

# Latest developments and upcoming innovations in Py4CATS - Python for Computational Atmospheric Spectroscopy

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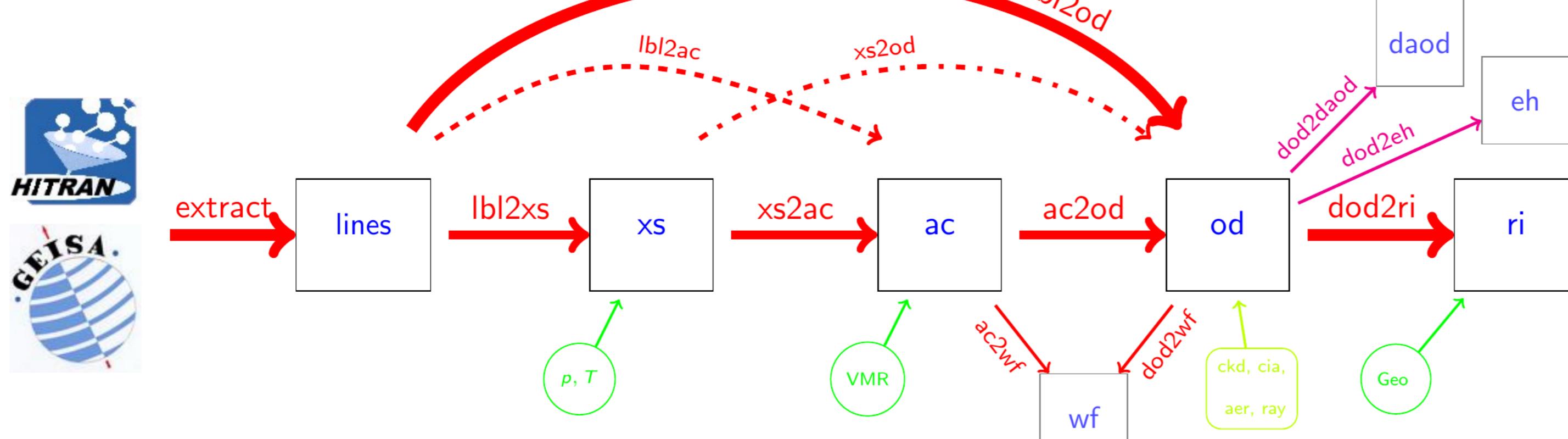
## Introduction

- ▶ High resolution IR- $\mu$ W atmospheric radiative transfer.
- ▶ Numeric/Scientific Python for optimized array processing.
- ▶ Rapid prototyping and testing of new algorithms.
- ▶ Study interaction of radiation with planetary atmospheres.
- ▶ Easy installation via pip after downloading the wheel file <https://atmos.eoc.dlr.de/tools/Py4CATS/>

## Radiative Transfer

Forward model setup (and design) using Py4CATS:

- ▶ Schwarzschild equation and Beer's law.
- ▶ Efficient and fast multi-grid line-by-line calculations.
- ▶ Spectroscopic databases such as HITRAN and GEISA.
- ▶ Various line shapes (e.g. Voigt, Rautian [2], ...).
- ▶ Continuum absorption (water vapor MT-CKD and CIA).
- ▶ Downlooking, uplooking, limb viewing geometries.
- ▶ Spherical or plane-parallel atmospheres.
- ▶ Passive and active sensors.



## Inversion

Retrieval algorithm for corresponding forward model:

- ▶ Infer atmospheric composition or/and temperature.
- ▶ Single or multiple spectral intervals.
- ▶ Weighted or classical least squares.
- ▶ Linear inversion schemes:
  - ▶ Singular Value Decomposition (SVD)
  - ▶ Matched Filter (MF) ...
- ▶ Nonlinear inversion schemes:
  - ▶ Variable Projection (VARPRO, separable least squares)
  - ▶ Levenberg-Marquardt (LM) ...

## Ongoing & Outlook

- ▶ Cloud and aerosol scattering (see Fig. 2).
- ▶ Atmospheric refraction.
- ▶ Jacobians via automatic differentiation.
- ▶ Parallelization to improve computational efficiency.
- ▶ Field-of-View instead of pencil-beam calculations.

## Molecular concentrations from SWIR

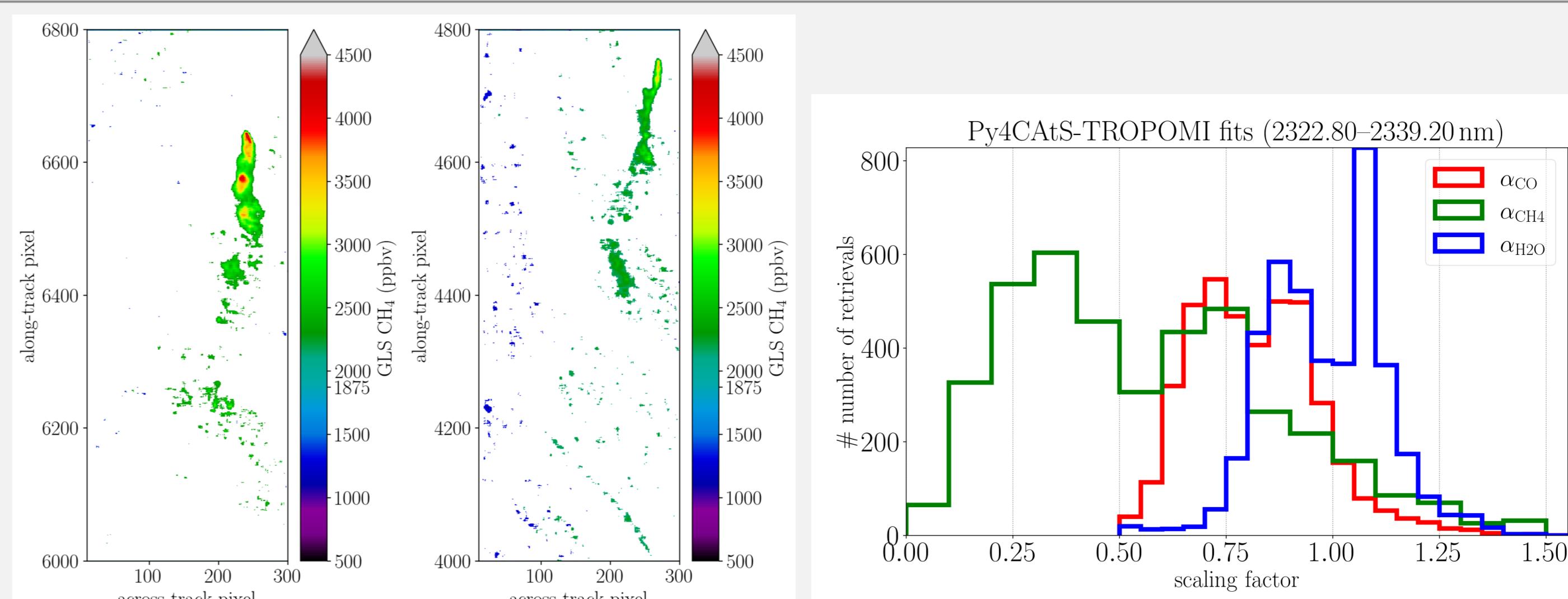


Figure 1: (Left) CH<sub>4</sub> plumes HySpex [1]. (Right) CO, CH<sub>4</sub>, and H<sub>2</sub>O fits TROPOMI.

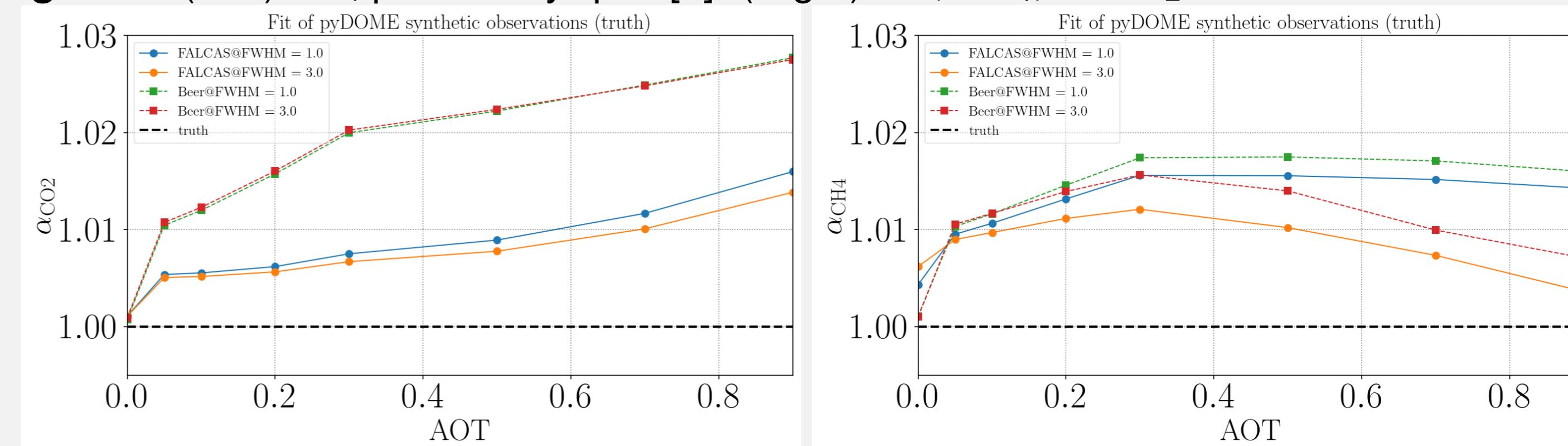


Figure 2: CO<sub>2</sub> and CH<sub>4</sub> fits with FALCAS accounting for aerosols vs. Beer model.

## Temperature profiles from TIR

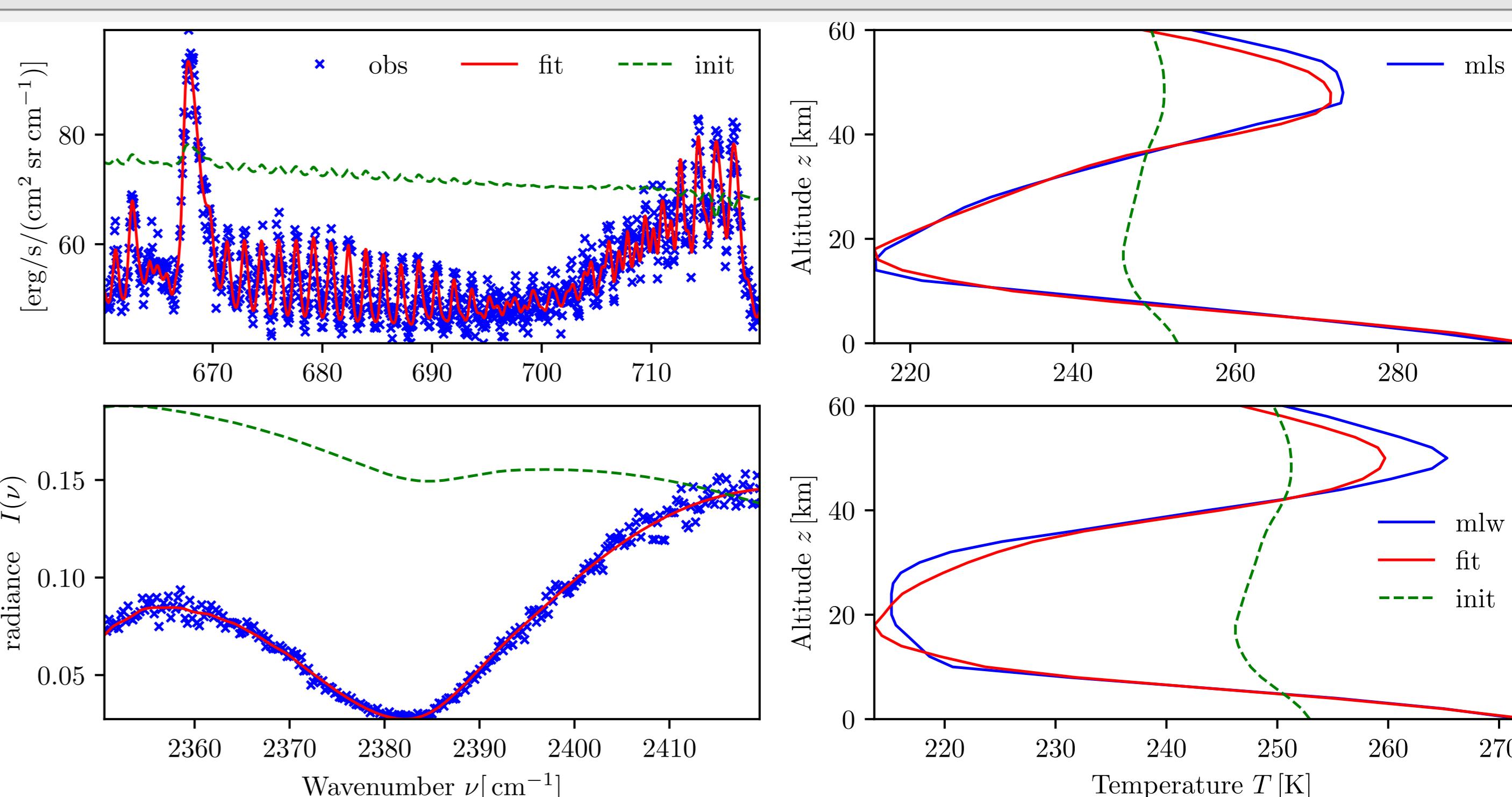


Figure 3: SVD to estimate temperatures from (exo-)planetary emission spectra [4].

## LIDAR and MIR Observations

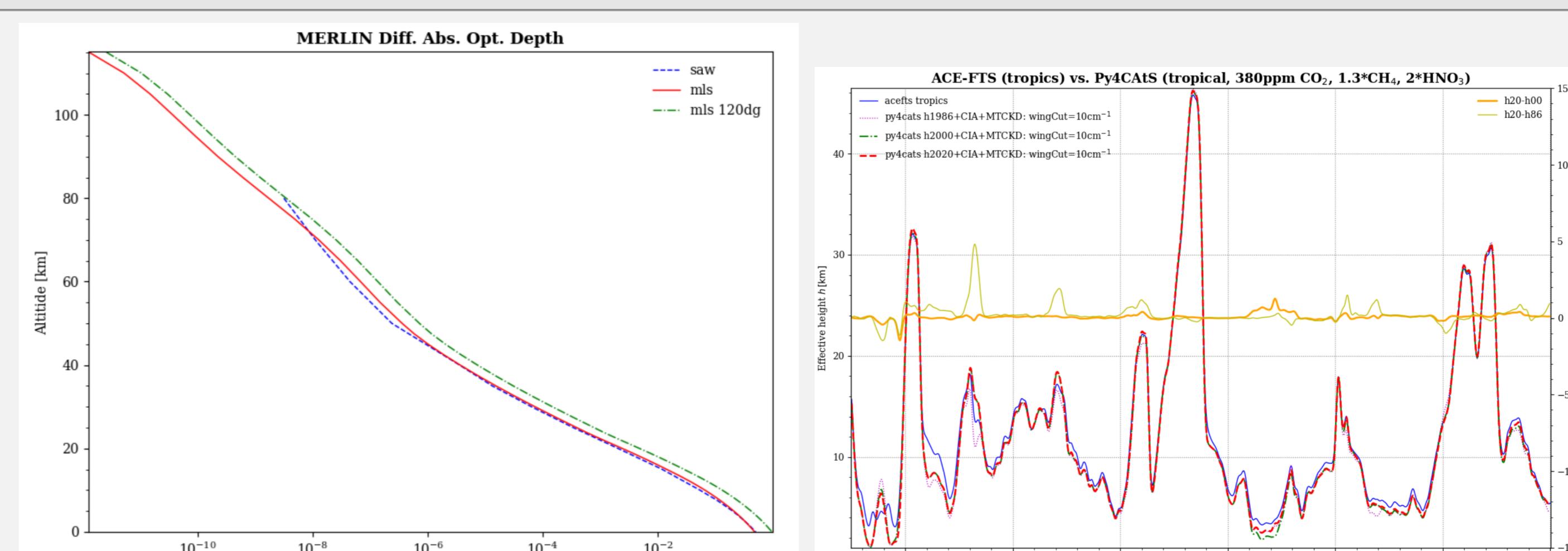


Figure 4: (Left) DAOD with center frequency  $\nu_{\text{on}}=1645.552 \text{ nm}$  and wing frequency  $\nu_{\text{off}}=1645.846 \text{ nm}$ . (Right) ACE-FTS effective heights from limb transmissions [3].

## References:

- [1] P. Hochstaffl, F. Schreier, C. H. Köhler, A. Baumgartner, and D. Cerra, Methane retrievals from airborne HySpex observations in the shortwave infrared. *Atmos. Meas. Tech.*, 16(18):4195–4214, Sept. 2023. doi: 10.5194/amt-16-4195-2023.
- [2] F. Schreier and P. Hochstaffl, Computational aspects of speed-dependent Voigt and Rautian profiles. *J Quant Spectrosc Radiat Transf.*, 258:107385, 2021. doi: 10.1016/j.jqsrt.2020.107385.
- [3] F. Schreier, S. Städter, F. Wunderlich, M. Godolt, and J. L. Grenfell, Transmission spectroscopy with the ACE-FTS infrared spectral atlas of Earth: A model validation and feasibility study. *Molec. Astrophys.*, 11:1–22, 2018. doi: 10.1016/j.molap.2018.02.001.
- [4] F. Schreier, S. Städter, F. Wunderlich, M. Godolt, and J. L. Grenfell, SVEEETIES: Singular vector expansion to estimate Earth-like exoplanet temperatures from infrared emission spectra. *A&A*, 633:A156, Jan. 2020. ISSN 0004-6361, 1432-0746. doi: 10.1051/0004-6361/201936511.