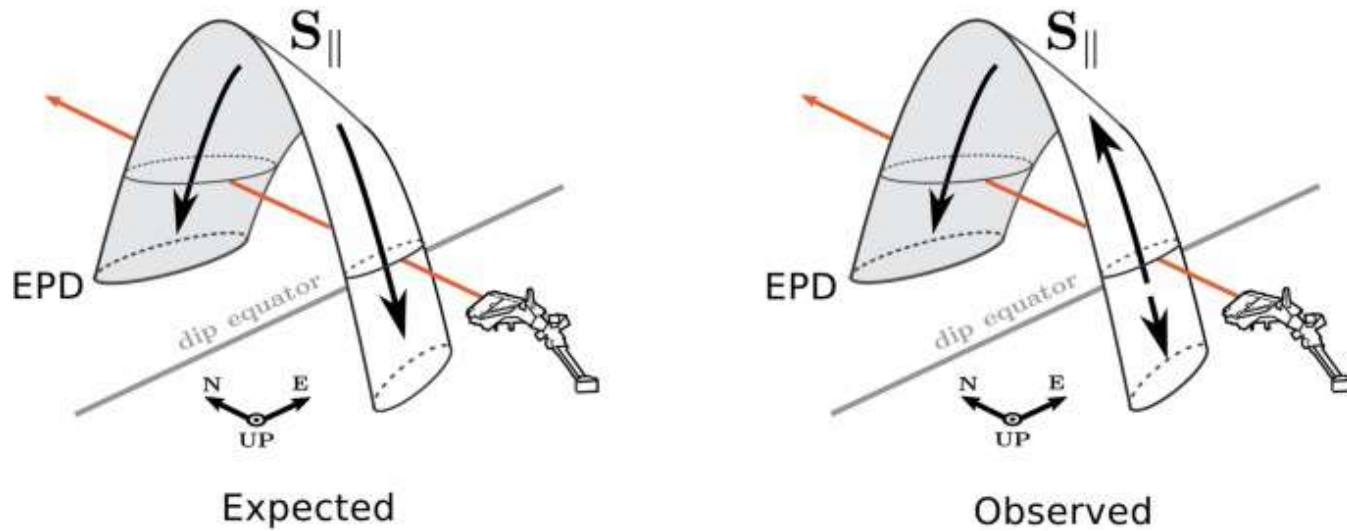


- F-region winds act as the dynamo source
- E-region conductivity acts as the load
- Well centred within a symmetric ionosphere

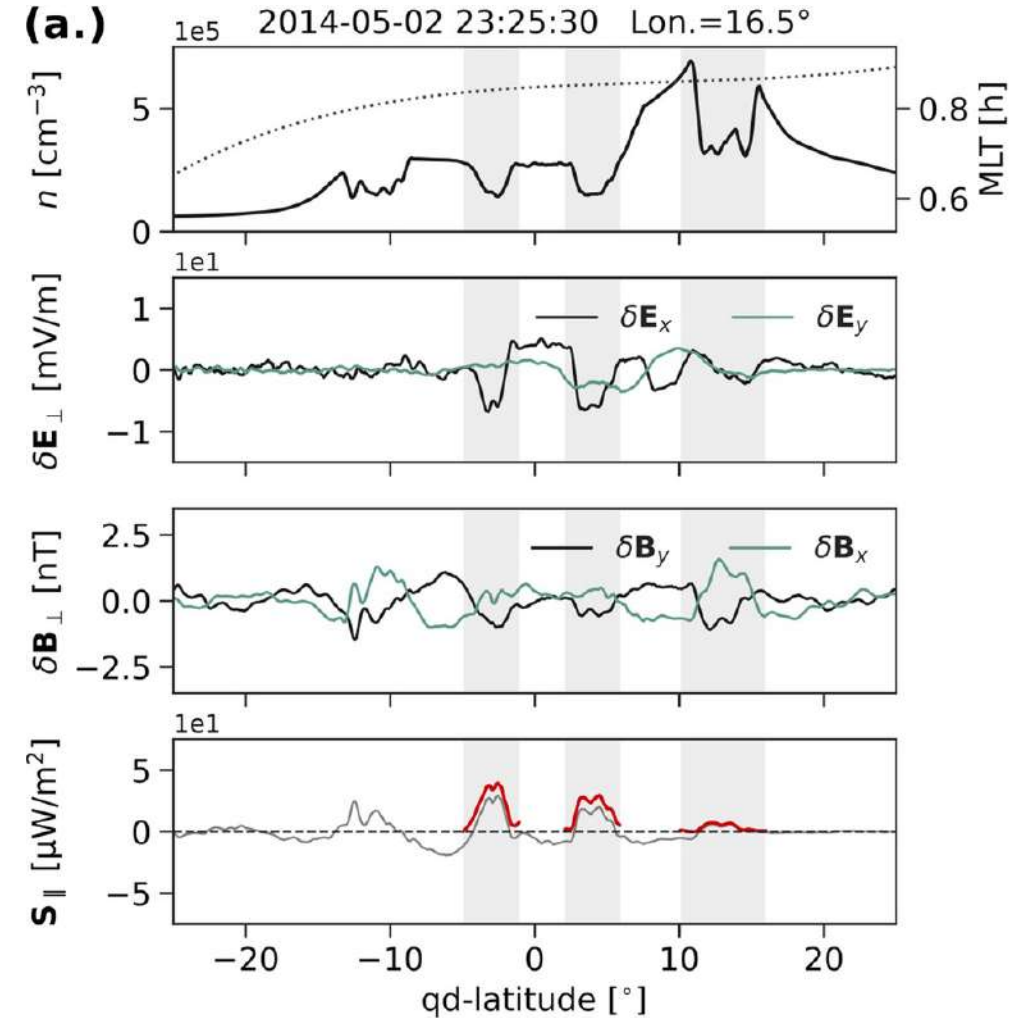
Rodríguez-Zuluaga et al., EPS, 2022; SciRep, 2019.

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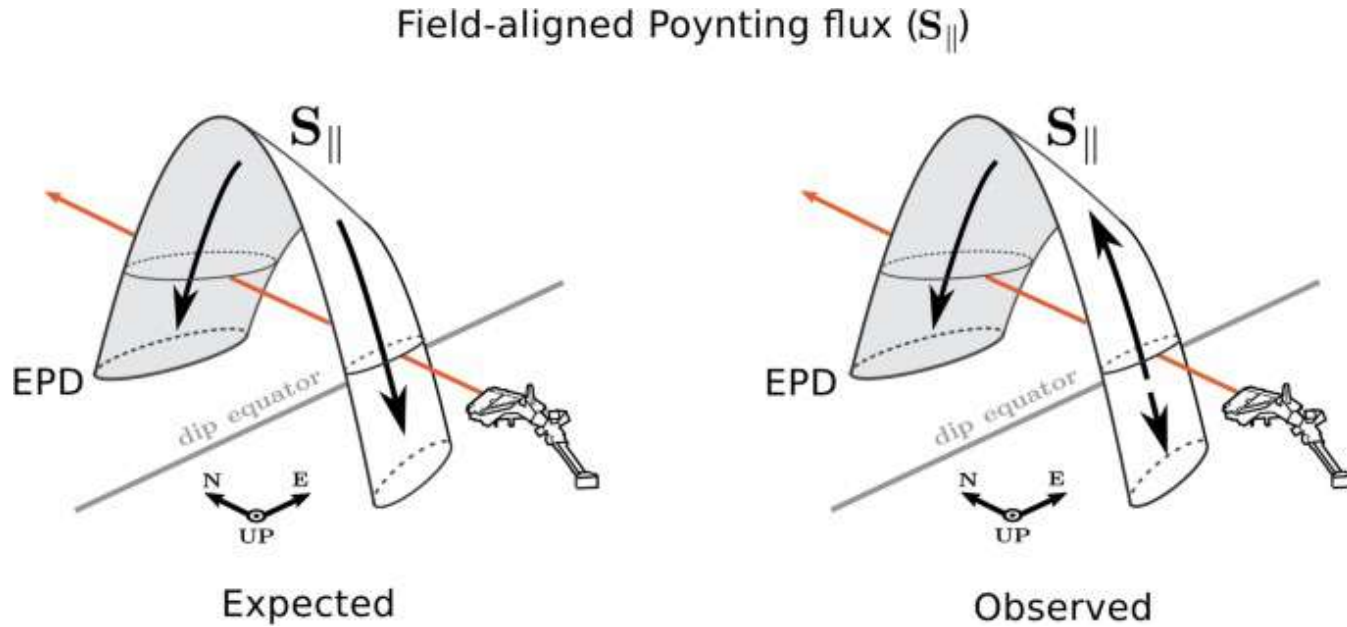
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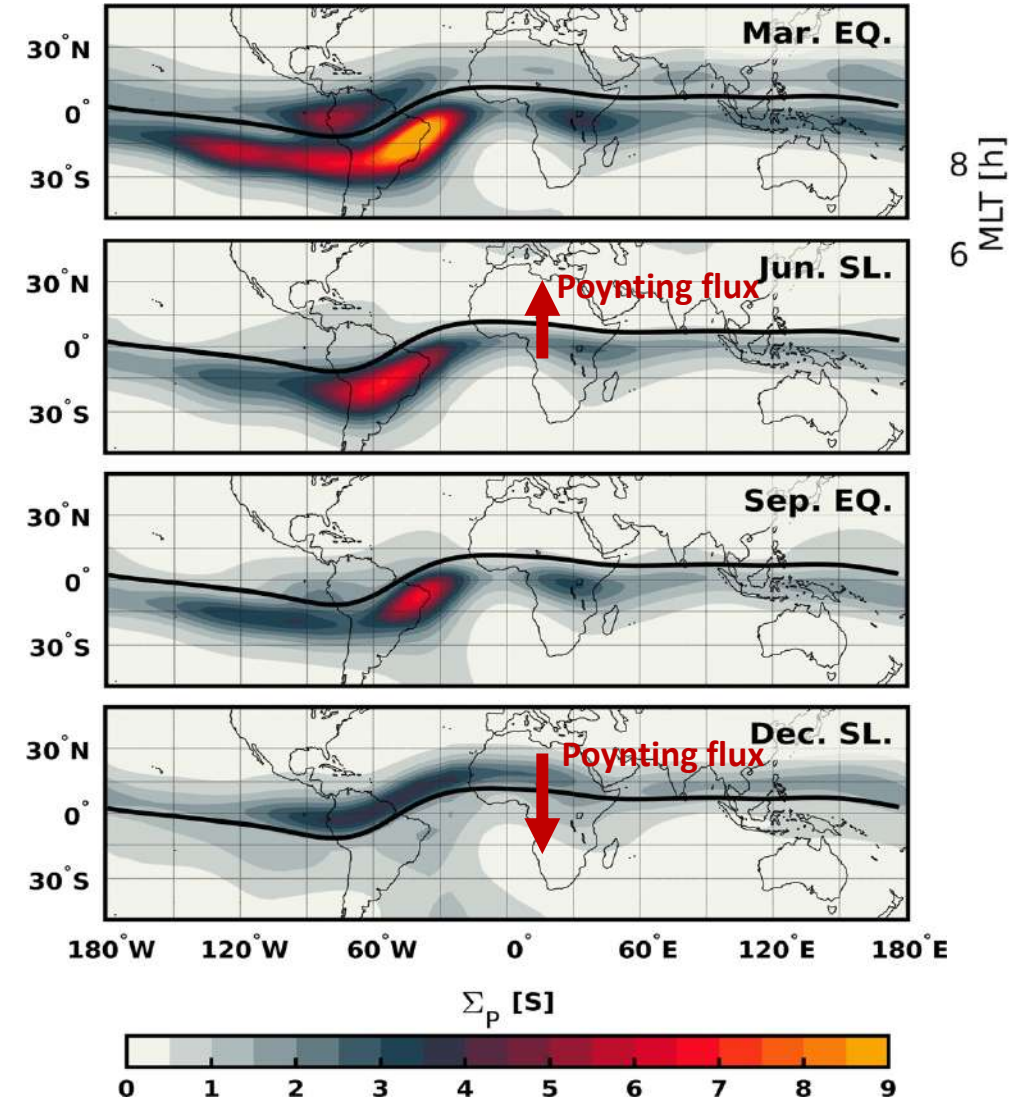
- **Observations:** Swarm mainly observes uni-directional Poynting flux



Determining the location of the dynamo source

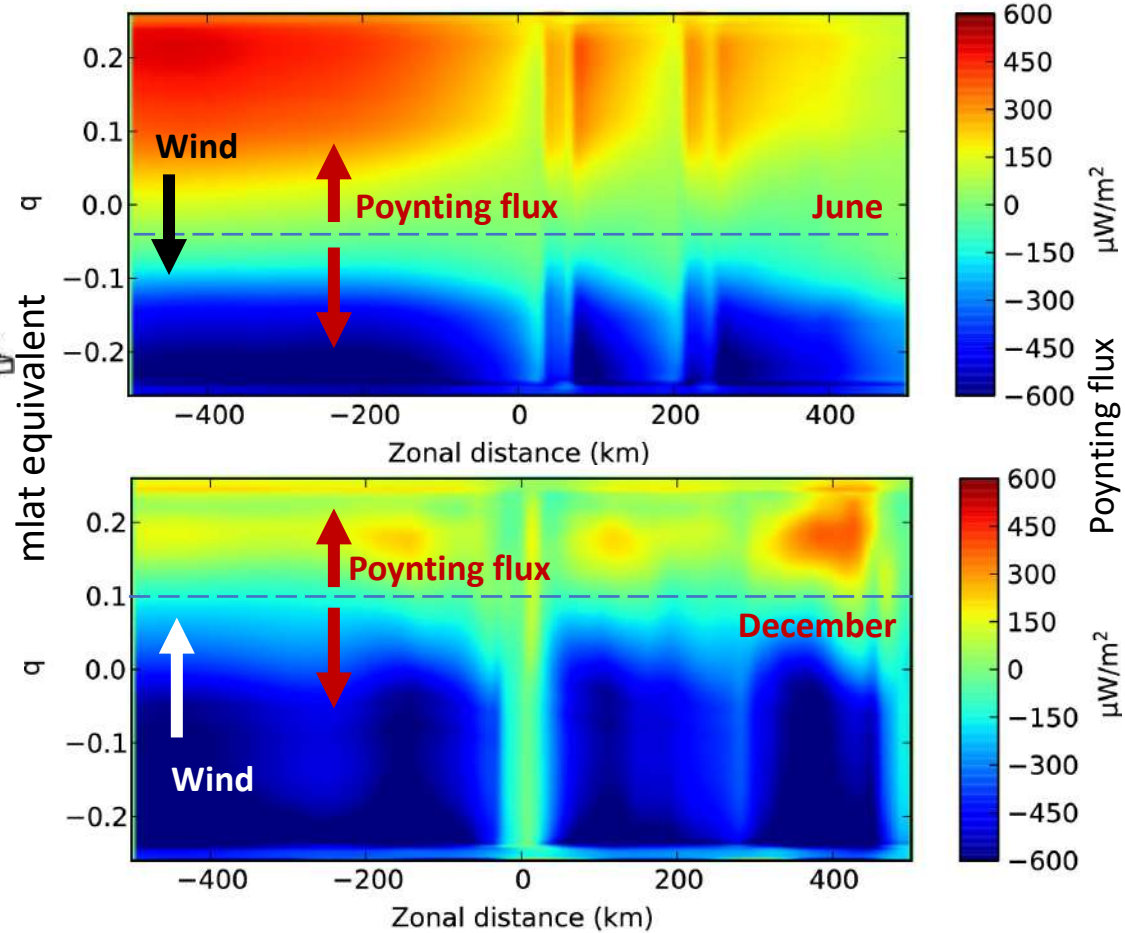
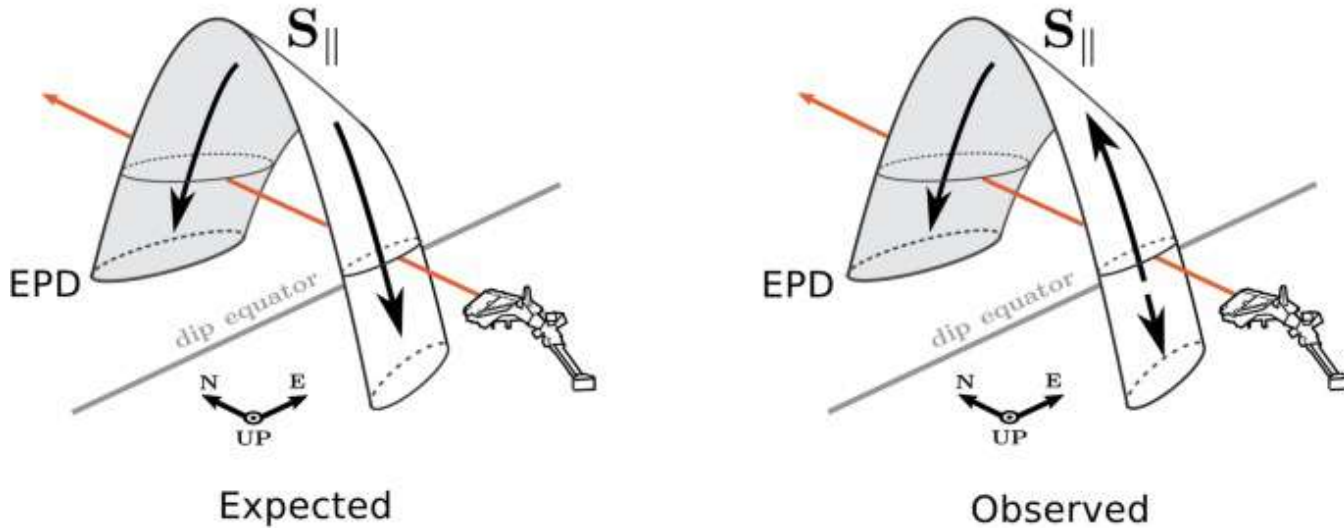


- **Observations:** Swarm mainly observes uni-directional Poynting flux
- **IGRF/IRI/MSIS:** Depending on season, conductivity is not symmetric (meridional winds expected to be the cause)
 - The dynamo source may be off-equator



Determining the location of the dynamo source

Field-aligned Poynting flux (S_{\parallel})



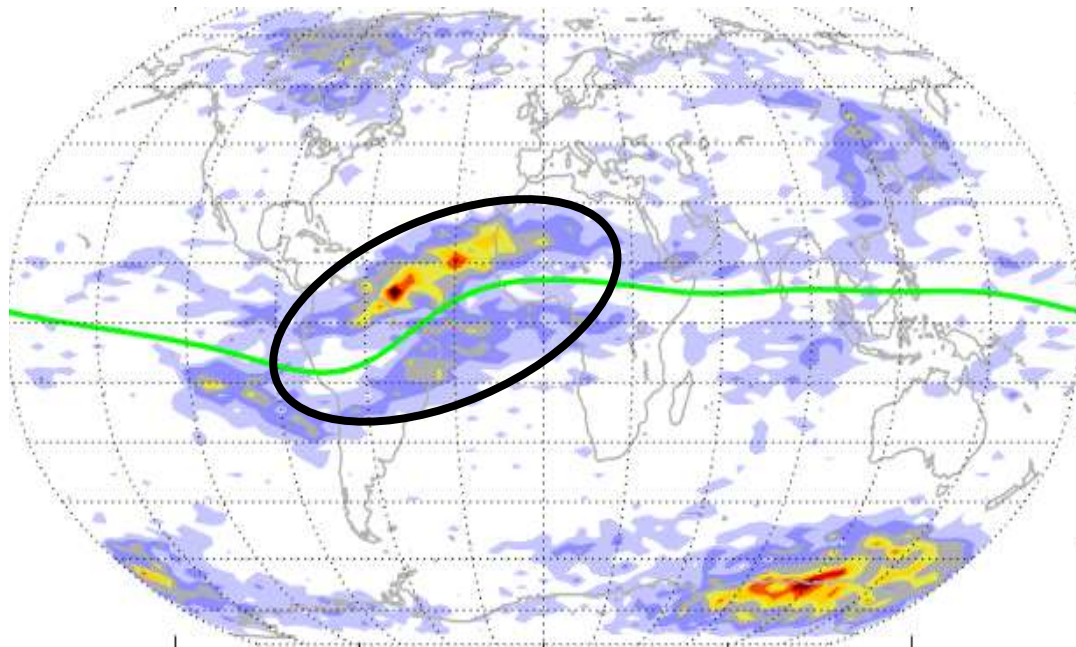
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 - The dynamo source may be off-equator
- **Cornell physical model:** meridional winds determine the “separatrix”

Rodríguez-Zuluaga et al., EPS, 2022; SciRep, 2019.

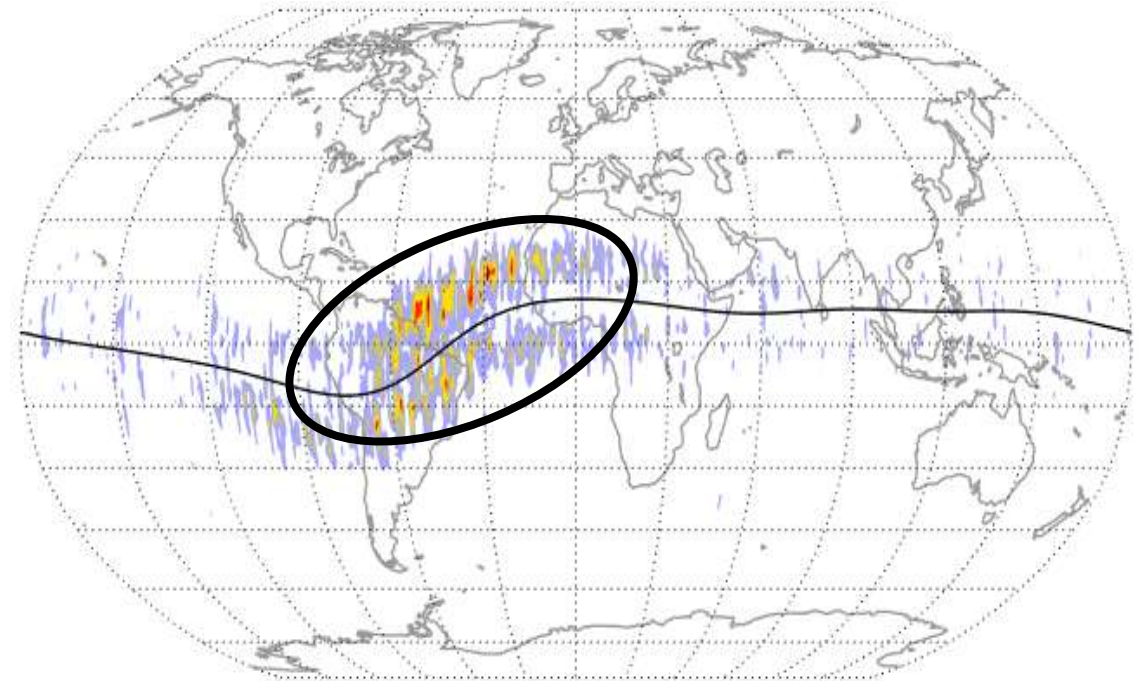
Co-located observations of plasma depletion and GPS observables

Swarm C 12/2013 – 11/2016

Occurrence rate of loss of GPS signal



Plasma irregularities (total count of positive IBI index)



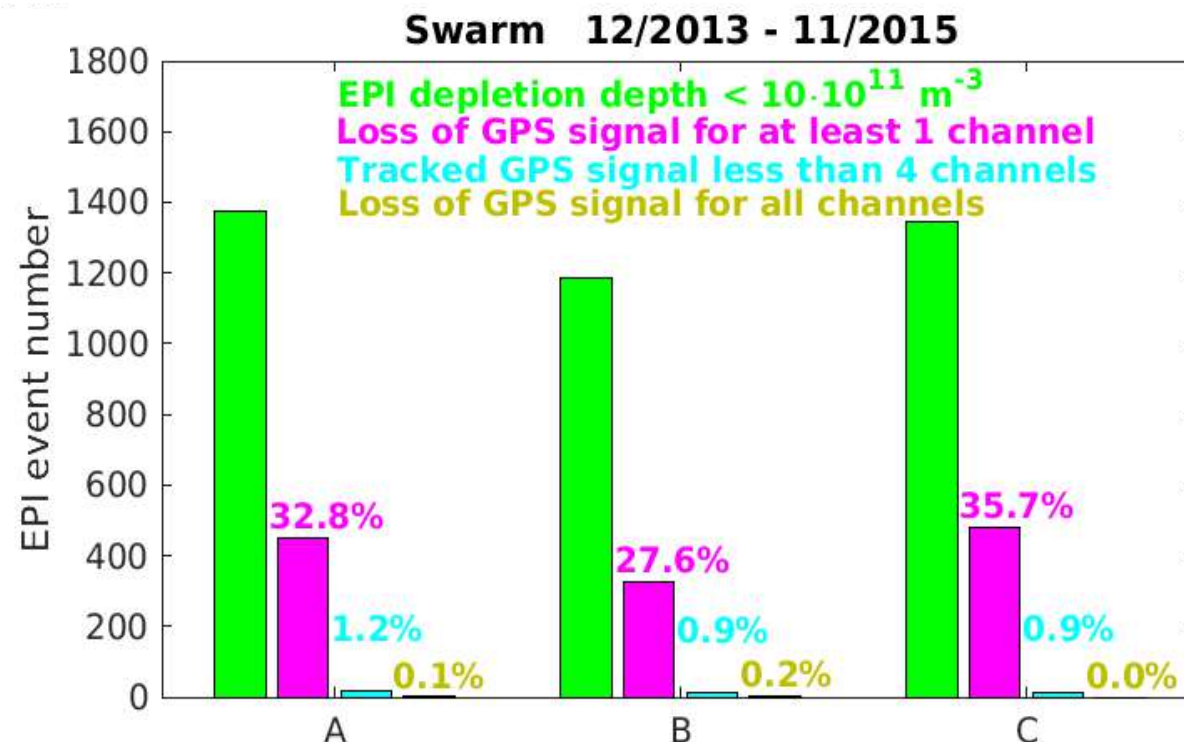
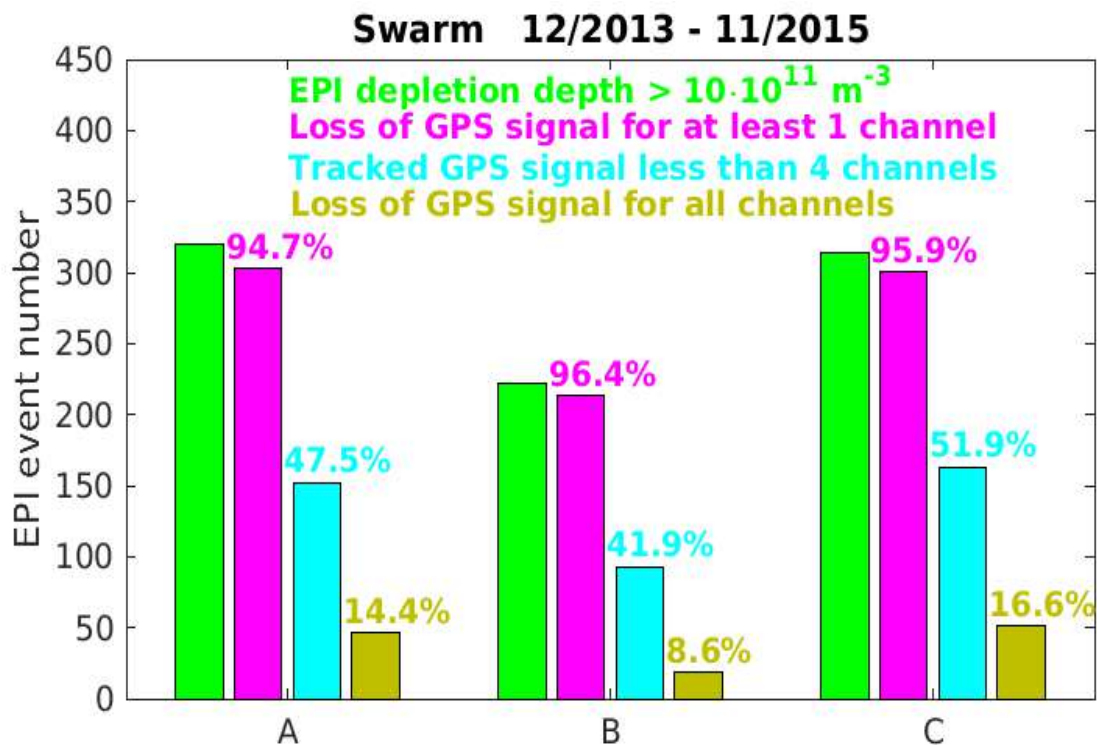
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Co-located observations of plasma depletion and GPS observables

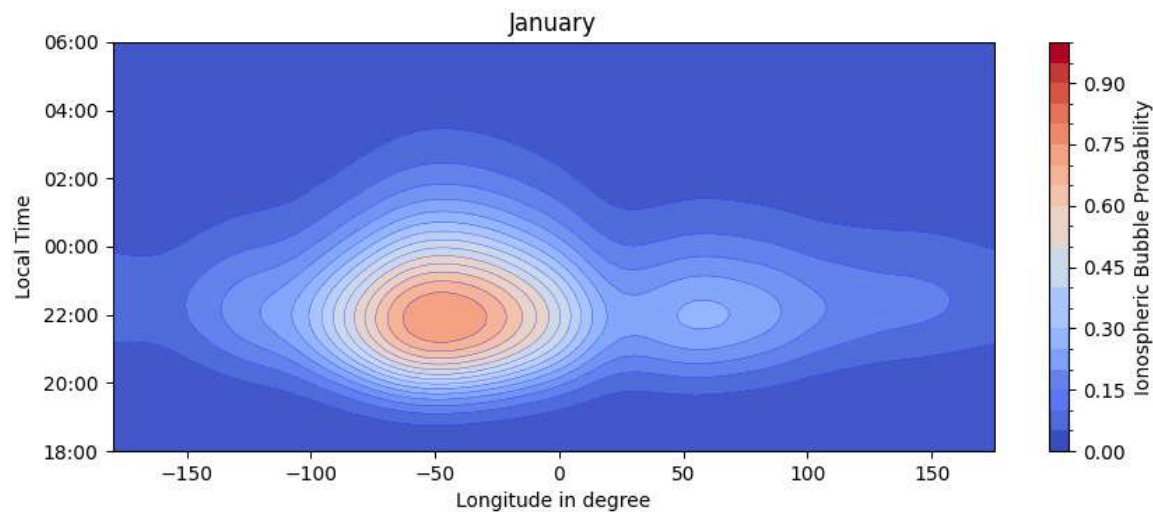
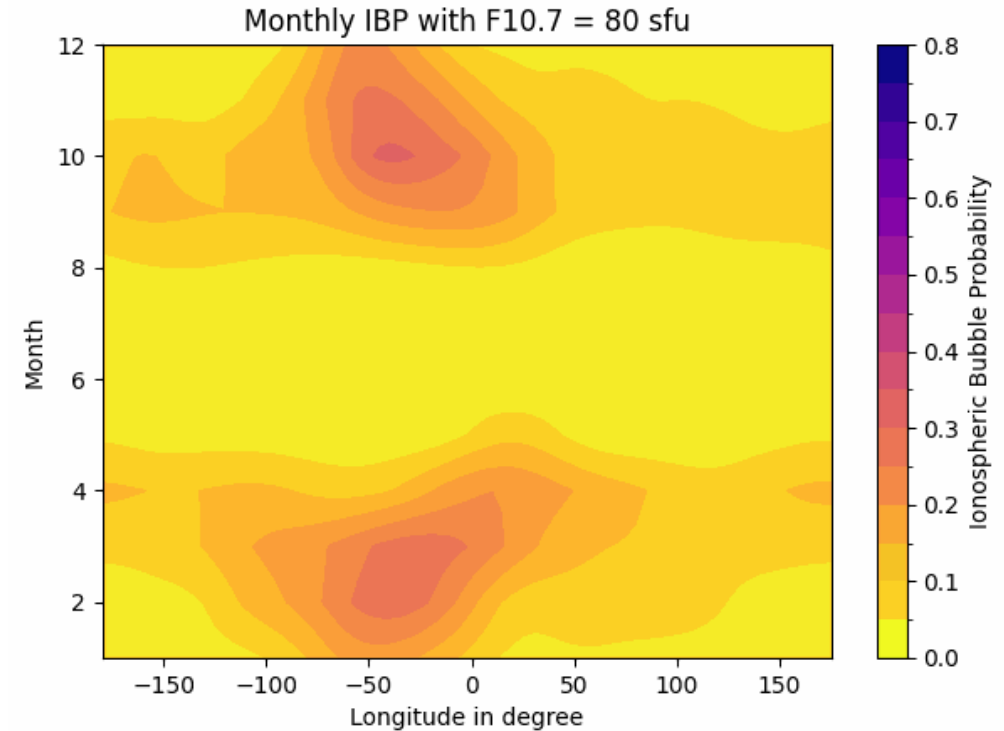
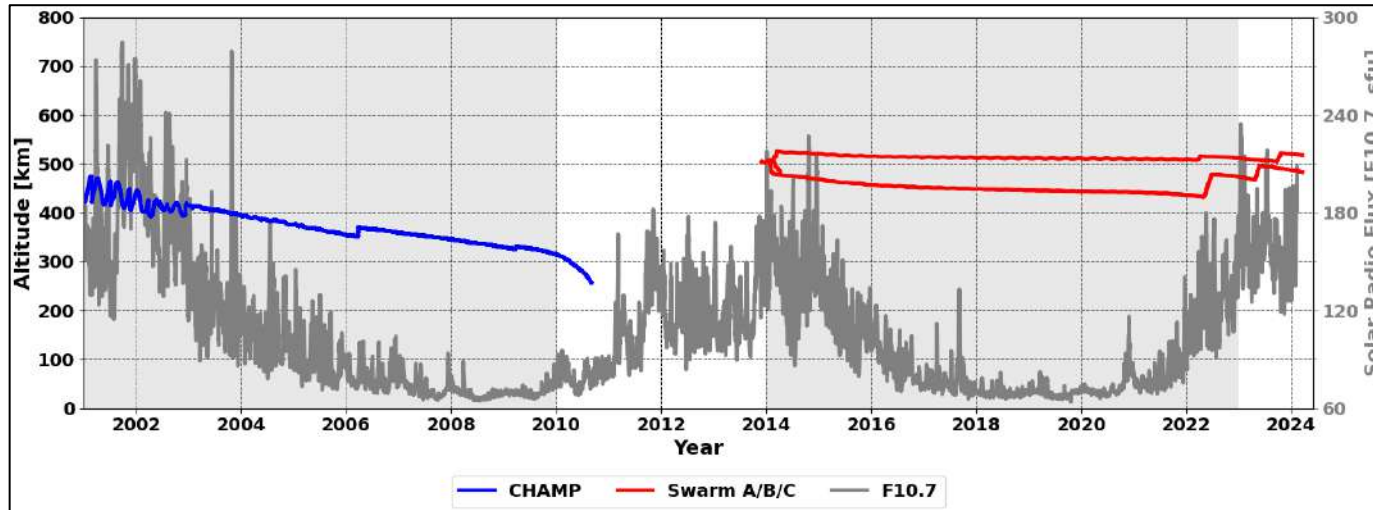
GPS loss occurrence for depletions $> 10 \times 10^{11} \text{ m}^{-3}$
 (For **steep** plasma density gradients)

GPS loss occurrence for depletions $< 10 \times 10^{11} \text{ m}^{-3}$
 (For **shallow** plasma density gradients)

Xiong et al., Space Weather, 2016; AnGeo, 2018.



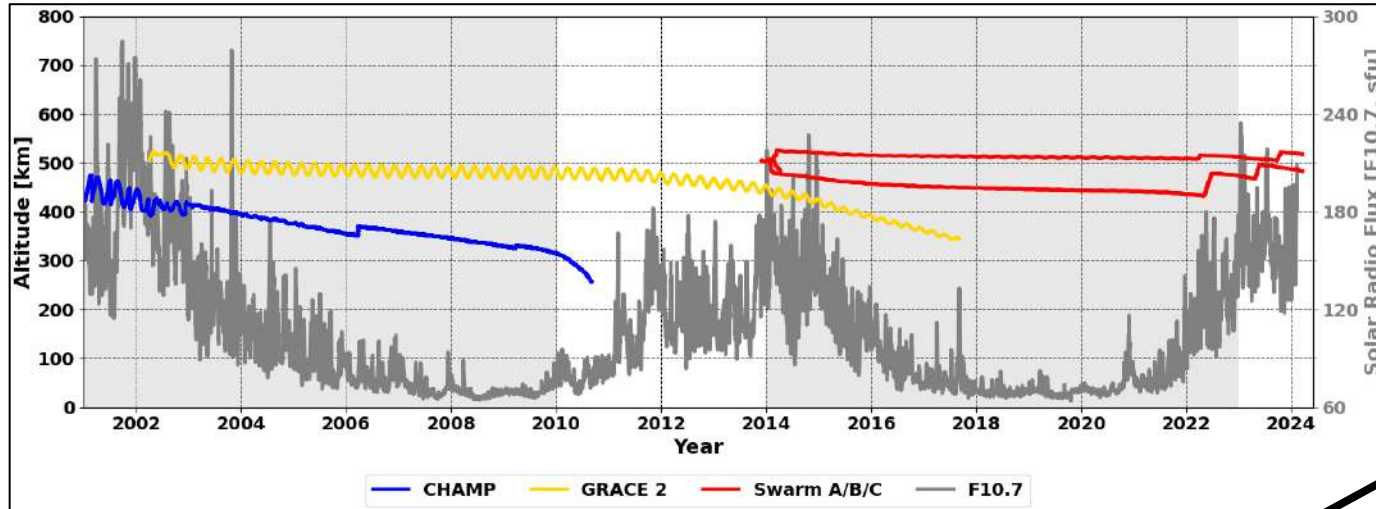
Statistical model of occurrence of irregularities



- Based on 9 years of CHAMP and 9 years of Swarm magnetic data (IBI)
- Estimate of probability (0-1) for a given local time, longitude, day of year and F10.7 solar flux
- <https://igit.iap-kborn.de/ibp/ibp-mode>

Stolle et al., Space Weather, under review, 2024.

Statistical model of occurrence of irregularities: validation



GRACE Ne: 2002 – 2017, $\pm 20^\circ$ mlat, 18-04LT,
F10.7 > 80, Hp30 < 3

A forecast for a binary event by a probabilistic model is better the larger the area, A , under the *ROC* curve is:

$$GC = 2 A - 1$$

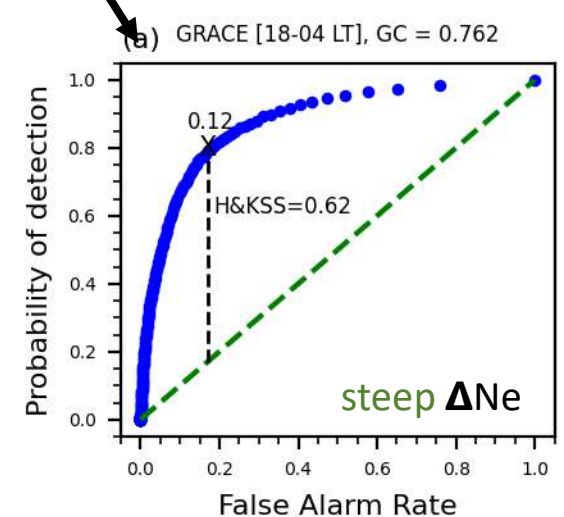
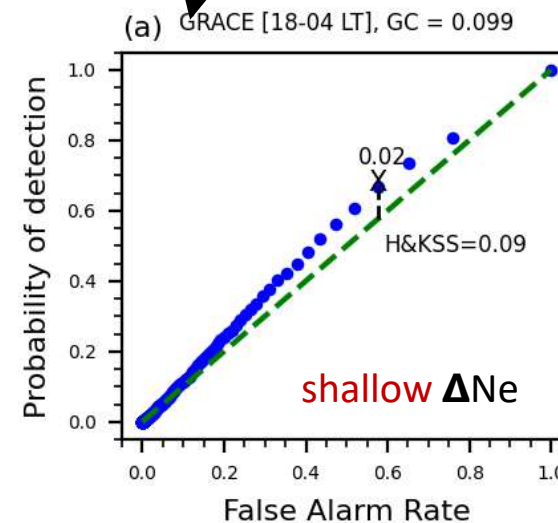
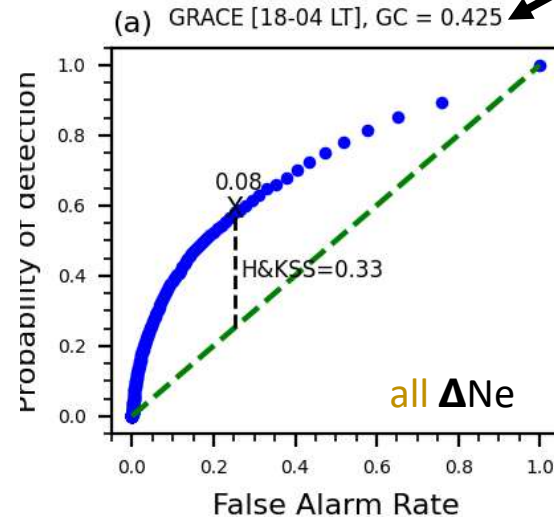
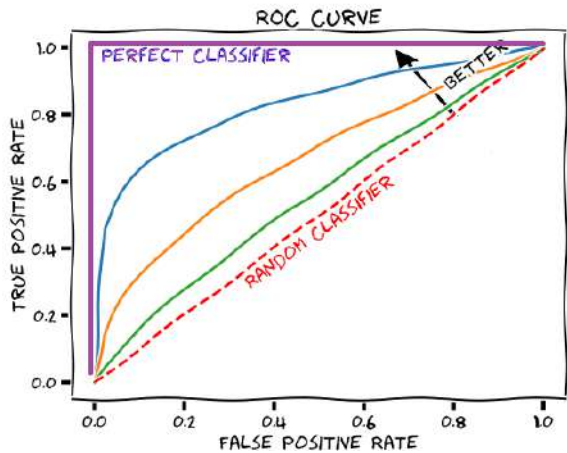
Good: GC= 1

Bad: GC=0

$$GC (\text{all } \Delta Ne) = 0.43$$

$$GC (\Delta Ne < 2 \cdot 10^5 \text{cm}^{-3}) = 0.10$$

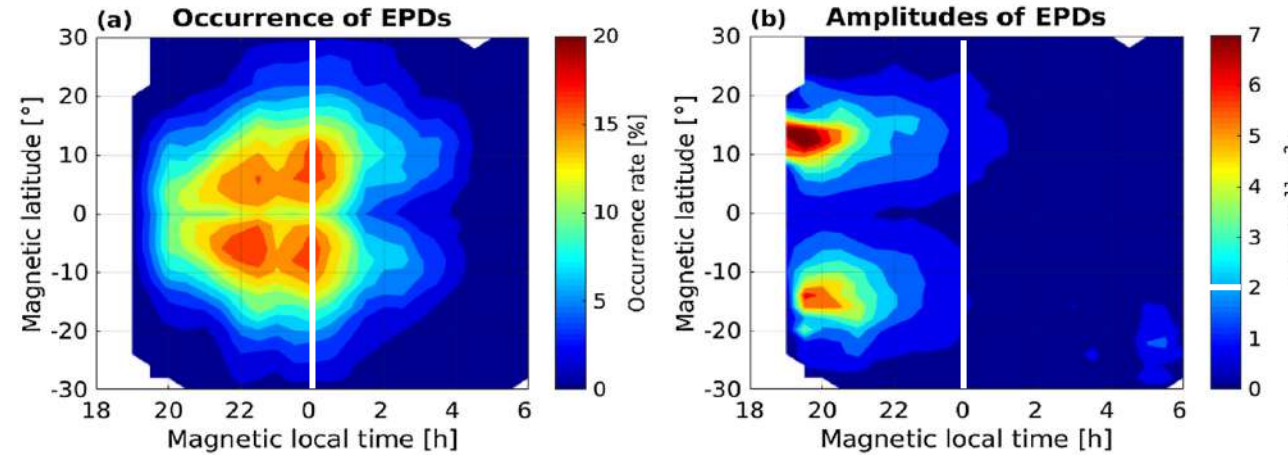
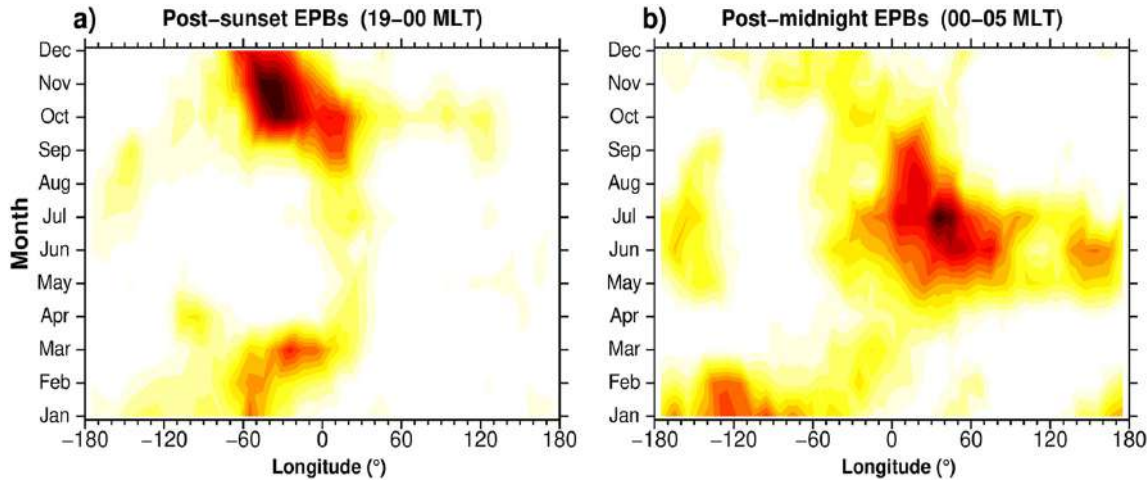
$$GC (\Delta Ne > 2 \cdot 10^5 \text{cm}^{-3}) = 0.76$$



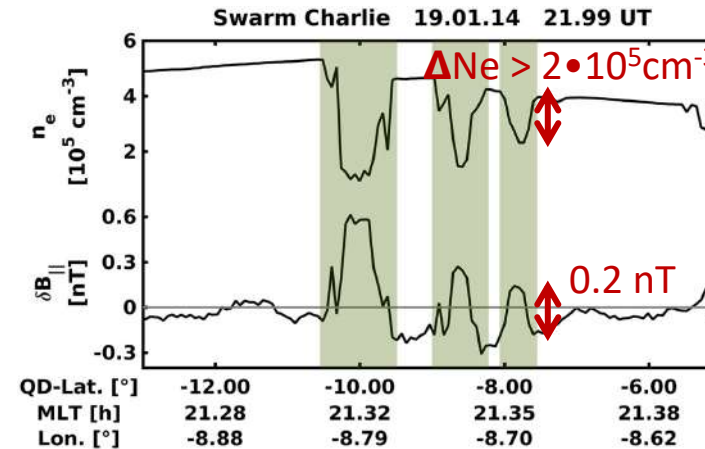
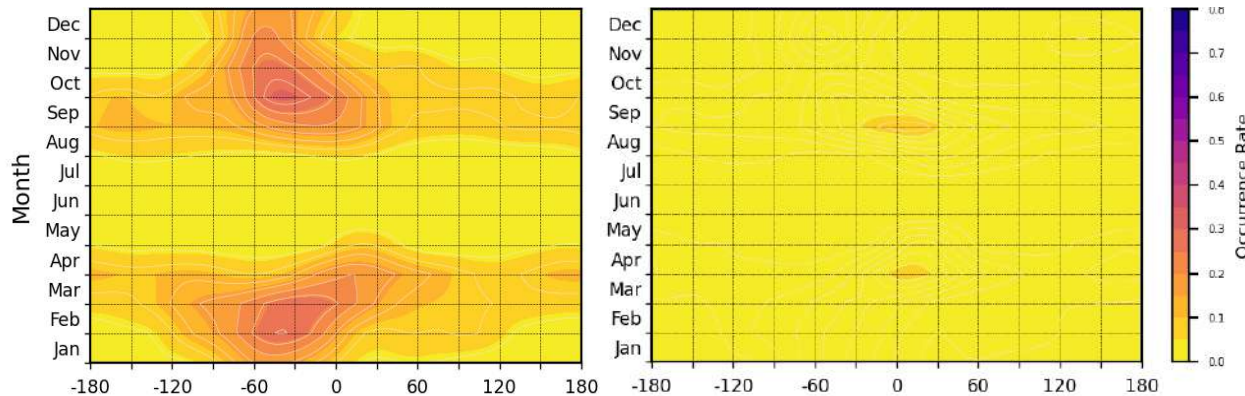
Post-midnight irregularities

Zakharenkova et al., *Space Weather*, 2023; **COSMIC-2**,
year 2021, F10.7 = 75 sfu, $\Delta N_e > 0.6 \cdot 10^5 \text{cm}^{-3}$

Wan et al., *JGR*, 2018; **Swarm A**,
2013-2017, Kp < 3, $\Delta N_e > 0.5 \cdot 10^5 \text{cm}^{-3}$



Pre-midnight **Swarm/CHAMP statistical model (IBI)** Post-midnight
F10.7 = 80 sfu

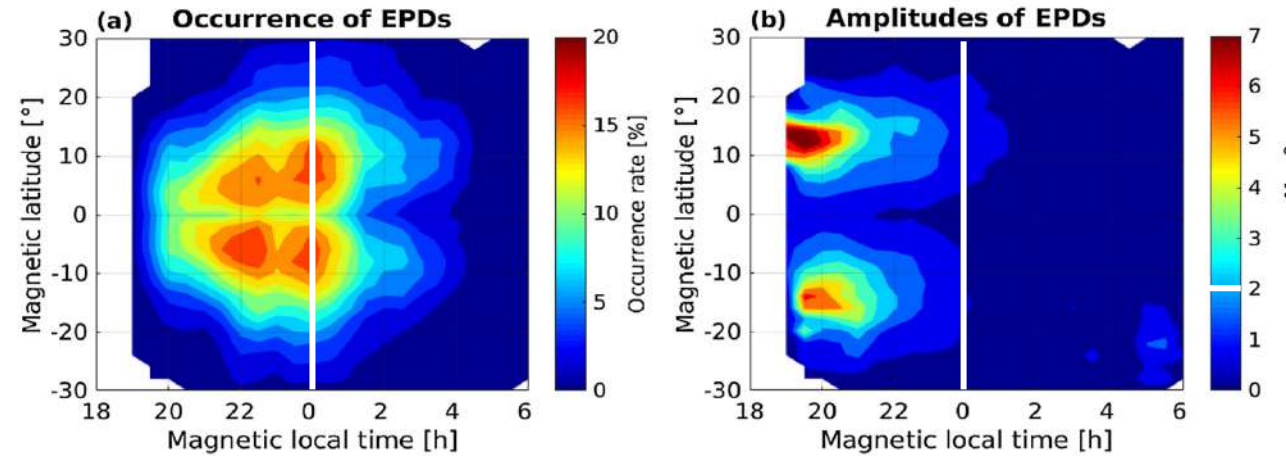
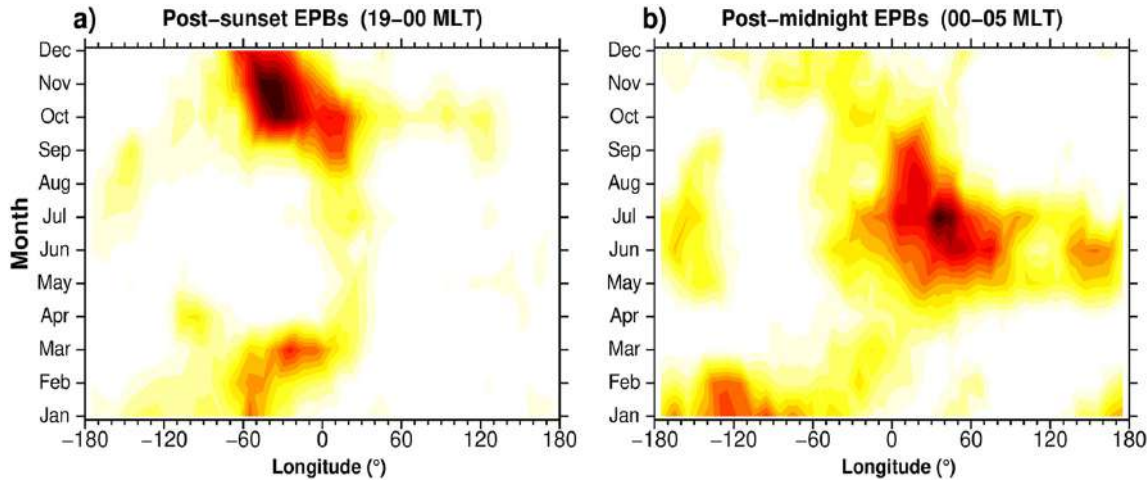


- The statistical model is sensitive to depletions of $\Delta N_e > 2 \cdot 10^5 \text{cm}^{-3}$
- Relates to magnetic signatures and GNSS scintillations!

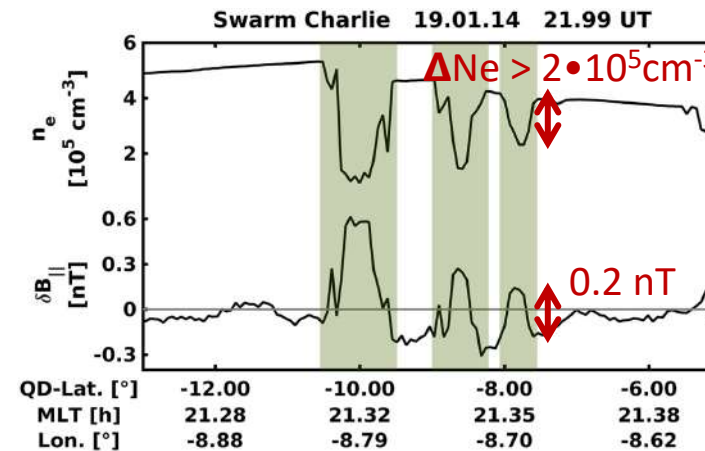
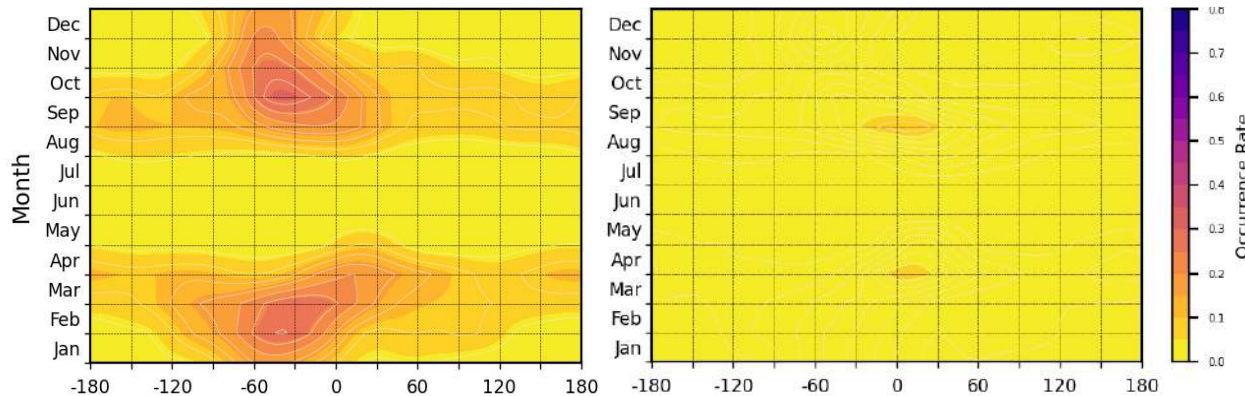
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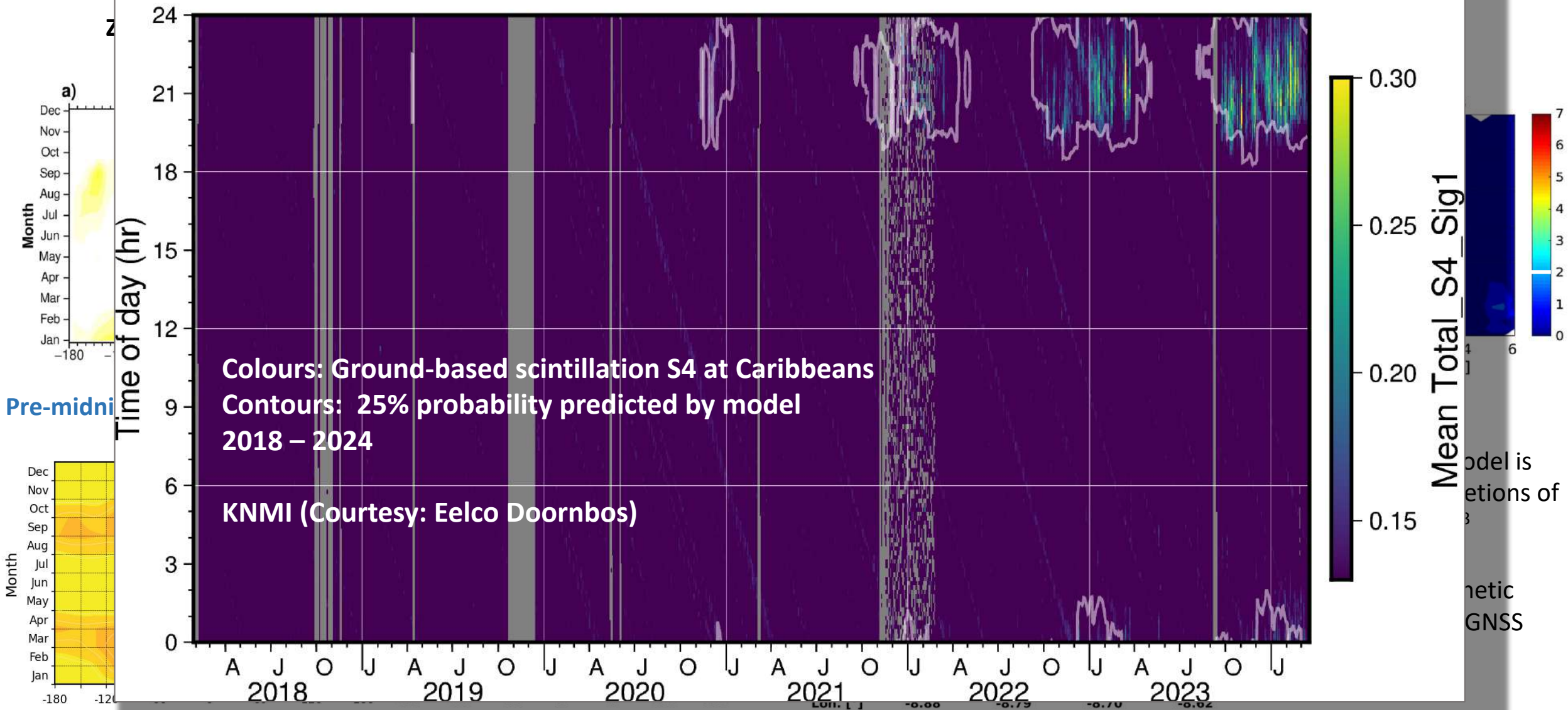


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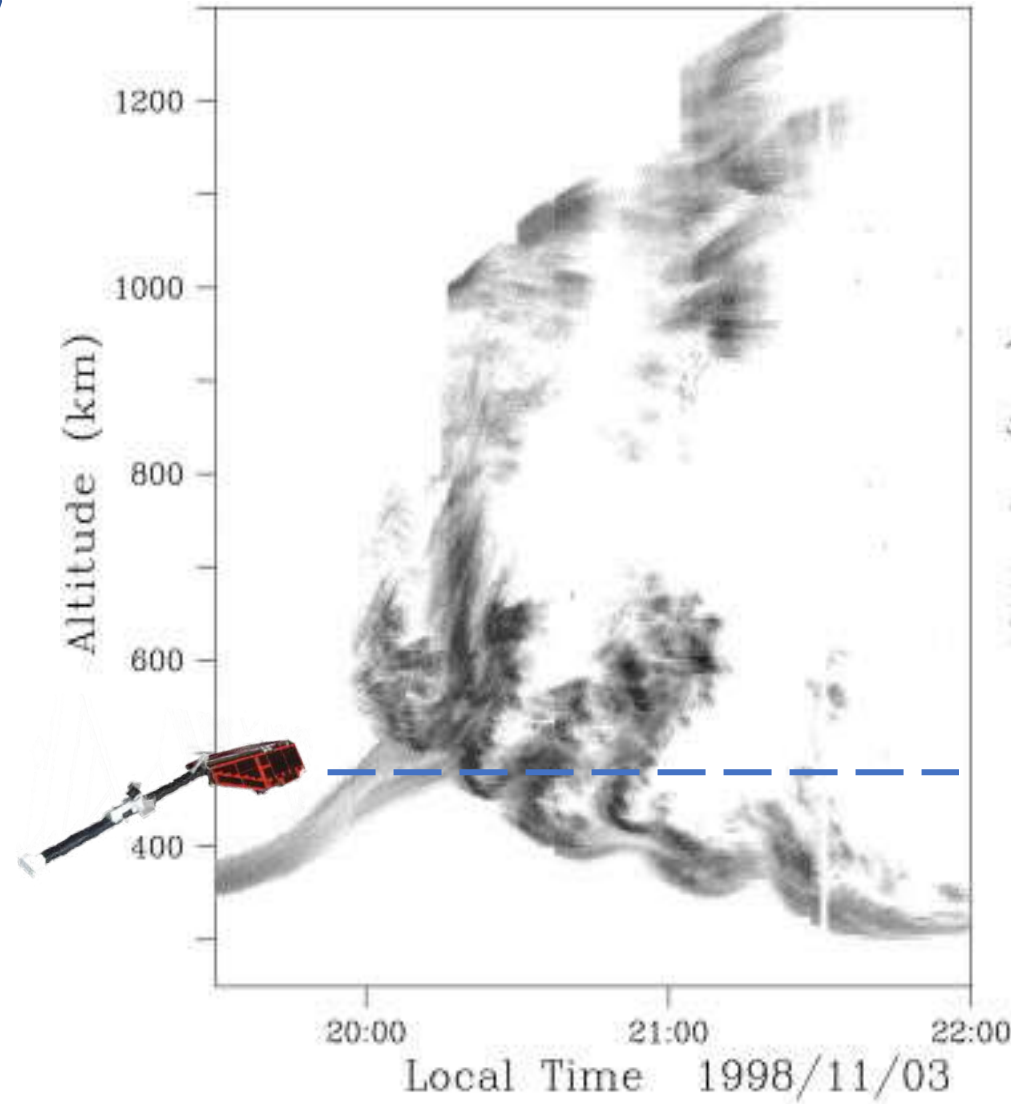
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SABA Mean Total_S4_Sig1 with IBP model 25% contours



Summary

- The **Swarm** mission **convince**s with its multi-parameter suite of high-precision instruments, its constellation and long time operation!
- The mission has led to **characterising** equatorial plasma depletions and their effects in **multiple** disciplines

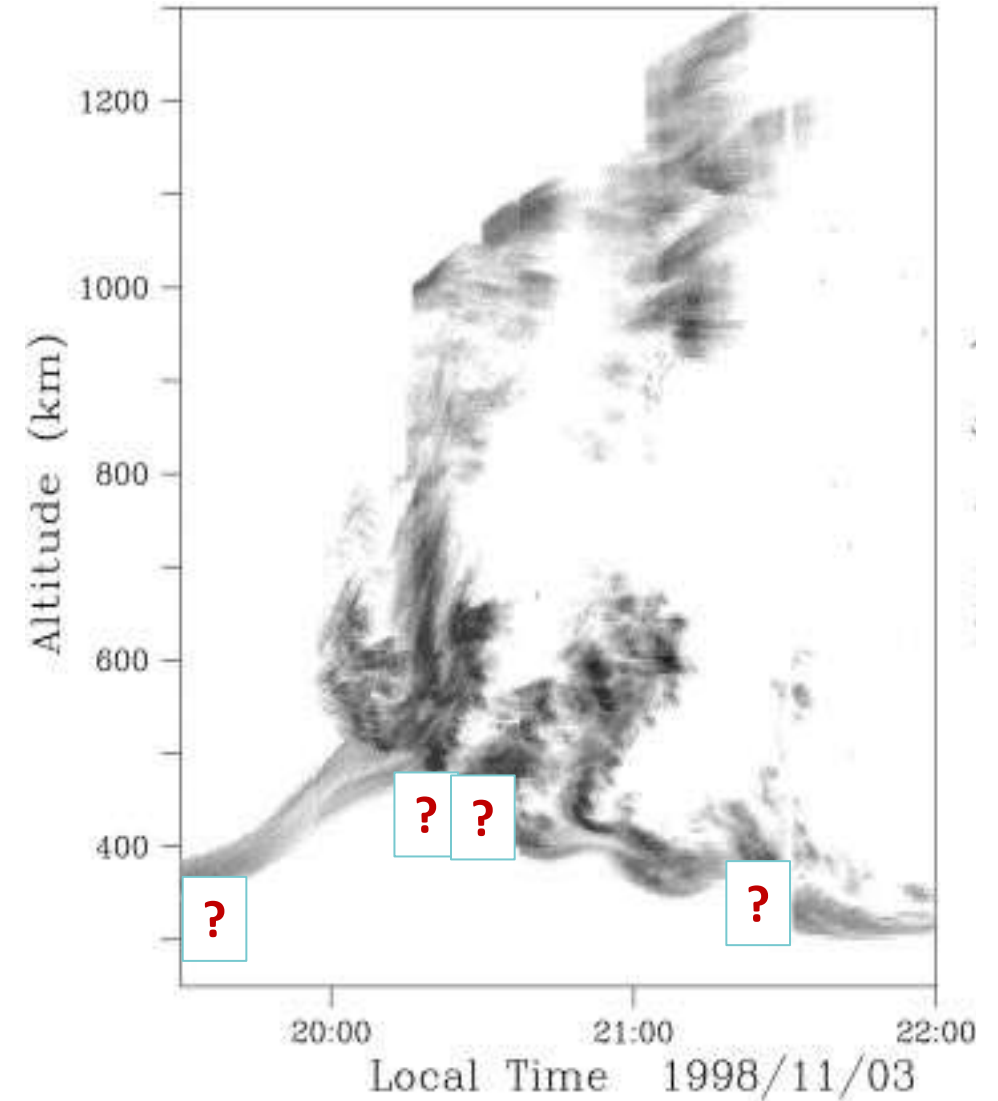


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Remaining targets:

- Explore **seeding** mechanisms of plasma irregularities in **combination** with low-inclination orbit satellites and remote sensing data, such as, the role of upward propagating atmospheric gravity waves



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Remaining targets:

- Explore **seeding** mechanisms of plasma irregularities in **combination** with low-inclination orbit satellites and remote sensing data, such as, the role of upward propagating atmospheric gravity waves
- Forecasting ionospheric **scintillation** related to these depletions on GNSS or radar applications at **ground** and **space**.

