



# L2 BM-RAD and BMA-FLX products verification

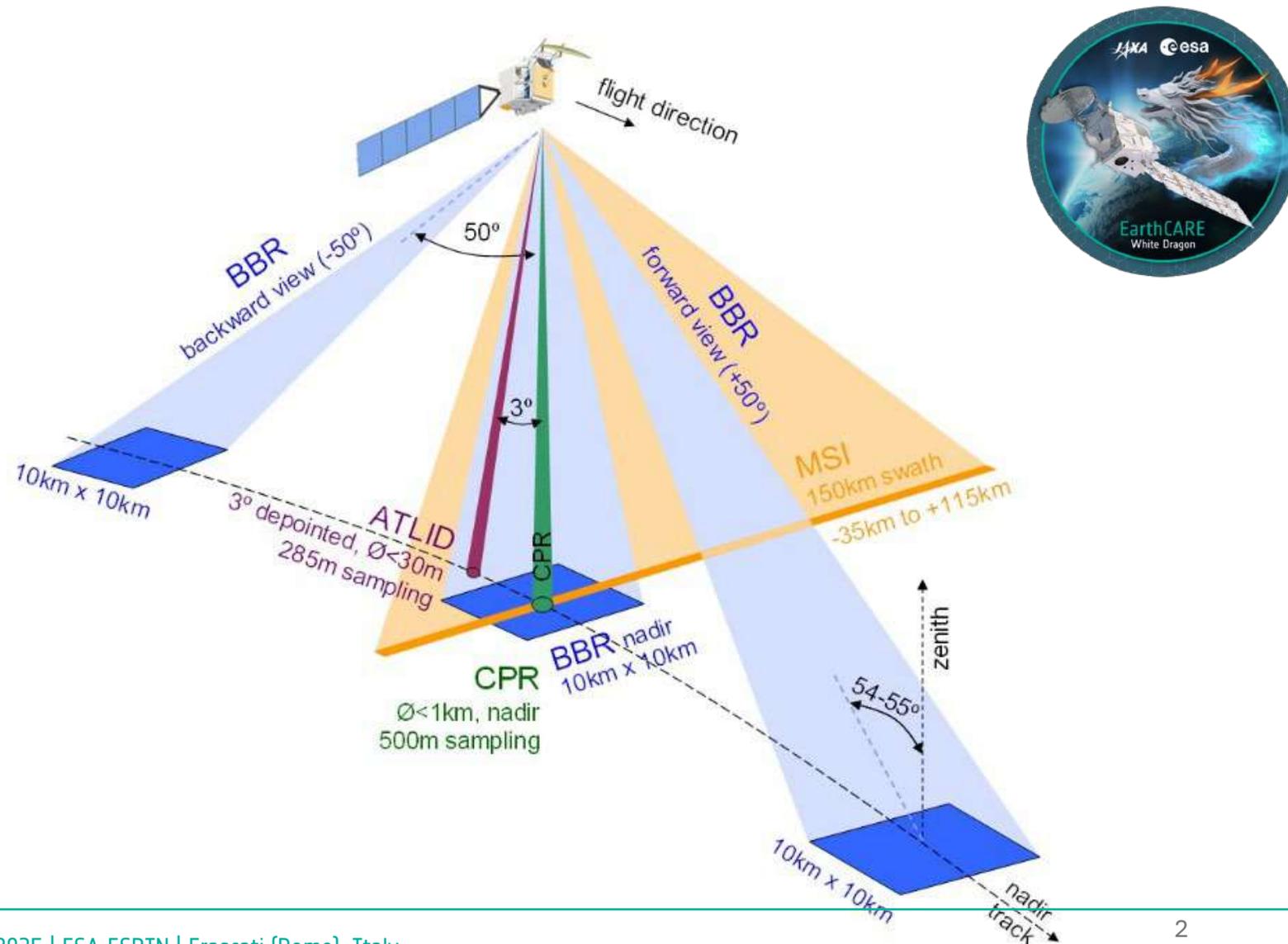
*Almudena Velázquez, Christine Aebi, Edward Baudrez, Nicolas Clerbaux – RMIB  
Carla Salas, Carlos Domenech – GMV-ES*

2<sup>nd</sup> ESA-JAXA EarthCARE In-Orbit Validation Workshop  
17 – 20 March 2025 | ESA-ESRIN | Frascati (Rome), Italy



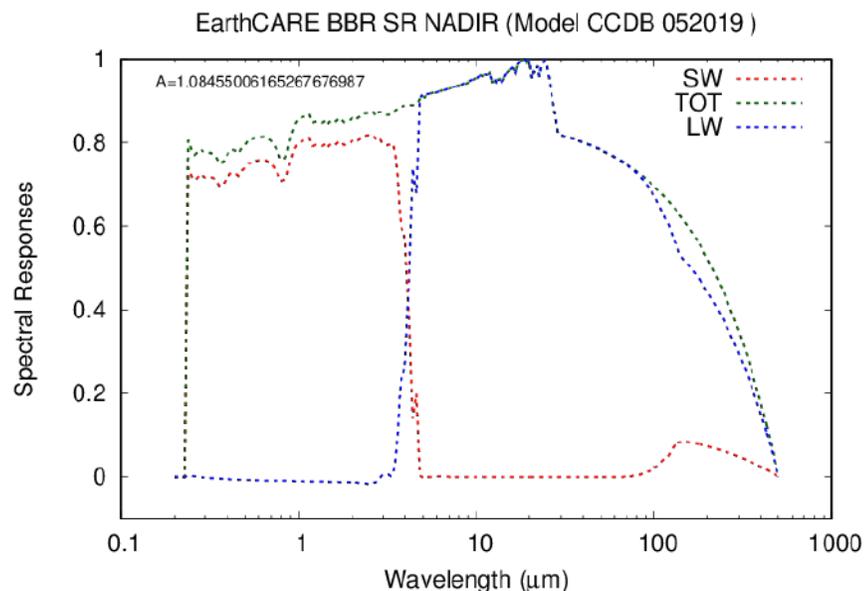
# Outline

1. BM-RAD and BMA-FLX
2. Product resolutions
3. Reference Level & coregistration
4. Latitudinal means
5. Satellite comparisons
  1. GERB-1 radiances
  2. CERES rad.& fluxes
6. Summary
7. Recommendations



## BM-RAD

- **BBR SW** and **TW** measurements are converted into **solar** and **thermal unfiltered radiances**
- Two unfiltering algorithms:
  - **Stand-alone** : BBR L1 and land use classification (SW & LW)
  - **SW MSI-based**: SW BBR L1 and MSI cloud mask and cloud phase



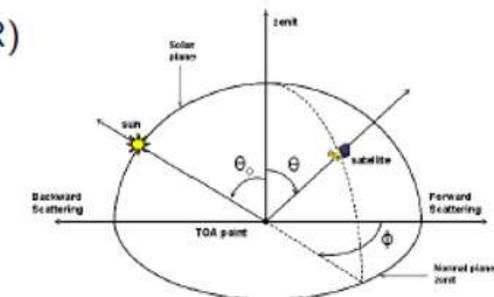
## BMA-FLX

- **BBR solar** and **thermal radiances** are converted to **fluxes**

$$F(\theta_0) = \int_{\theta=0}^{\frac{\pi}{2}} \int_{\phi=0}^{2\pi} L(\theta, \phi) \cos(\theta) \sin(\theta) d\theta d\phi$$

- ▶ Need of Angular Dependency Models (R)

$$F = \frac{\pi L(\theta, \phi)}{R(\theta, \phi)}$$



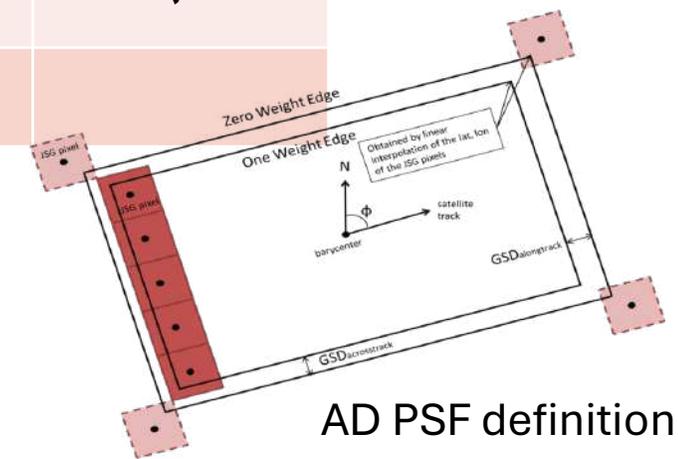
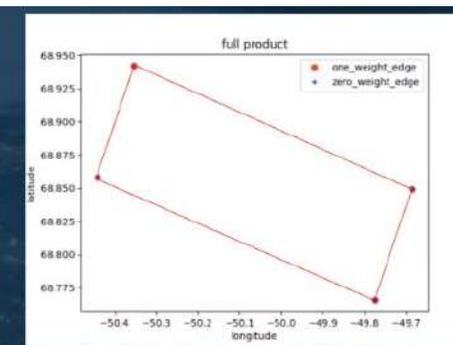
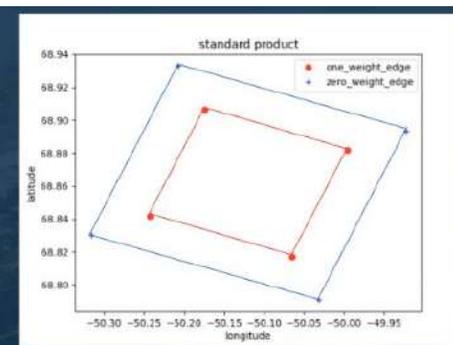
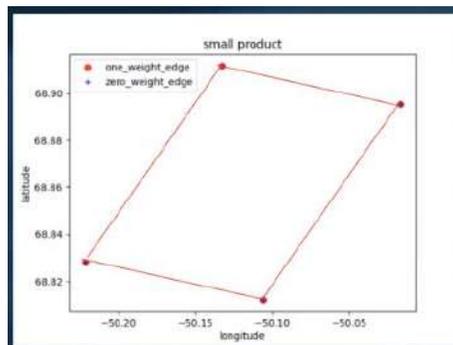
- In the LW:  $R(\theta) = a_0 + a_1 z_1 + a_2 z_2 + a_3 z_1^2 + a_4 z_1 z_2 + a_5 z_2^2$   
 $z_1 = BT_{10.8 \mu m}$  and  $z_2 = BT_{12.0 \mu m} - BT_{10.8 \mu m}$
- In the SW: feed-forward back-propagation ANN using CERES and MODIS data

Inputs: **BM-RAD**, **M-RGR**, **M-CM**, **M-COP**, **A-CTH**, **X-MET**

# BM-RAD / BMA-FLX product resolutions



L1 PSF	Size (across x along track)	Reference grid	L1 filtered radiances	BM-RAD unfiltered radiances	BMA-FLX fluxes	BMA-FLX combined flux
Standard	10x10 km <sup>2</sup>	BBR	B-NOM	✓	✓	✓
Small	5x10 km <sup>2</sup>	BBR	B-NOM	✓	✓	✓
Full	17x10 km <sup>2</sup> (nadir) 28x10 km <sup>2</sup>	BBR	B-NOM	✓	✓	✓
Assessment domain	5x21 JSG (configurable)	JSG	B-SNG	✓	✓	✓
JSG*	1x1 JSG (or 3x3 JSG)	JSG	B-SNG	✓	✓	✓



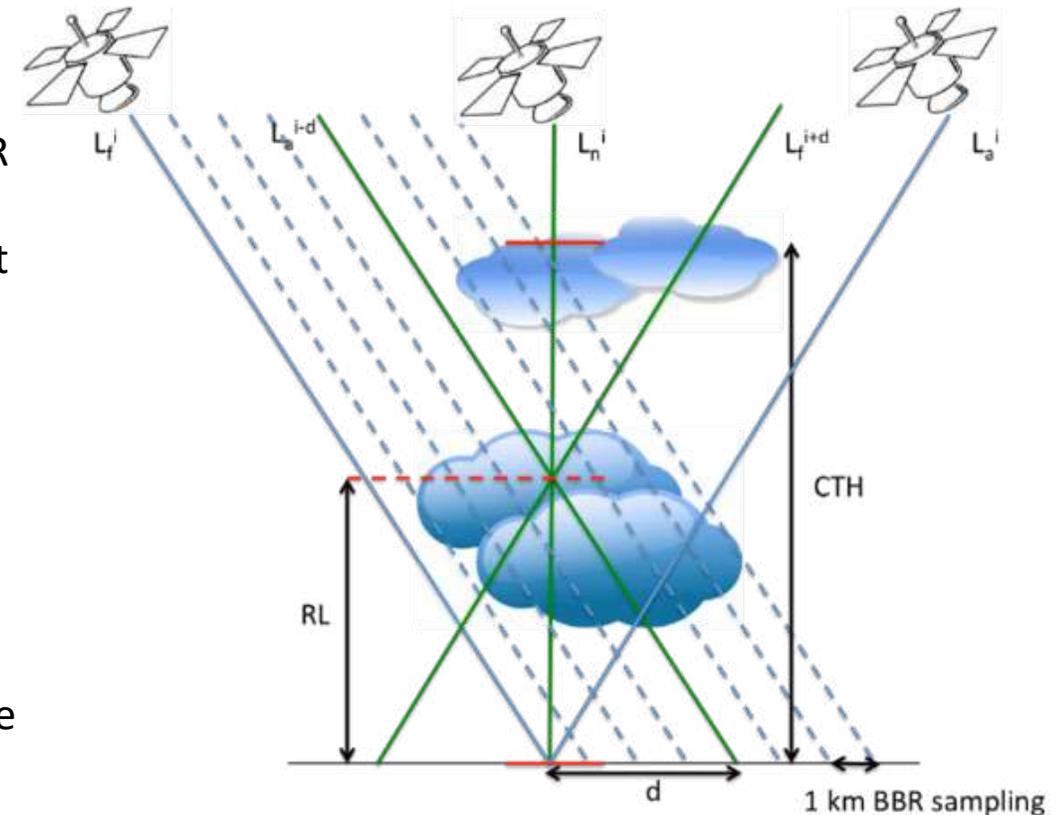
AD PSF definition

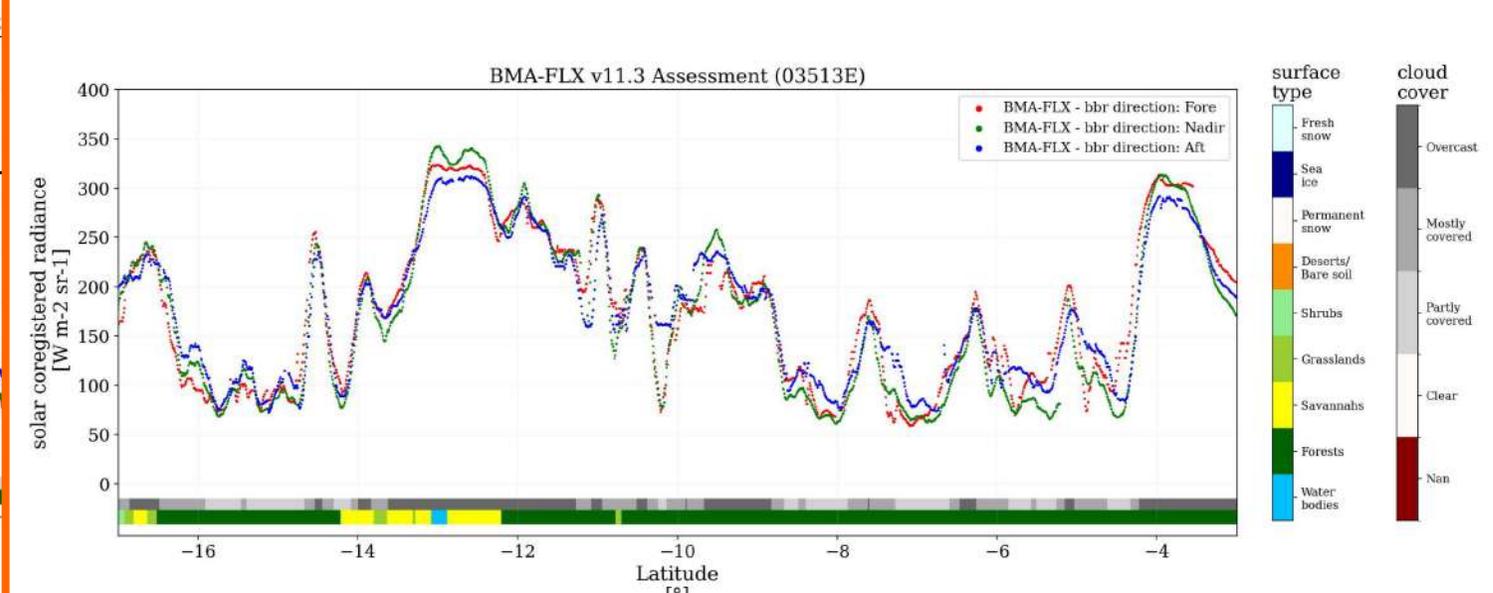
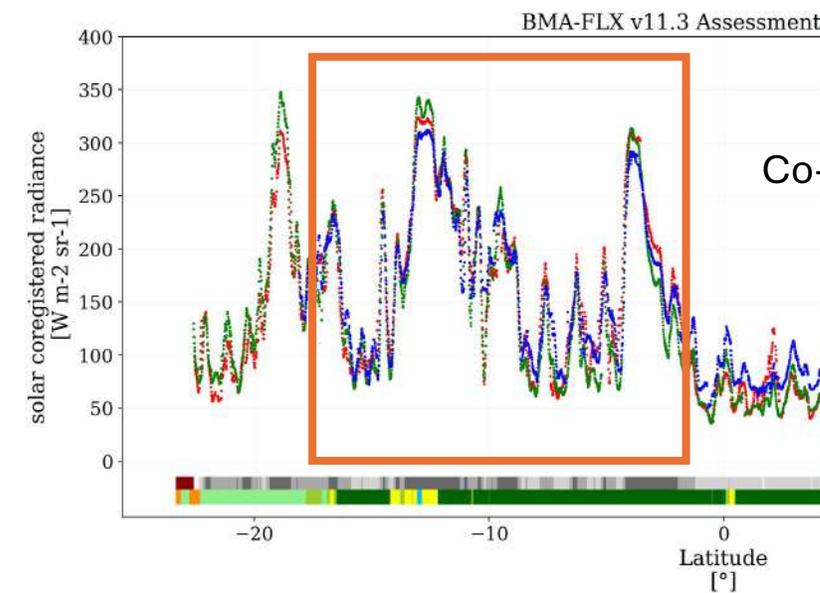
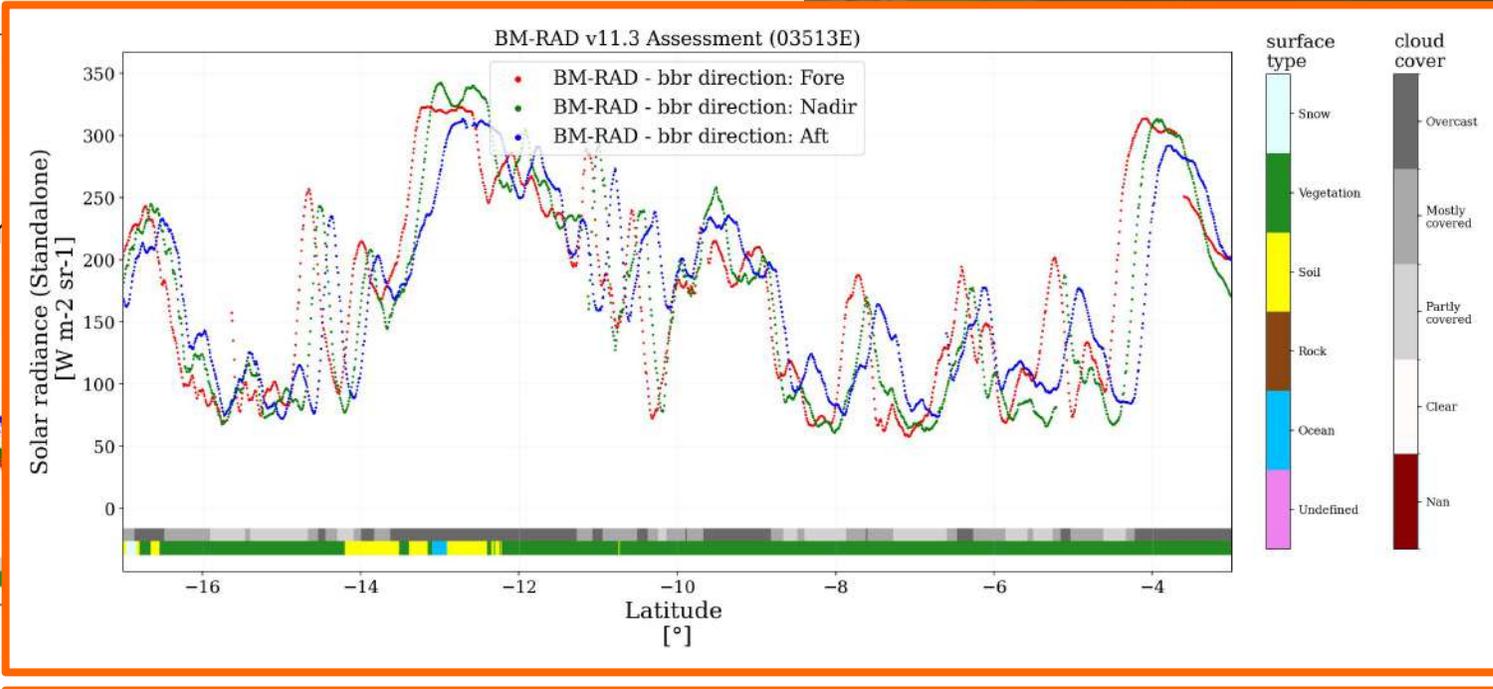
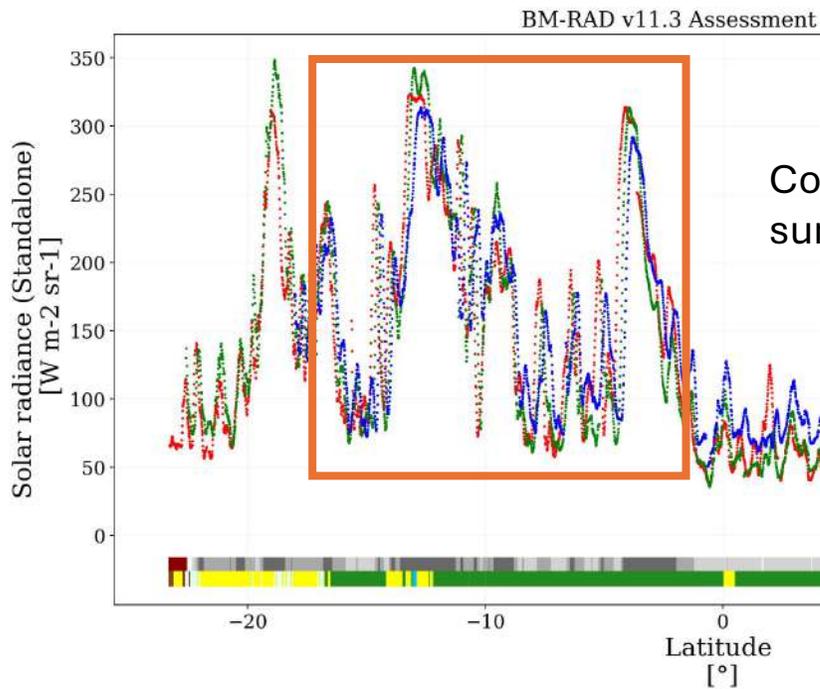
## Co-registration of LW radiances

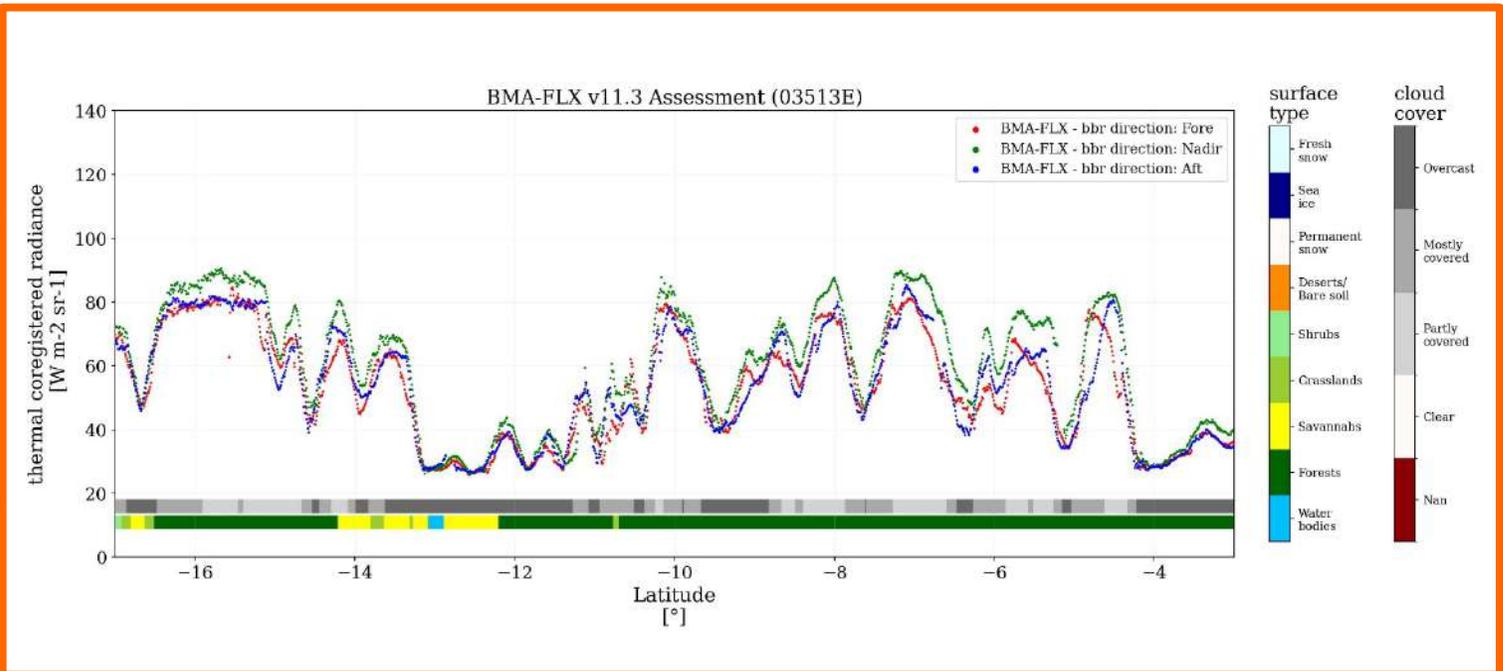
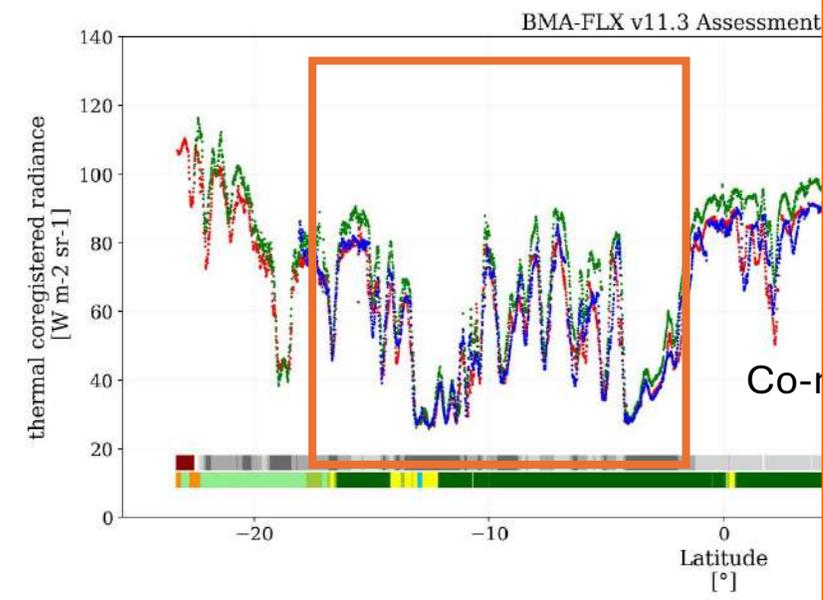
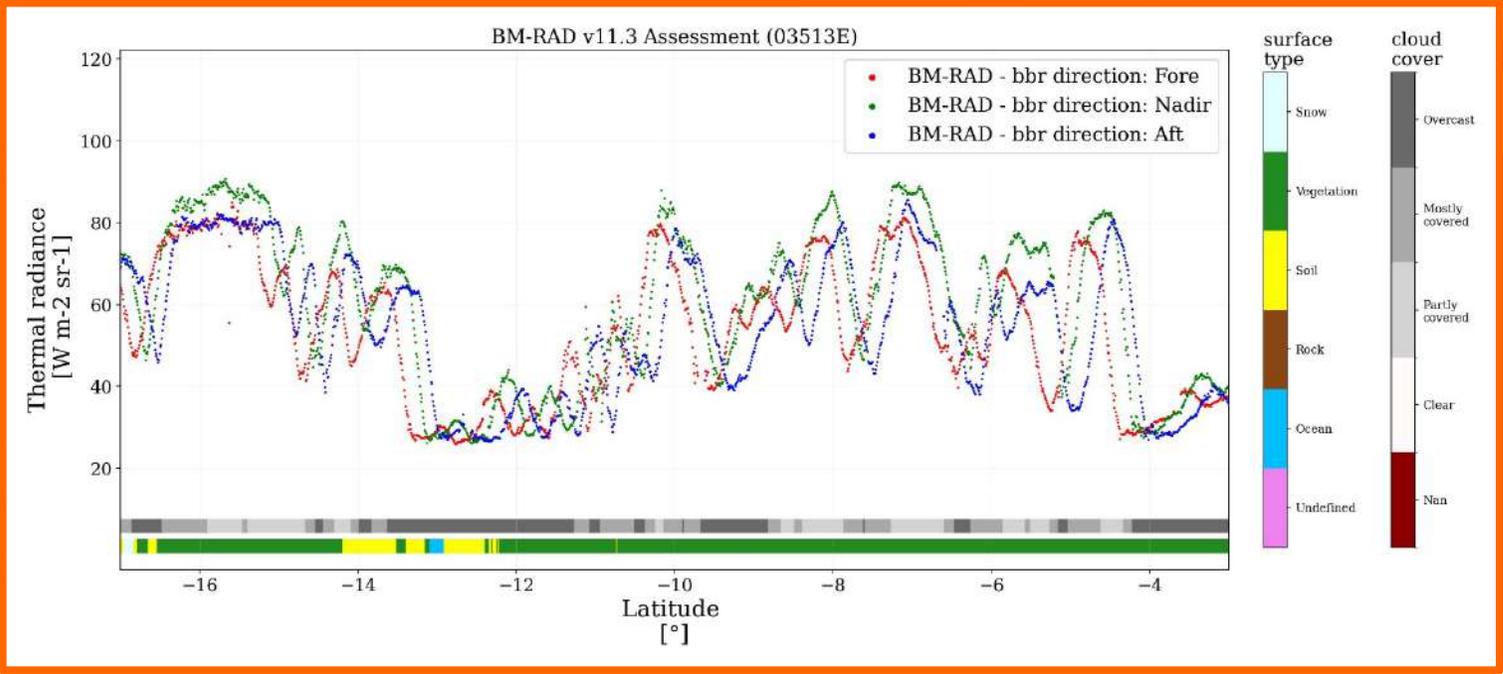
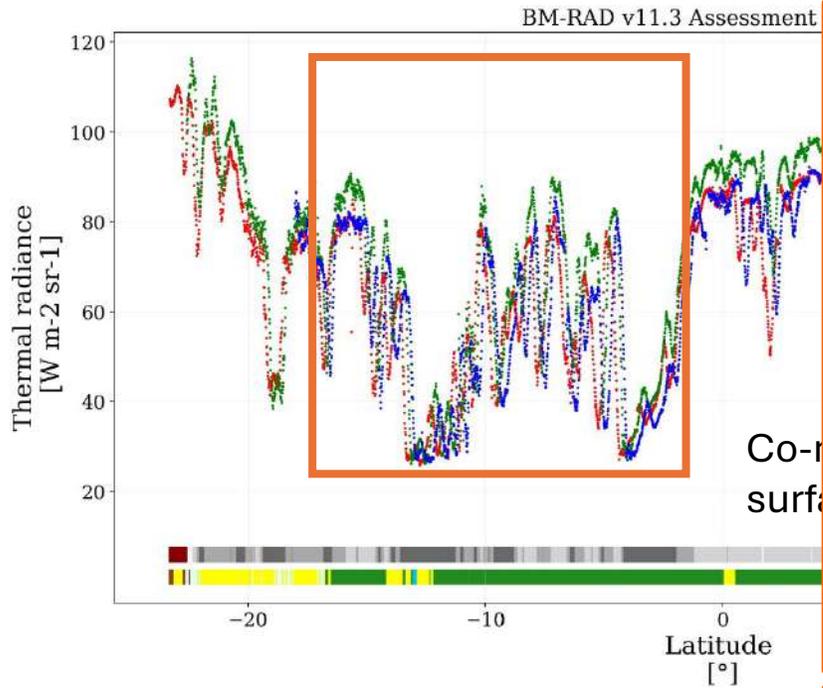
- The height defined by the CTH derived from brightness temperature ( M-COP) is a good estimator to co-register the BBR radiances in LW.
- RL is defined as the 90th percentile of the altitude of the highest cloud in the BBR domain.

## Co-registration of SW radiances

- SW radiances cannot be co-registered using BT-based CTH. In average co-registration errors are higher than using the default surface co-registration.
- Views are co-registered at a RL defined as the vertical level in the nadir domain that minimizes the flux differences between the nadir, aft and fore flux retrievals



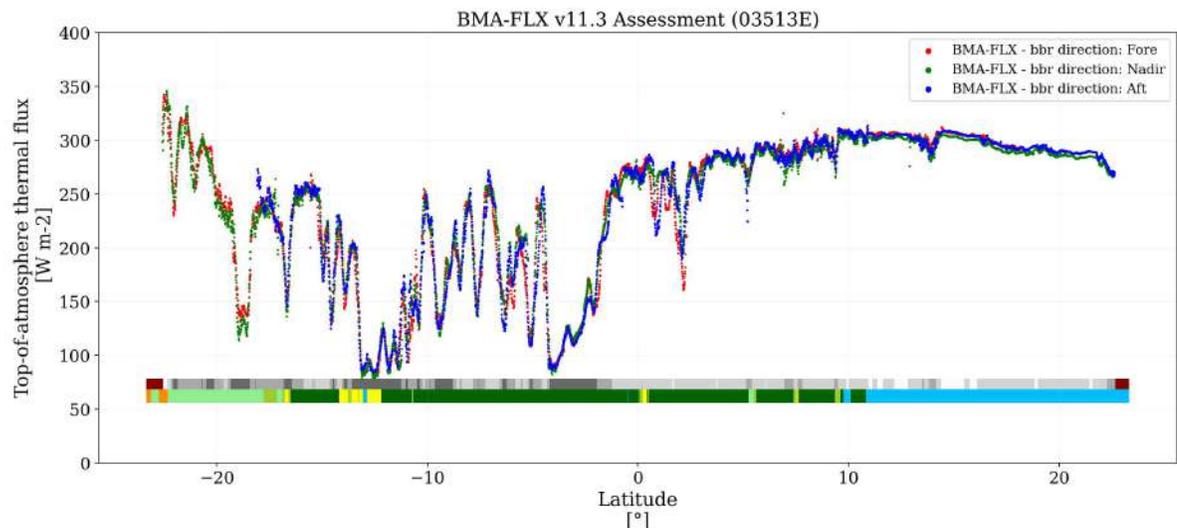
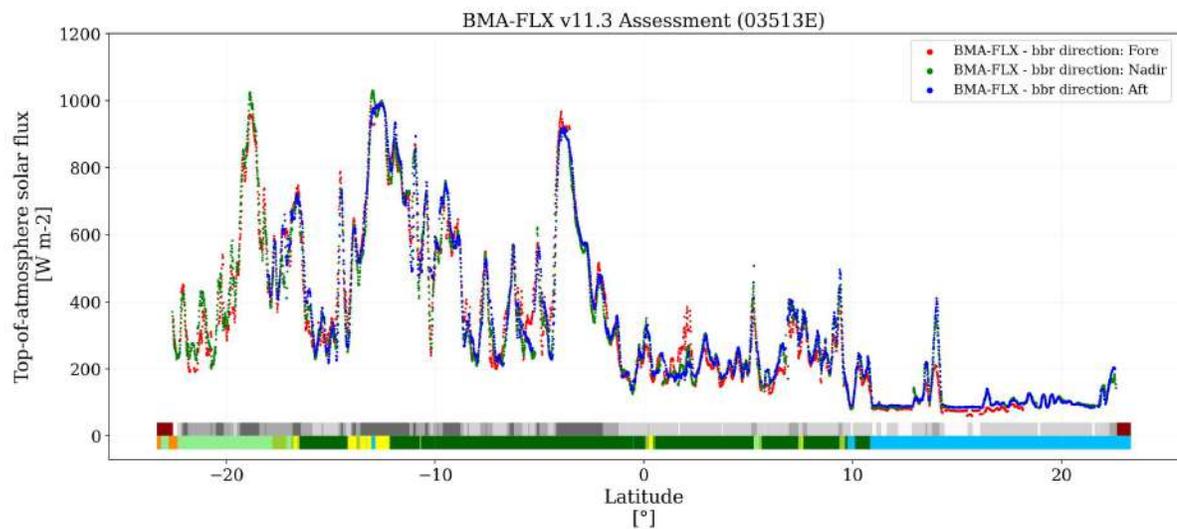




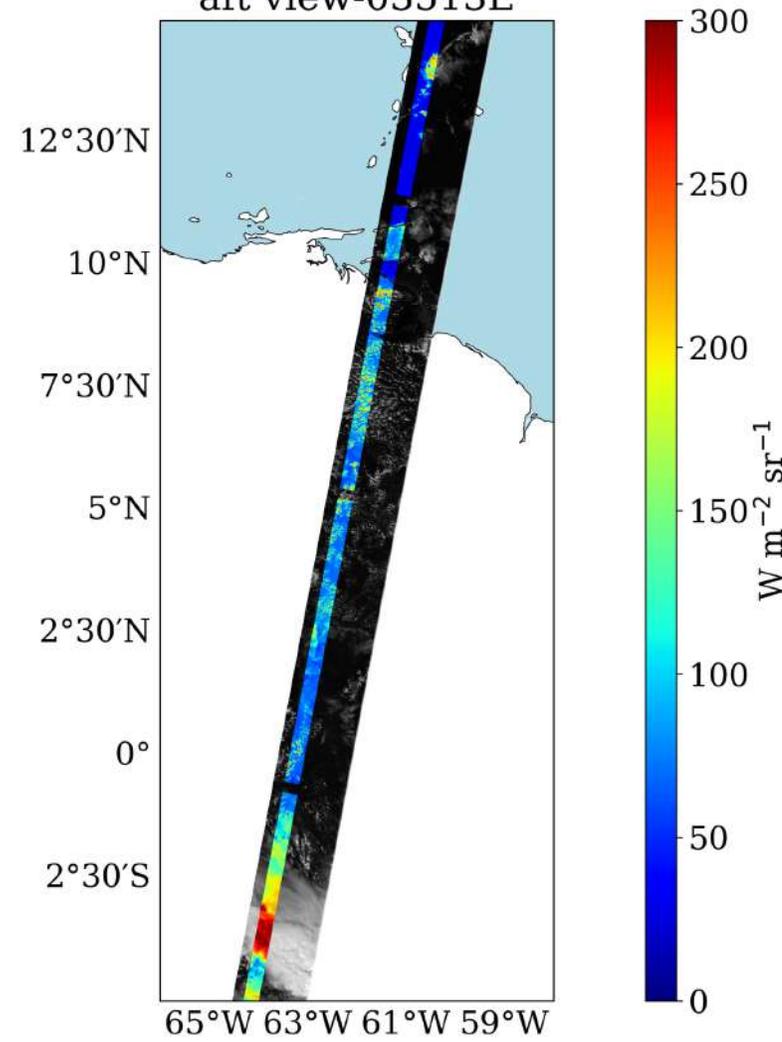


# SOLAR

# THERMAL



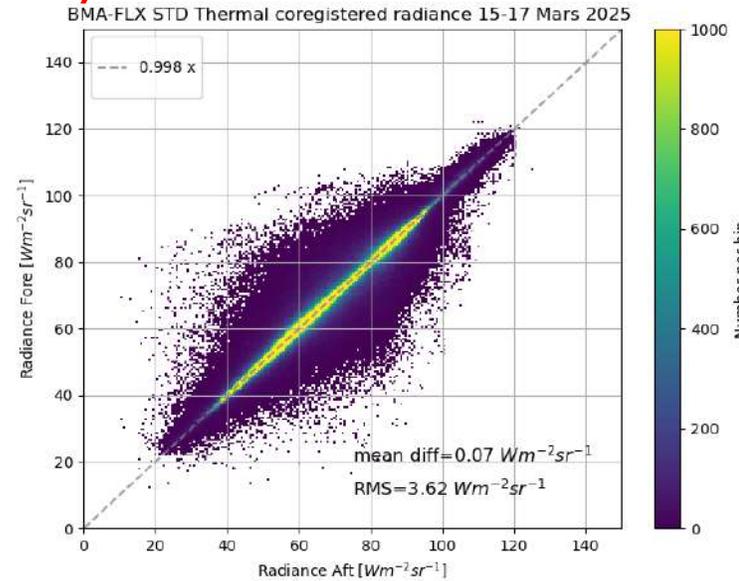
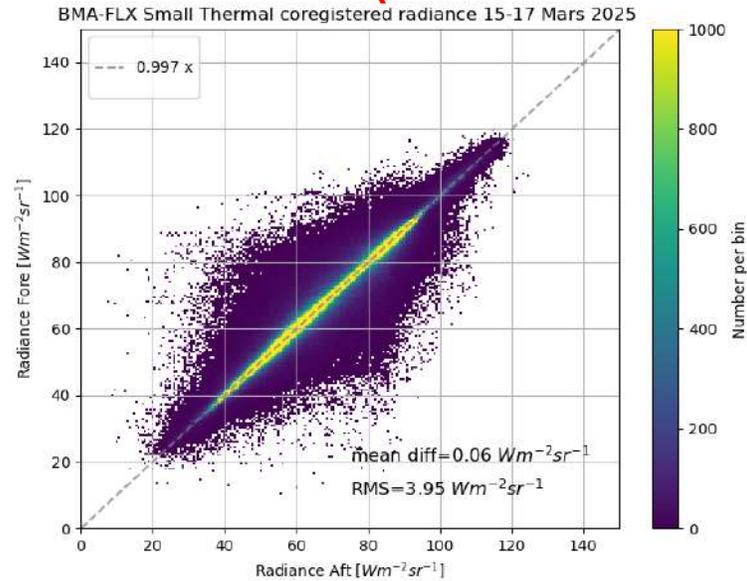
### B-SNG SW radiances aft view-03513E



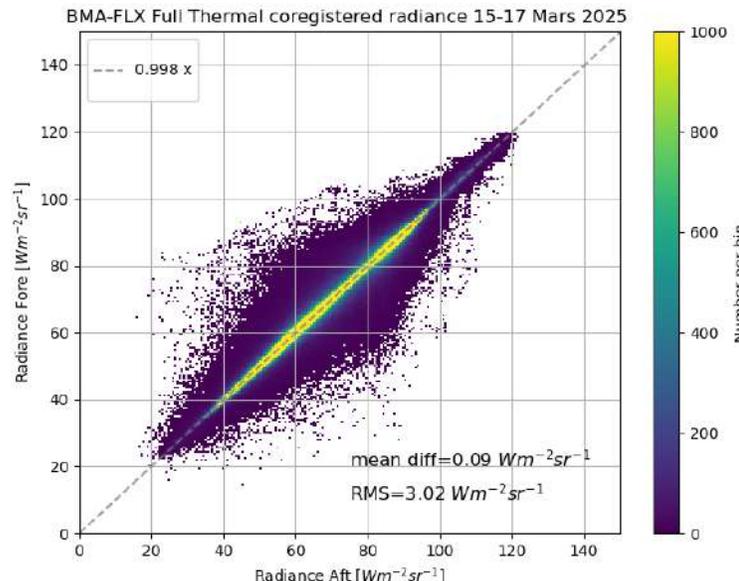
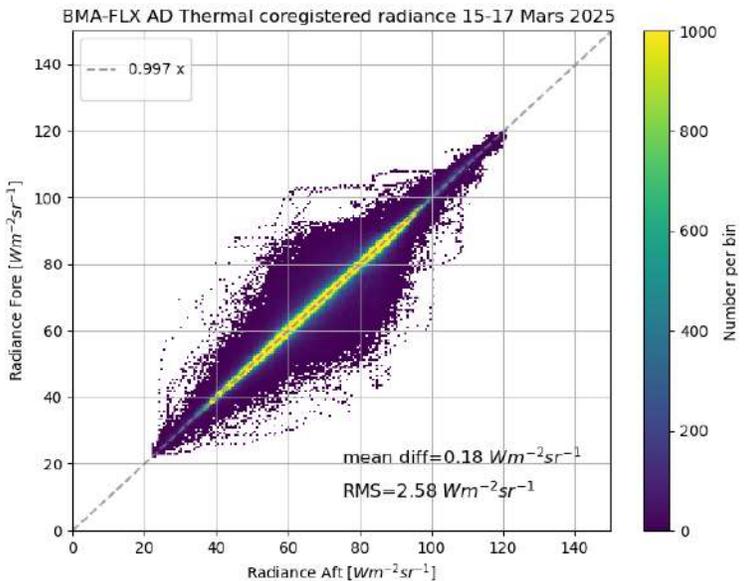
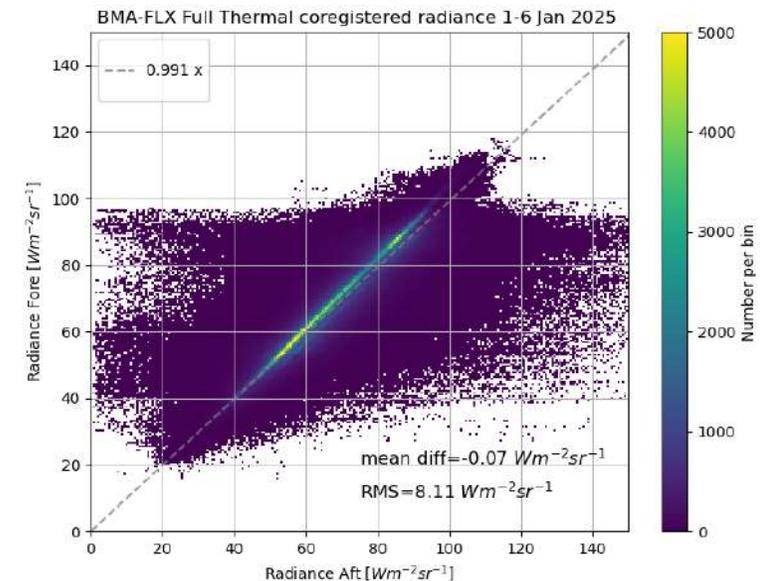
# Off-nadir thermal radiances (BMA-FLX AA)



From B-NOM AD (as from 14th January 2025)



From B-NOM AC

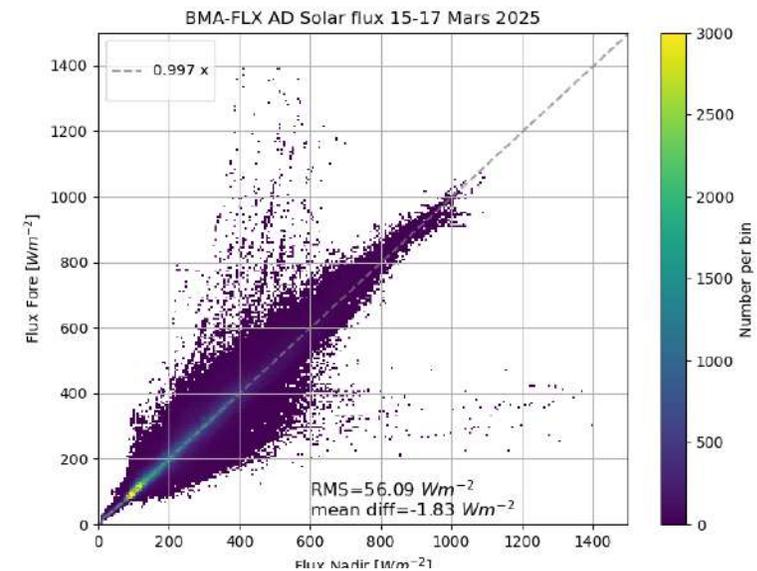
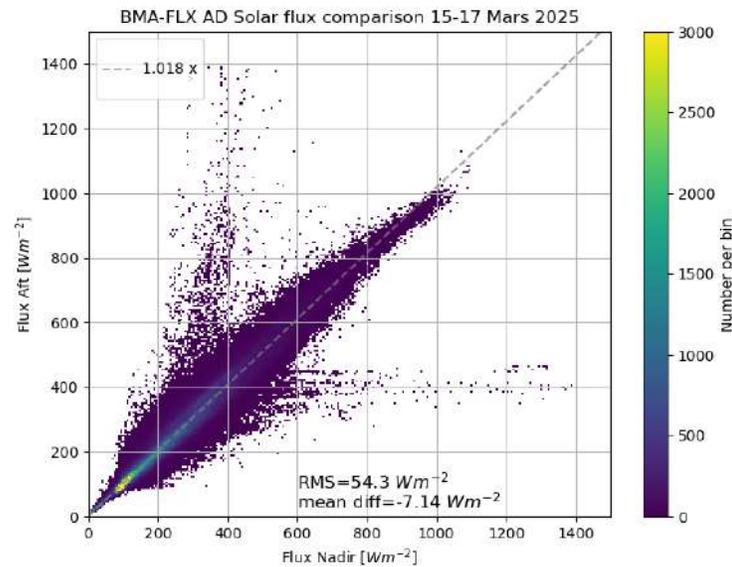
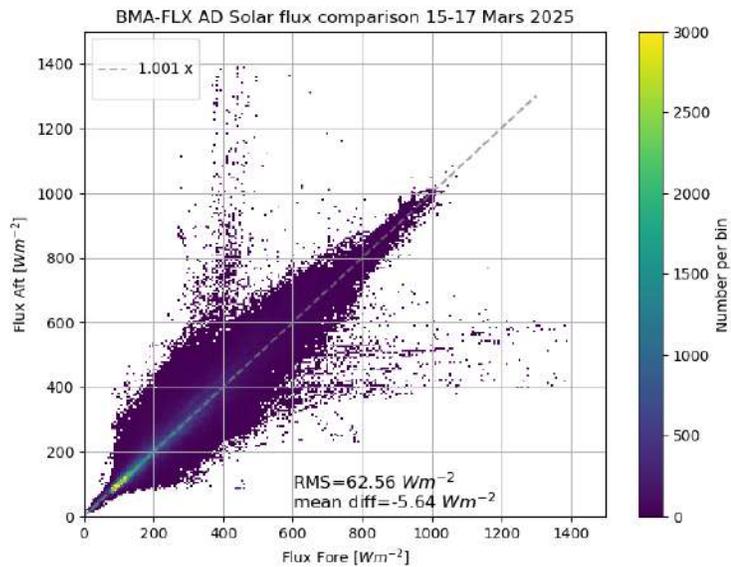


Unless reprocessed data, before 14th of January, risk that the dead pixel is used in the fore view for the STD and FULL resolutions in BM-RAD and BMA-FLX

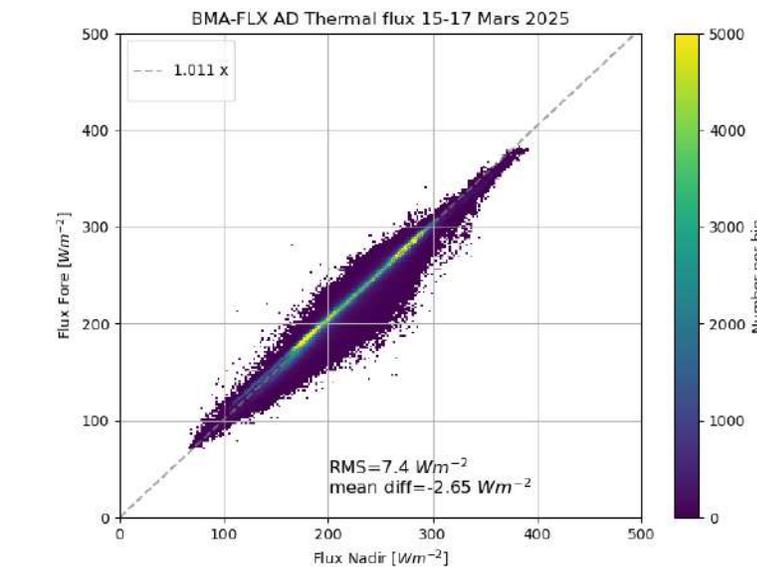
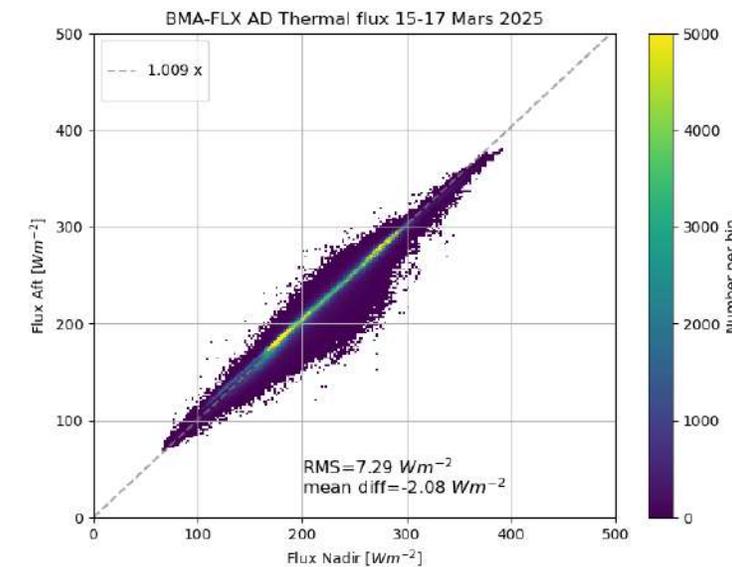
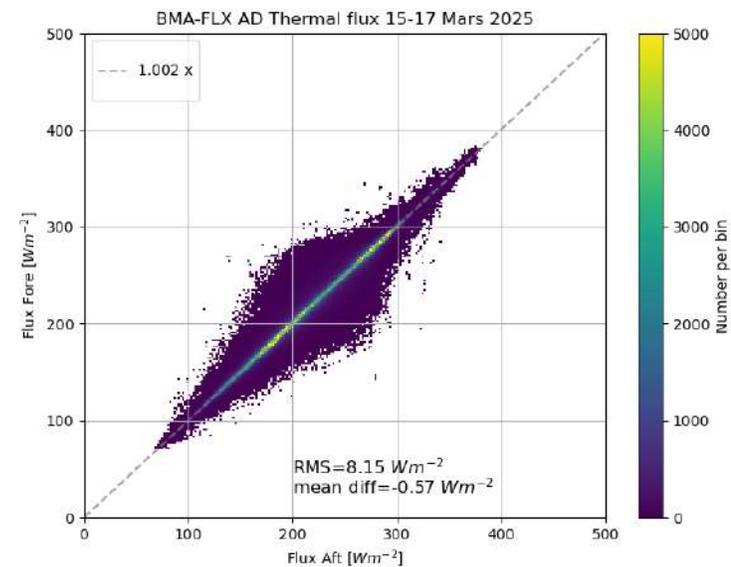
# BBR FLUXES COMPARISON: Assessment Domain



SOLAR

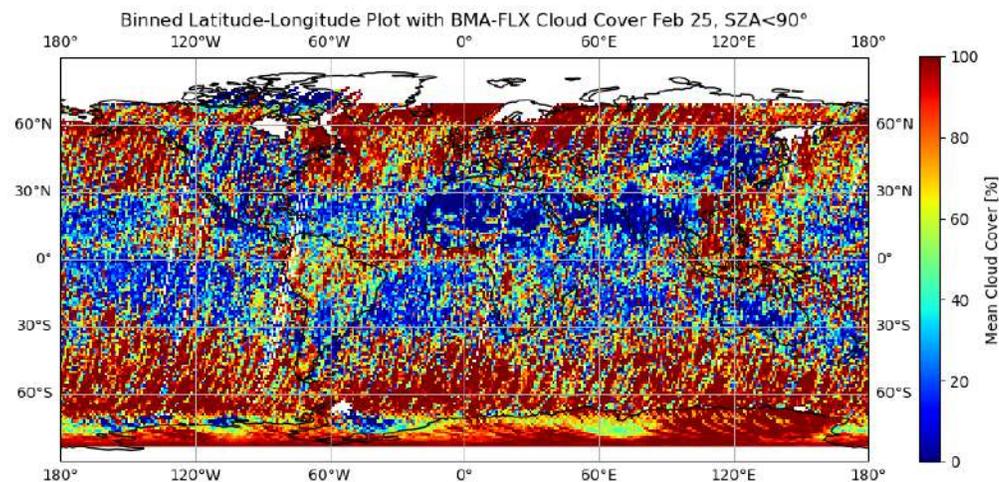
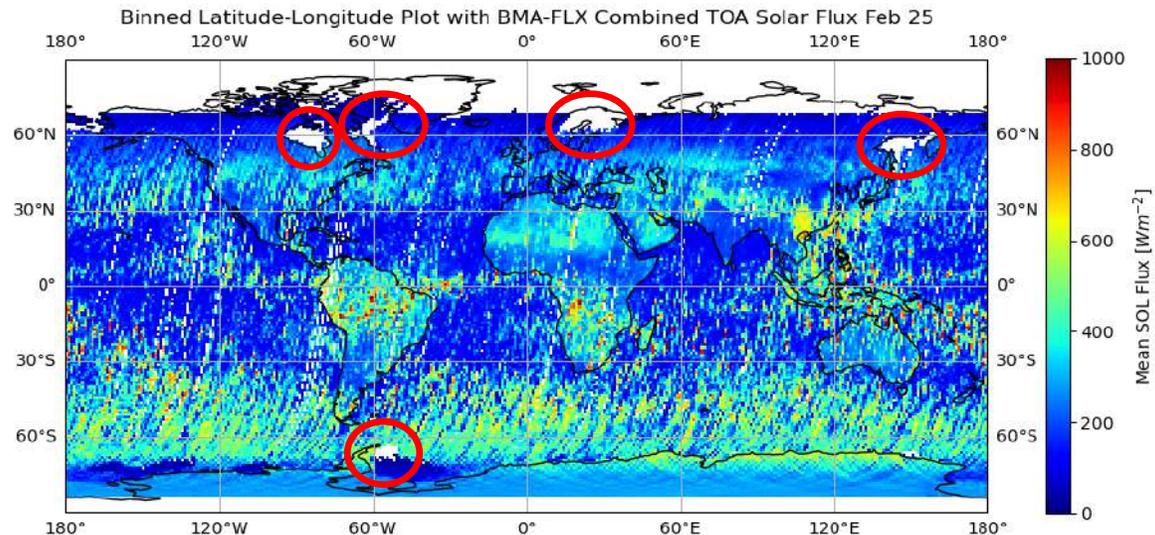


THERMAL

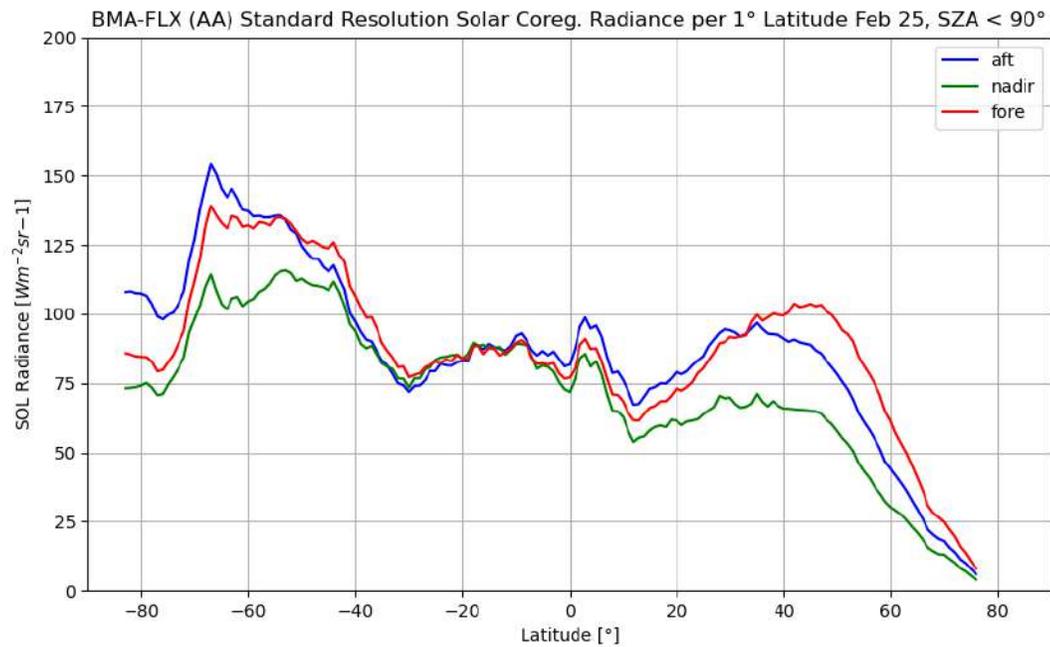




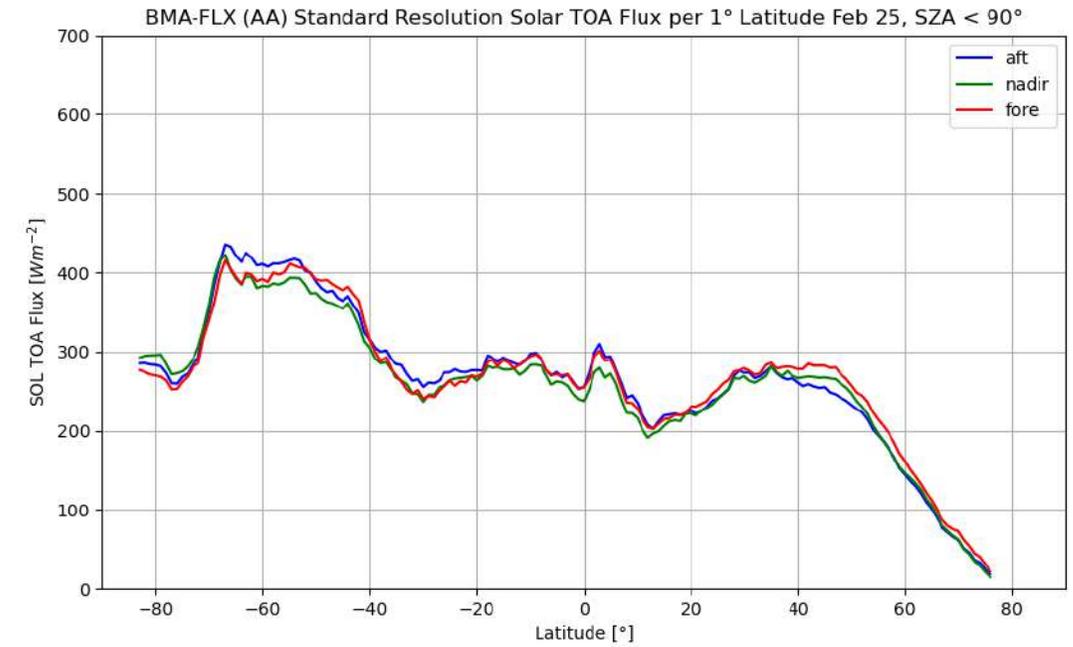
- Solar fluxes are very sensitive to the quality of the inputs used in the BMA-FLX processor.
- MSI L1 calibration → M-CM L2a & M-COP L2a
- Missing values in solar fluxes are due to:
  - Fill values in cloud cover (in red)
  - Unphysical solar fluxes over snow/ice surfaces
- Solar fluxes are expected to improve when MSI L1 calibration does.



## Radiances



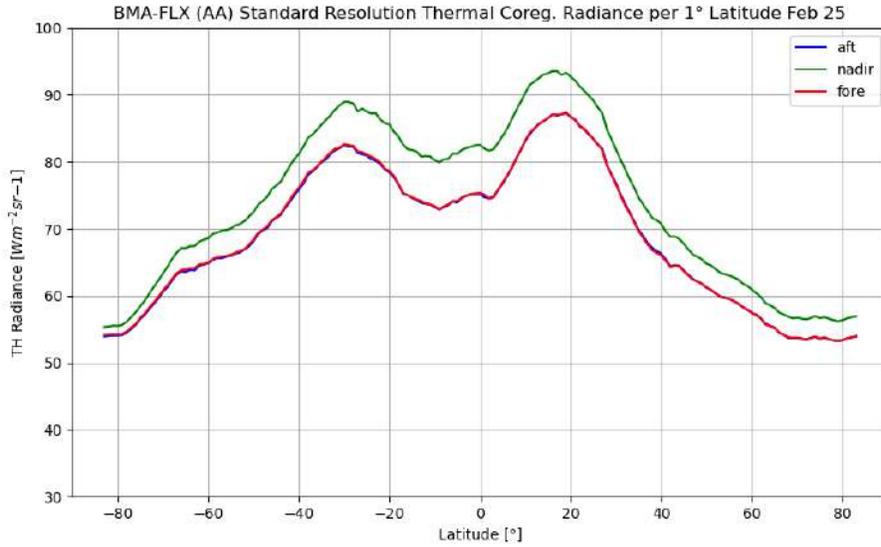
## Fluxes



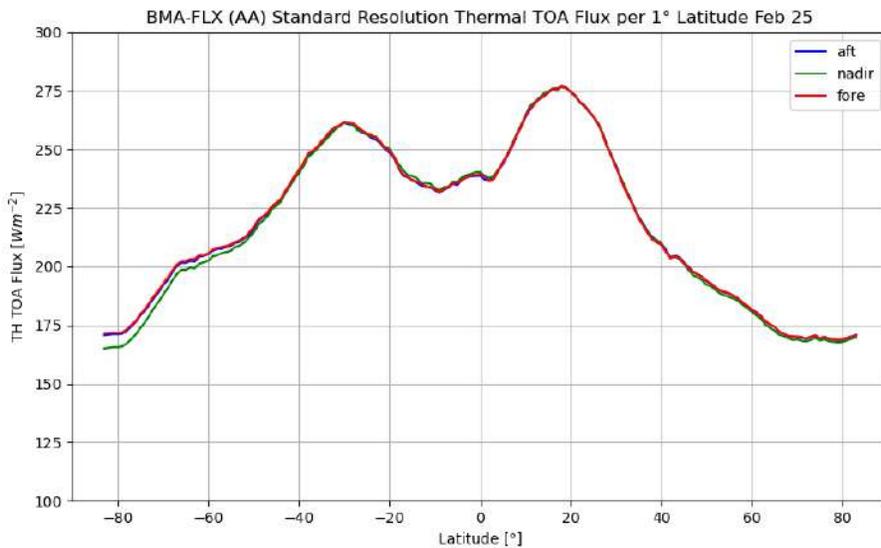
# Latitudinal means: Thermal radiances & fluxes



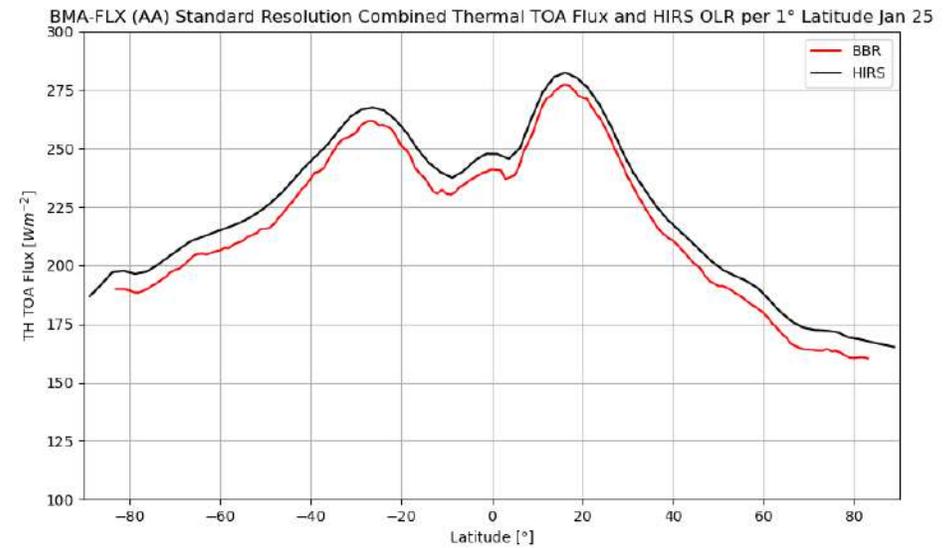
Radiances



Fluxes



## HIRS NOAA OLR daily mean BBR 14:00 Descending node



# GERB-like/ BBR comparisons: SW and LW Radiances

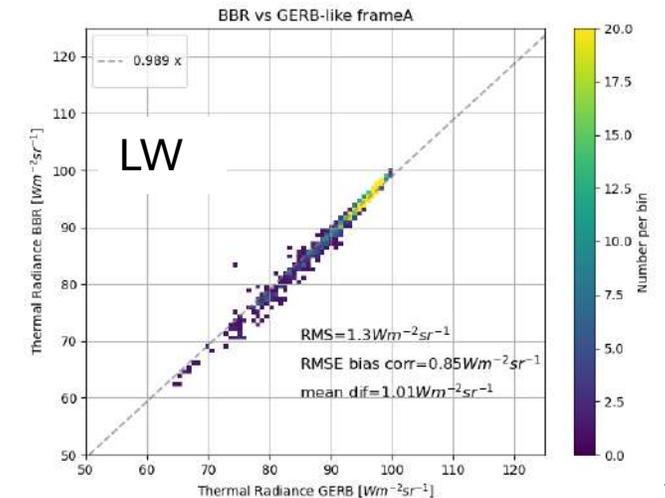
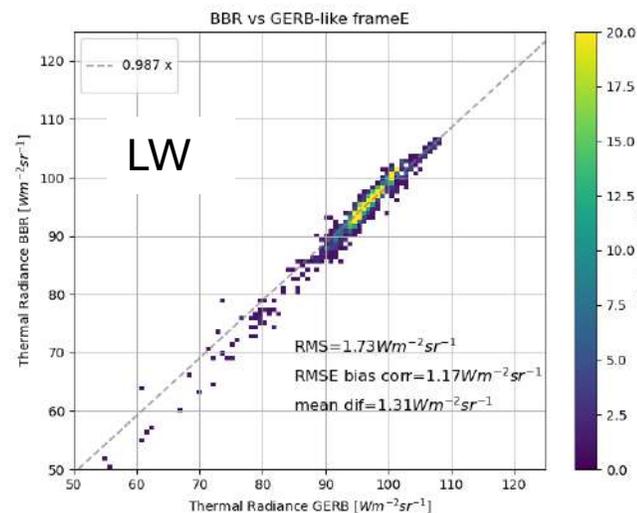
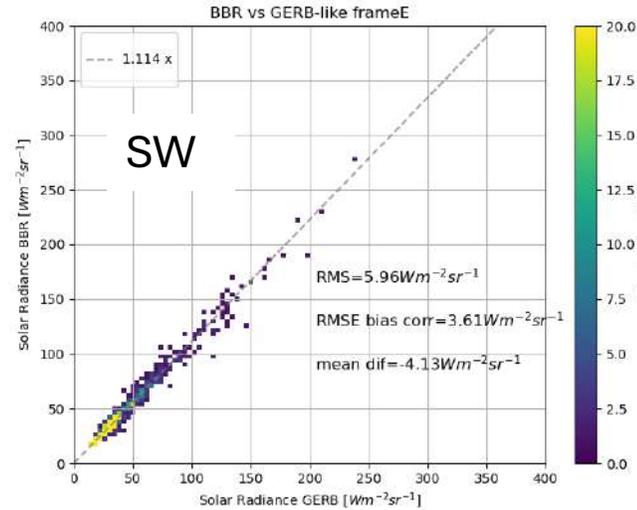


Daytime

Nighttime

- GERB-1 is over the Indian ocean
- Data from October 2024 to February 2025
- Only nadir comparisons

- BBR solar radiances are higher than GERB-like
- BBR thermal radiances are lower



GERB VZA < 3 deg  
 $\Delta$ dist < 3 km  
 $\Delta$ time < 5 min

# CERES FLASHflux comparison: SW and LW Radiance

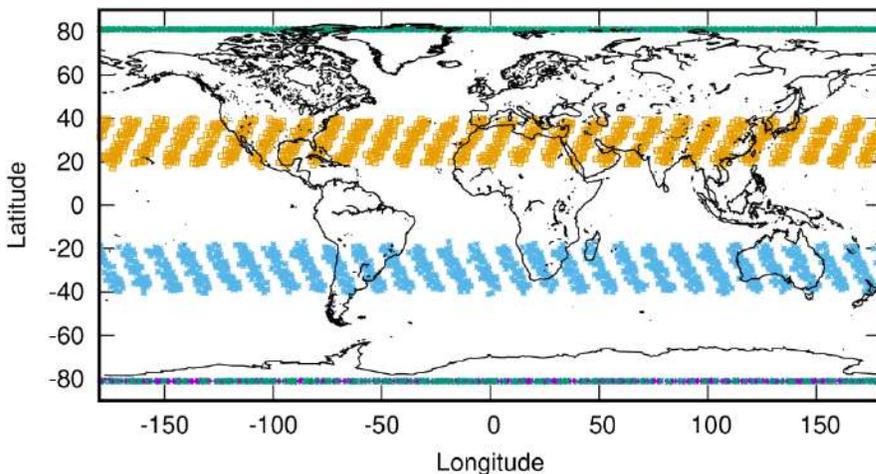


Daytime

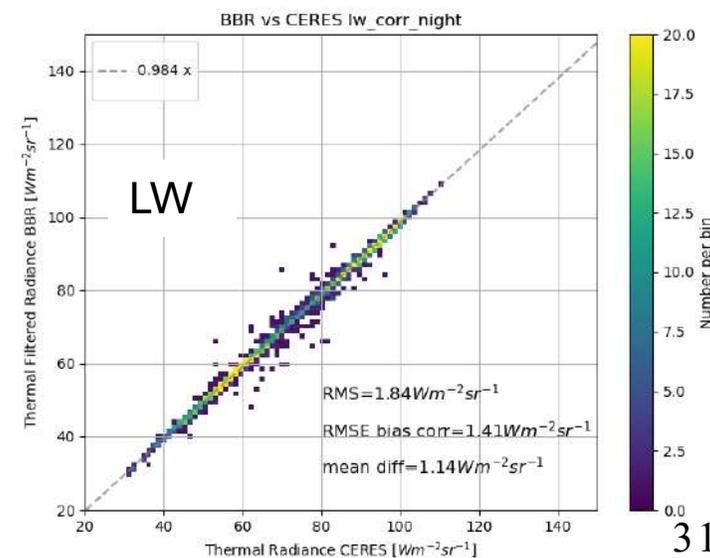
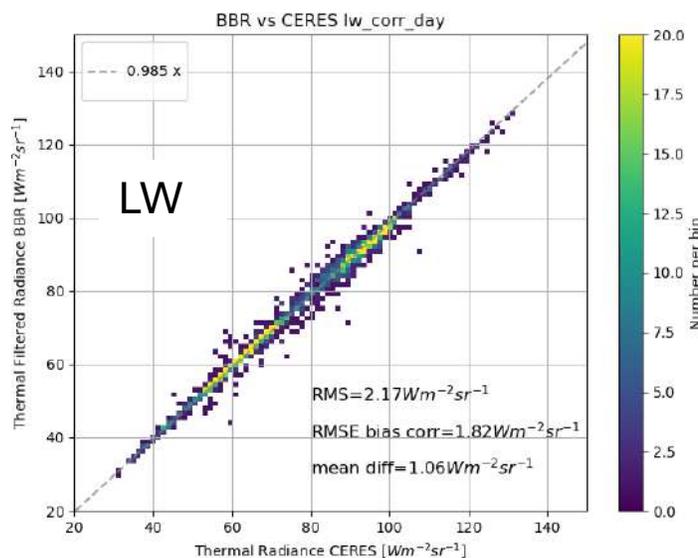
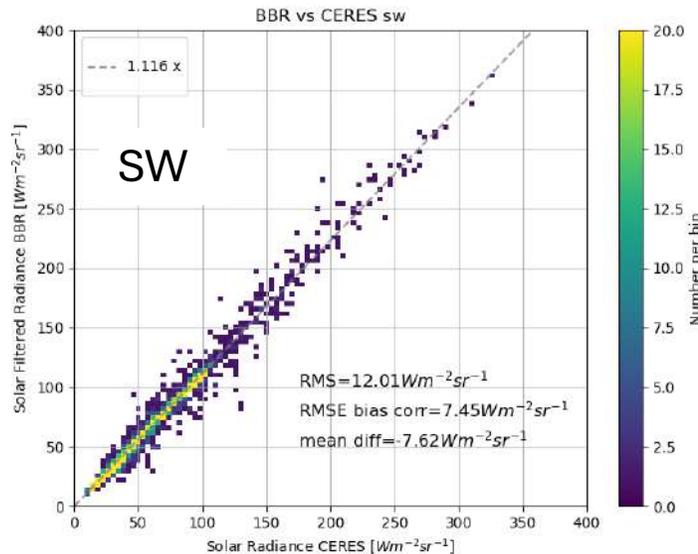
Nighttime

10 August 2024 to  
14 February 2025

Terra - day +  
Terra - night x  
NOAA20 - day \*  
NOAA20 - night □



**CERES VZA < 3°**  
**Δdist < 3 km**  
**Δtime < 5 min**

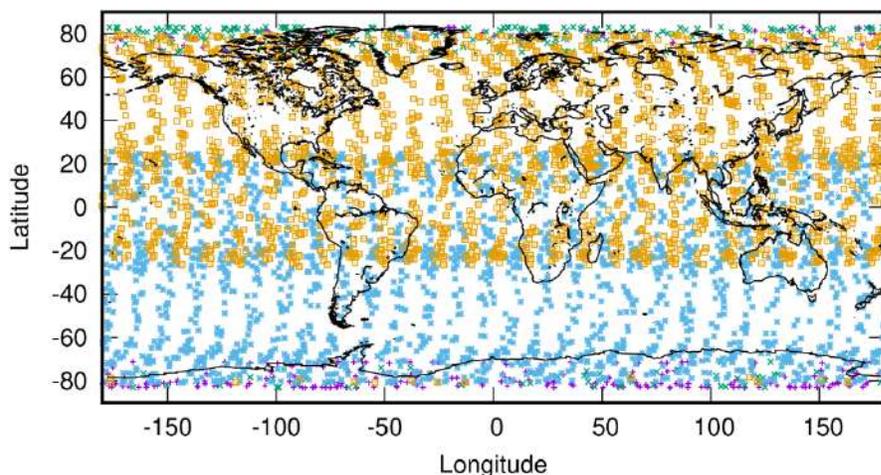


# CERES FLASHflux comparison: SW Flux, Daytime

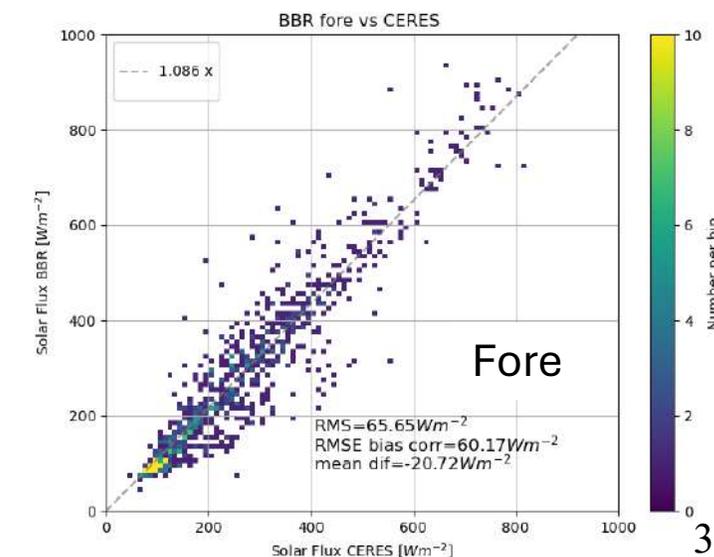
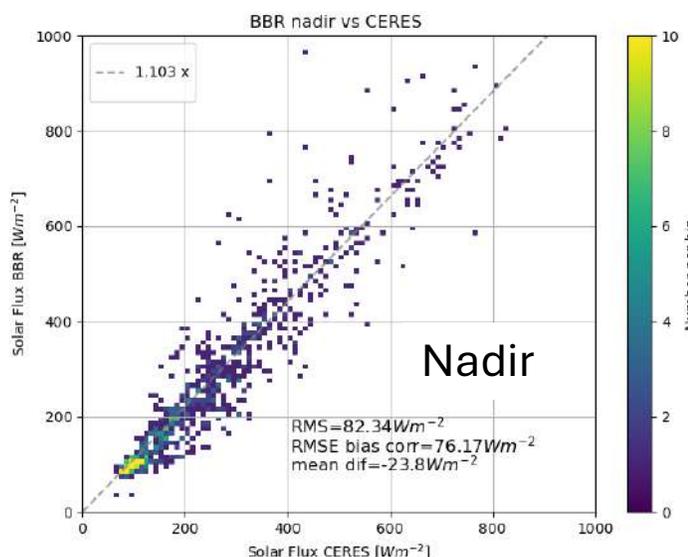
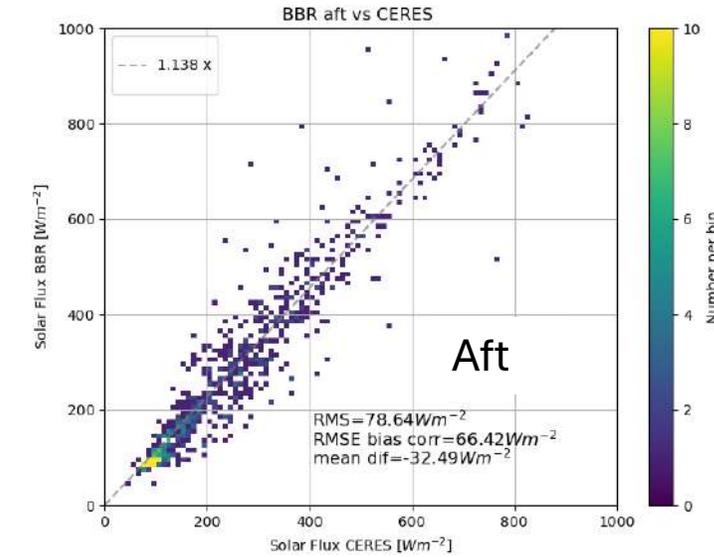
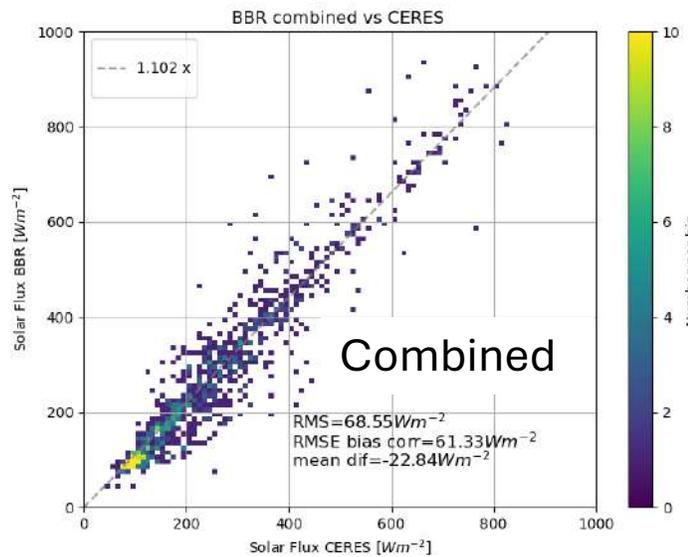


10 August 2024 to  
14 February 2025

Terra - day +  
Terra - night x  
NOAA20 - day \*  
NOAA20 - night □



**CERES VZA < 50°**  
**Δdist < 3 km**  
**Δtime < 5 min**

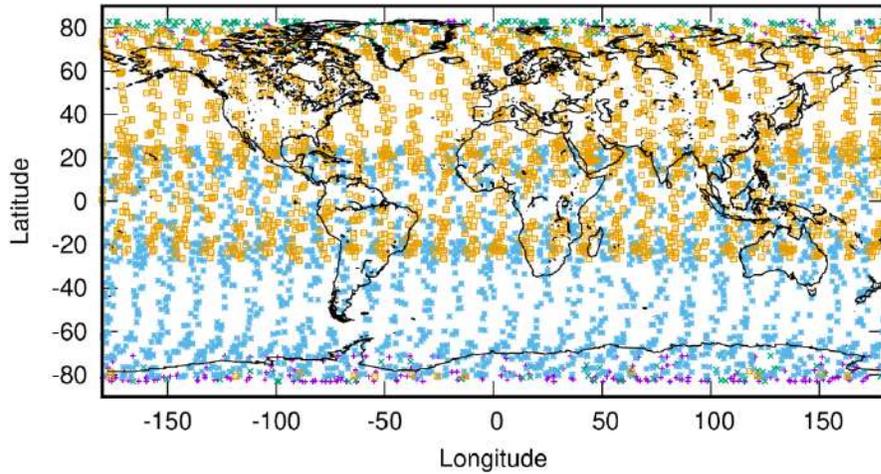


# CERES FLASHflux comparison: LW Flux

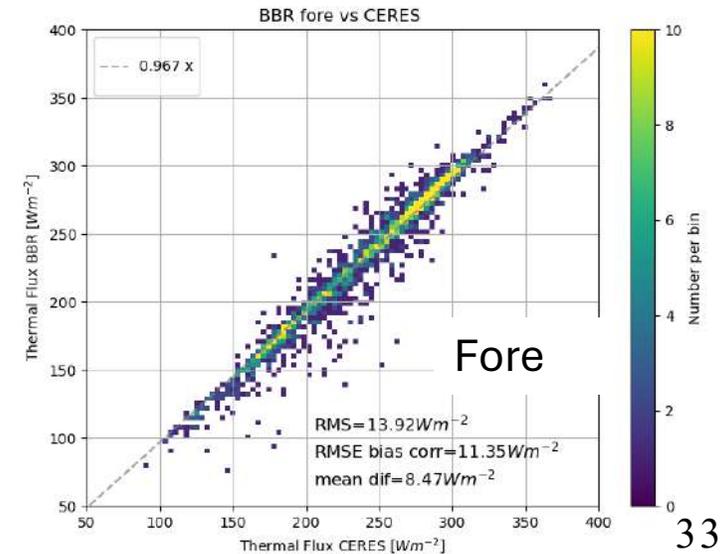
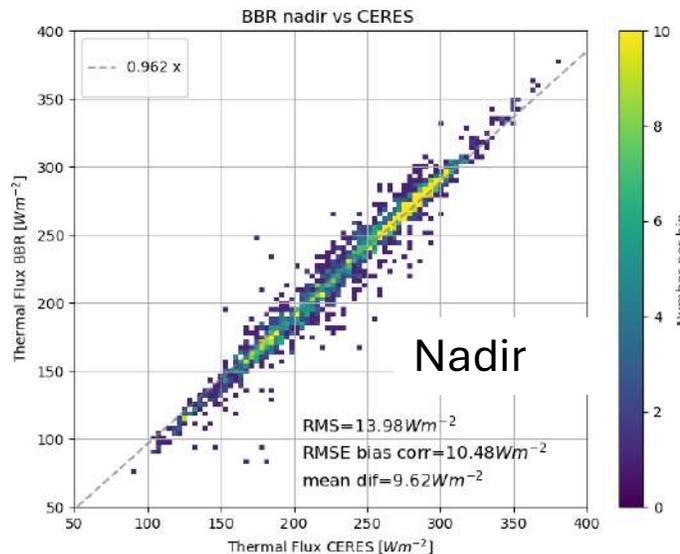
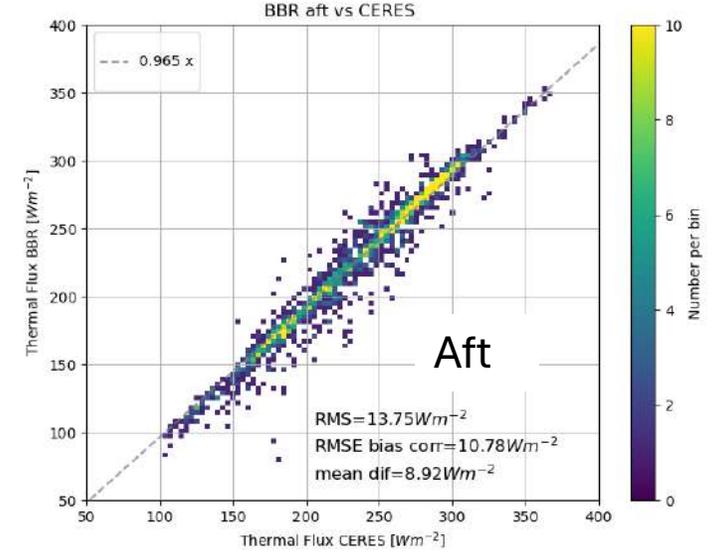
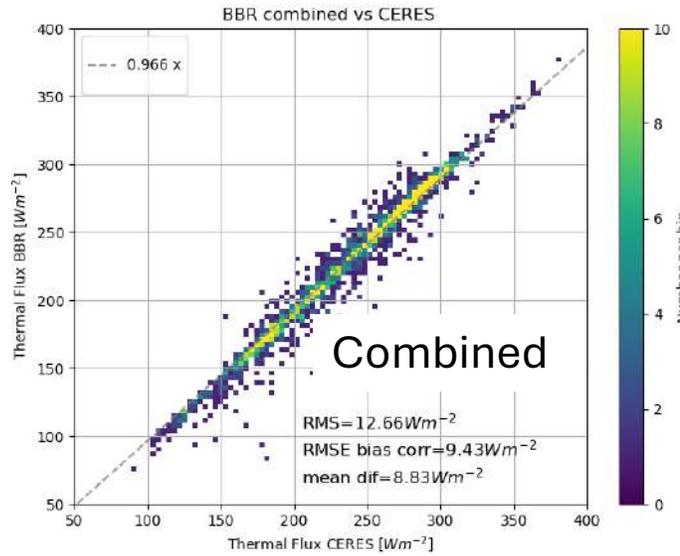


10 August 2024 to 14  
February 2025

Terra - day +  
Terra - night x  
NOAA20 - day \*  
NOAA20 - night □



**CERES VZA < 50°**  
**Δdist < 3 km**  
**Δtime < 5 min**



# Summary



- Very good agreement for the BBR LW off-nadir radiances in all the resolutions (STD, Small and AD and Full Resolution)
  - New baseline available (AD) for BBR L1 since 13/01/2025 frame 3571B masking the dead pixel
- LW fluxes show very good agreement between the 3 views
- SW fluxes comparison is more complicated → Sensitive to co-registration of the off-nadir views, cloud mask/properties, CTH, MSI radiances, ADMs, etc.
- Preliminary comparisons with CERES FLASHflux suggest that BBR SW fluxes are brighter than CERES and BBR LW fluxes are lower → to be verified when CERES SSF data is available.
  - RMSE in the SW likely affected by collocation and not representative of the instantaneous combined flux error
- Good agreement with GERB-1 radiances over Indian Ocean. BBR solar radiances brighter than GERB, thermal lower.
- Future work: evaluate impact of proposed update on the BBR L1 calibration on the radiances and fluxes.

# Recommendations to users



- Choose resolution -> No combined flux in the Full resolution
- If radiance comparisons using the 3 BBR-views -> BMA-FLX co-registered radiances
- Check quality status flags
- Read disclaimers for BM-RAD (available in ECVT) and BMA-FLX (to be published soon)
- BM-RAD and BMA-FLX description in:

Velázquez Blázquez, A., Baudrez, E., Clerbaux, N., and Domenech, C.: Unfiltering of the EarthCARE Broadband Radiometer (BBR) observations: the BM-RAD product, *Atmos. Meas. Tech.*, 17, 4245–4256, <https://doi.org/10.5194/amt-17-4245-2024>, 2024.

Velázquez Blázquez, A., Domenech, C., Baudrez, E., Clerbaux, N., Salas Molar, C., and Madenach, N.: Retrieval of top-of-atmosphere fluxes from combined EarthCARE lidar, imager, and broadband radiometer observations: the BMA-FLX product, *Atmos. Meas. Tech.*, 17, 7007–7026, <https://doi.org/10.5194/amt-17-7007-2024>, 2024.

- Questions? <https://ecvt.csde.esa.int/forum/>  
[almudena.velazquez@meteo.be](mailto:almudena.velazquez@meteo.be)

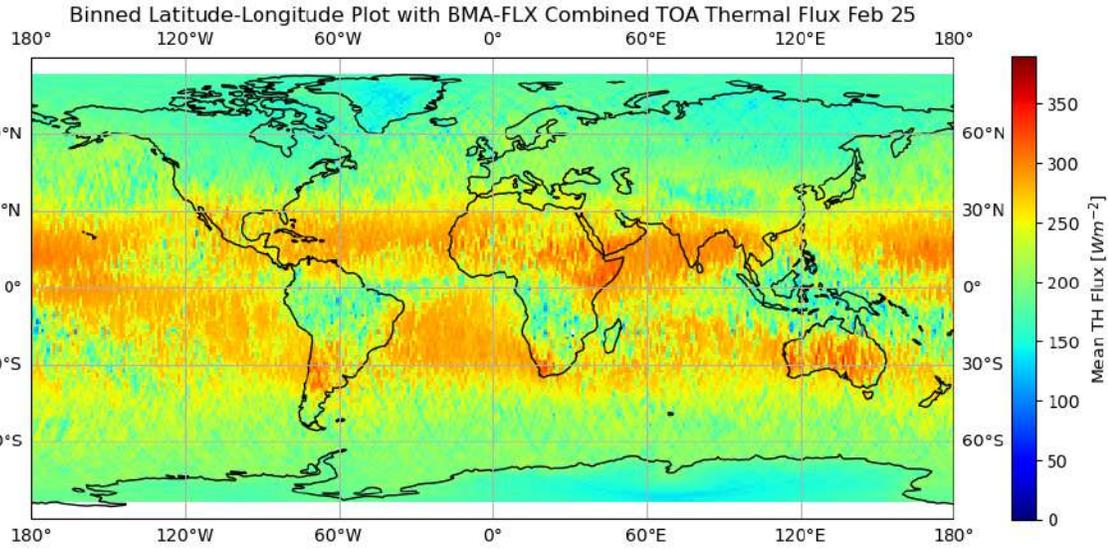


**Thank you!**

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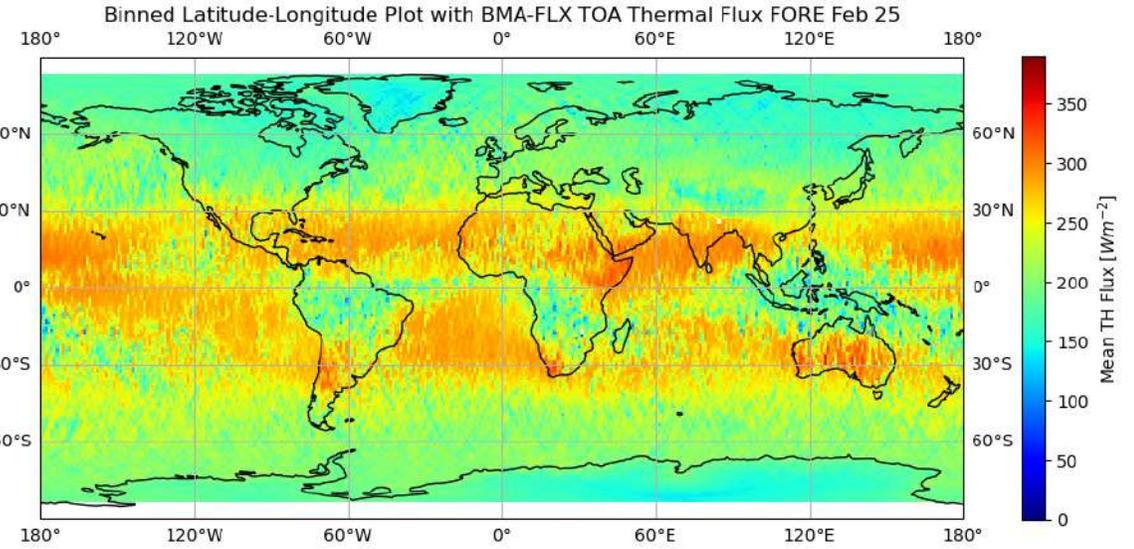
**2<sup>nd</sup> ESA-JAXA EarthCARE In-Orbit Validation Workshop**  
17 – 20 March 2025 | ESA-ESRIN | Frascati (Rome), Italy

COMBINED

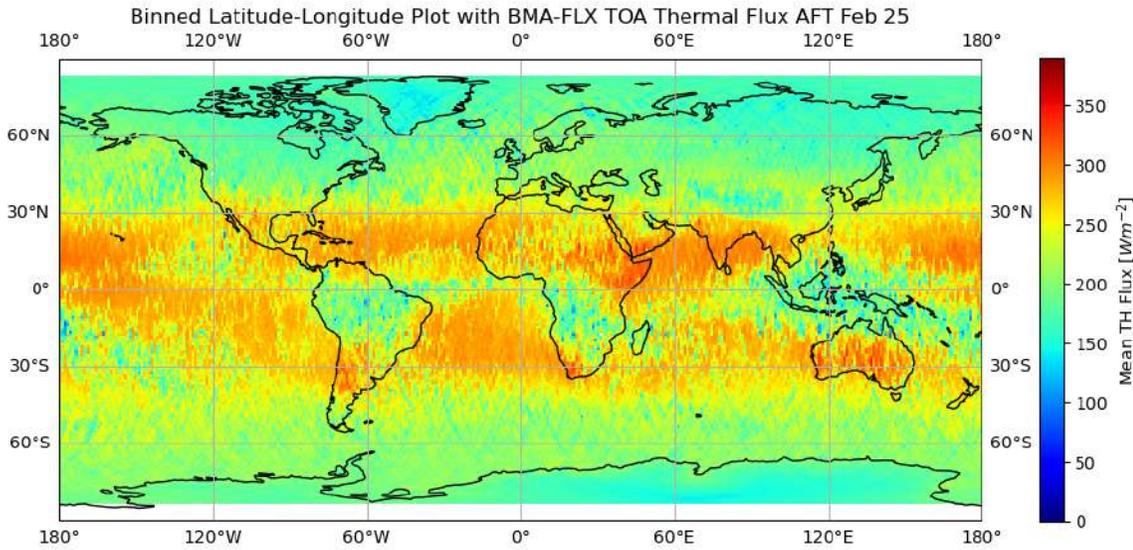


# THERMAL

FORE



AFT



NADIR

