

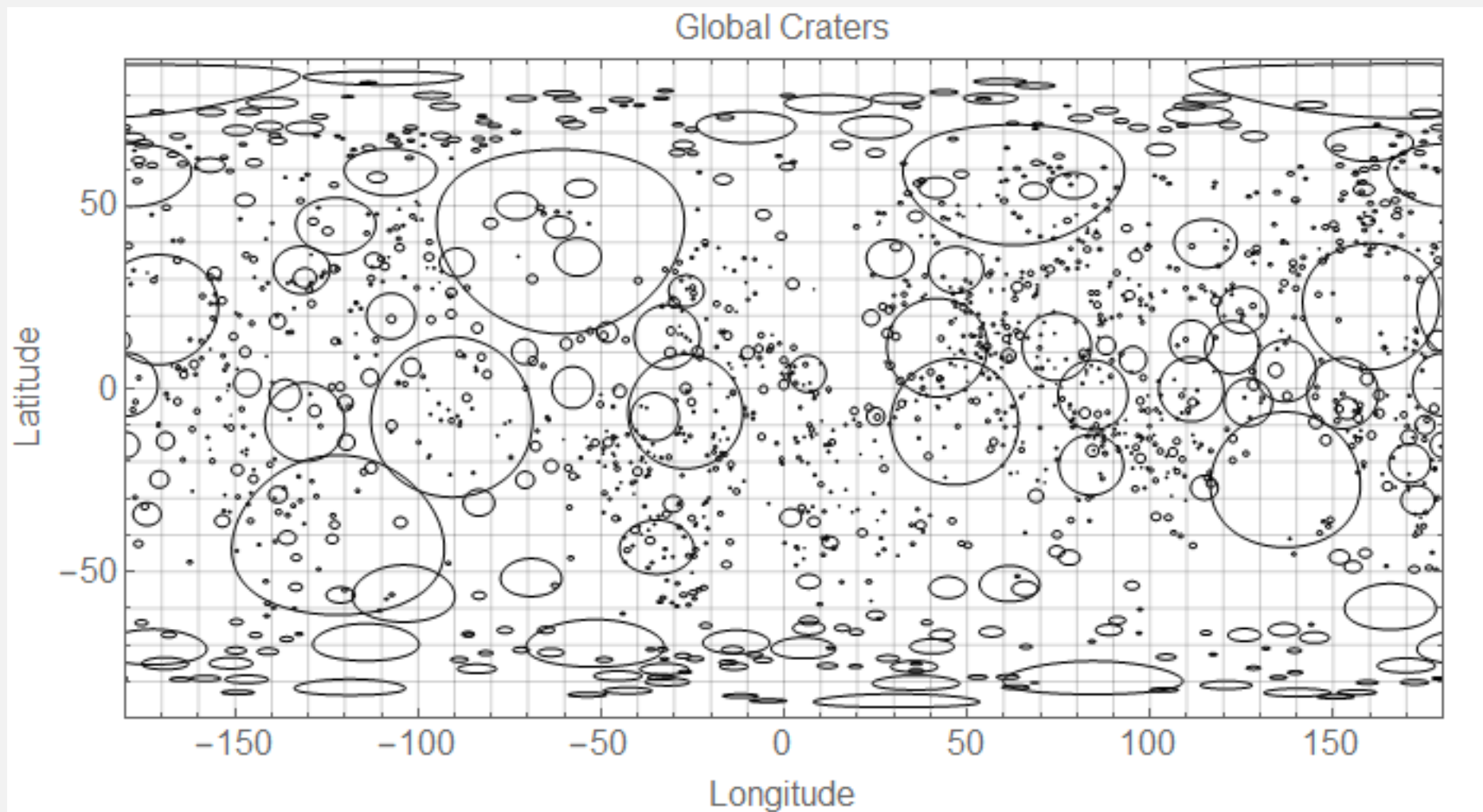


OSIRIS-REx
Activity Presentation

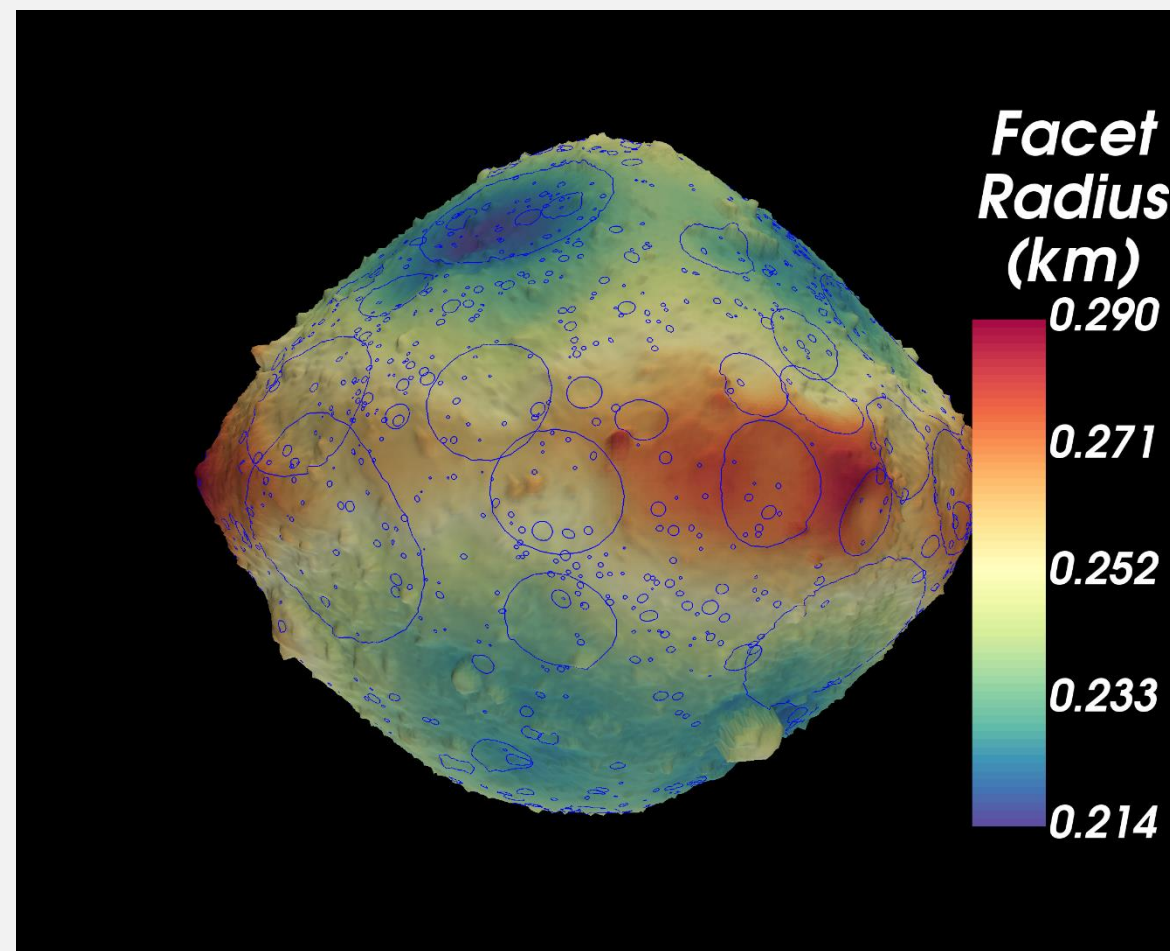
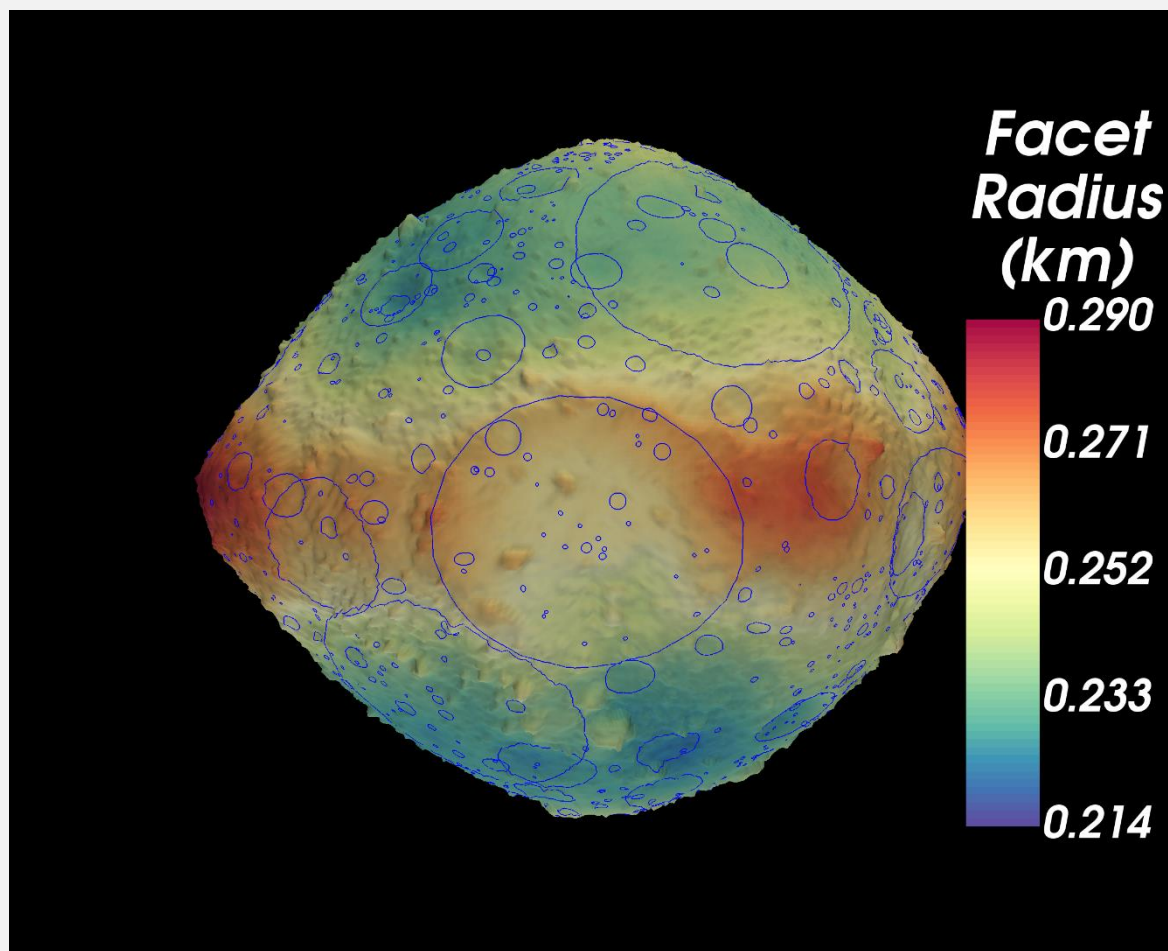
Bennu craters in the context of planetary defense

Beau Bierhaus
Lockheed Martin Space
26 April 2021

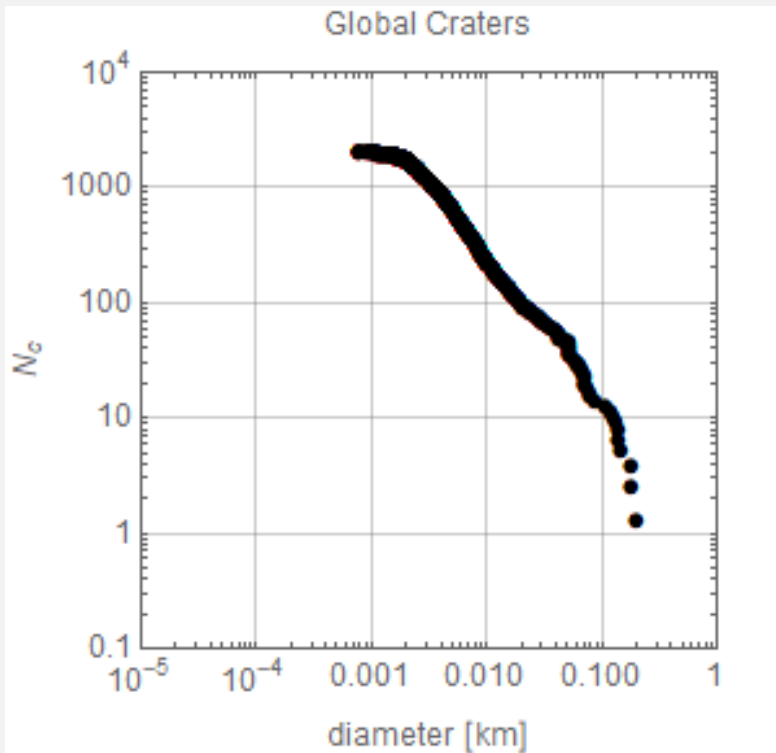
Bennu's craters



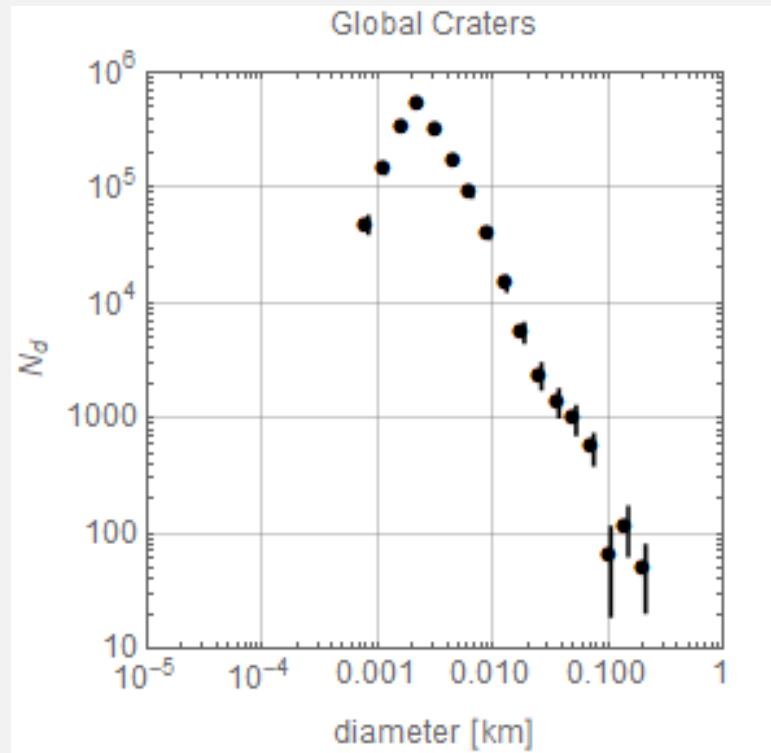
Projected on the shape model



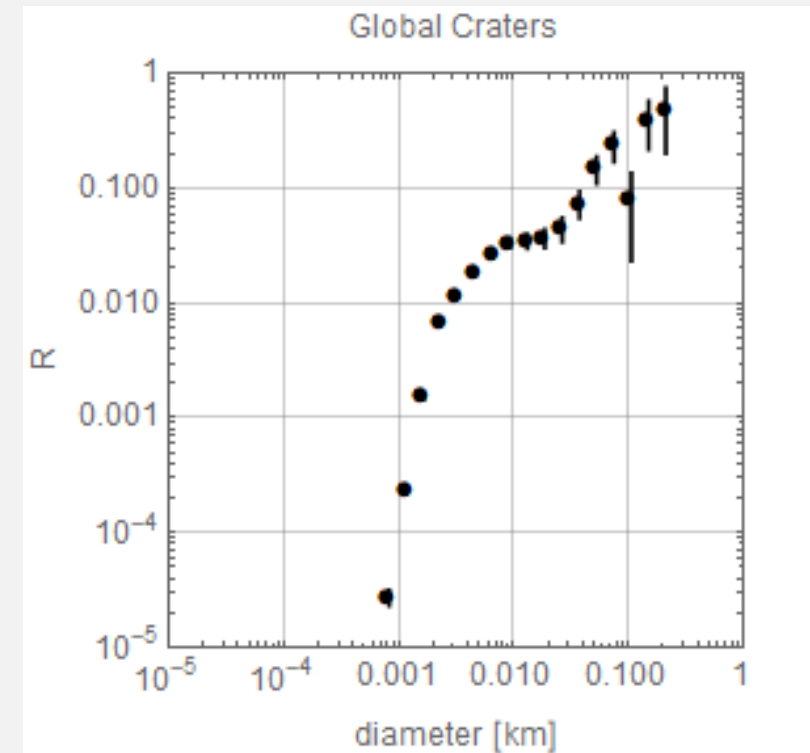
Crater SFD



cumulative



differential

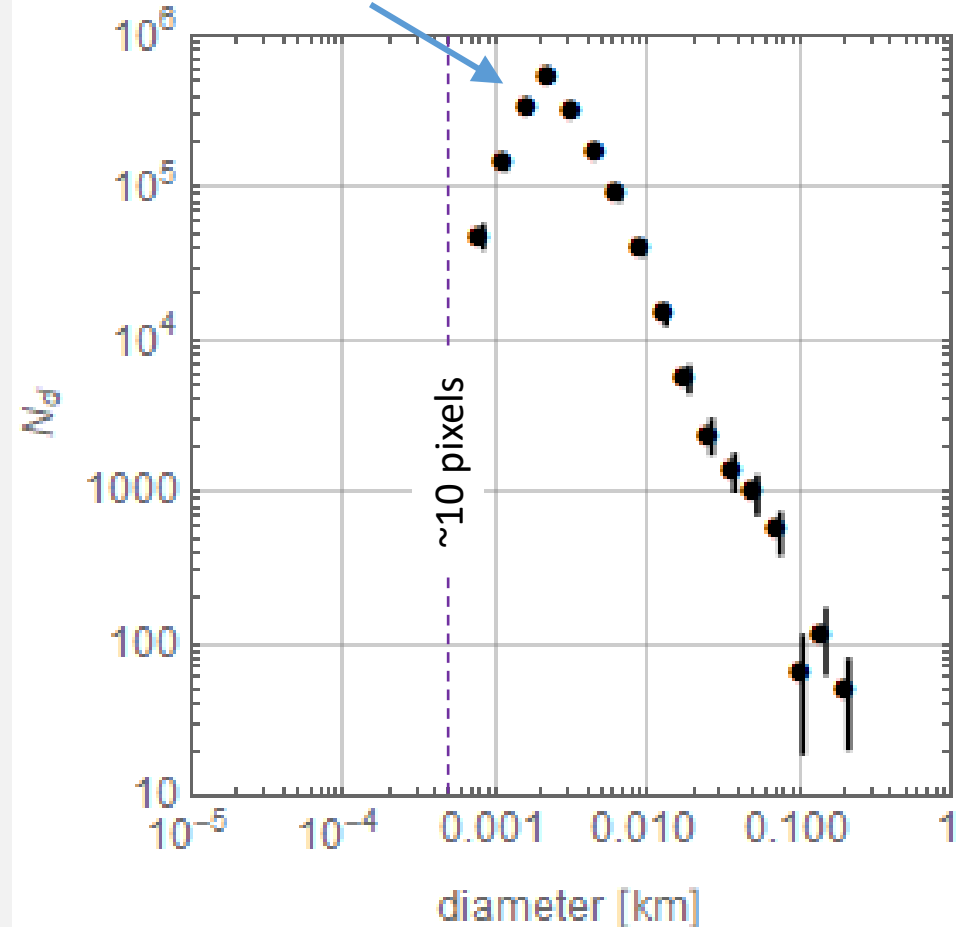


Relative

Completeness . . . or not ?!

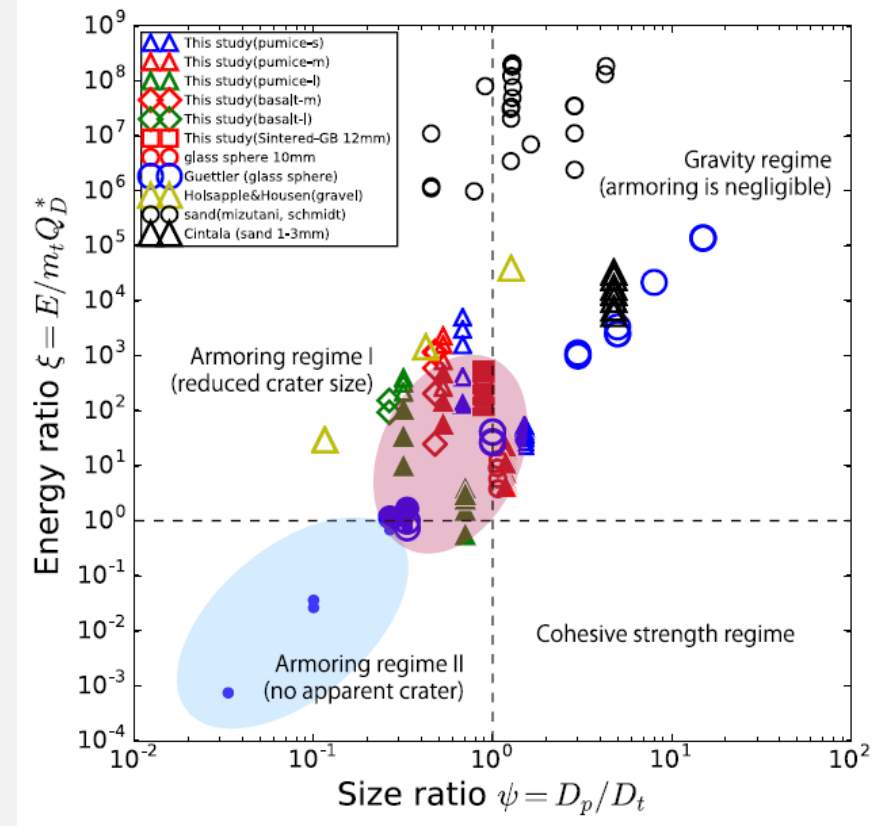
- The rapid fall-off at diameters < 2 m typically would be a sign of the completeness limit (decreasing ability to sample population because of finite image resolution)
- However . . .
- 2 m corresponds to 40 pixels (!) in the detailed survey images, which are typically ~ 5 cm/pix
 - This is well above typical completeness limit values of 5-10 pixels
- **The roll-over is a real observation, and not a completeness-limit effect!**

Rapid drop-off for diameters < 2 m **Global Craters**



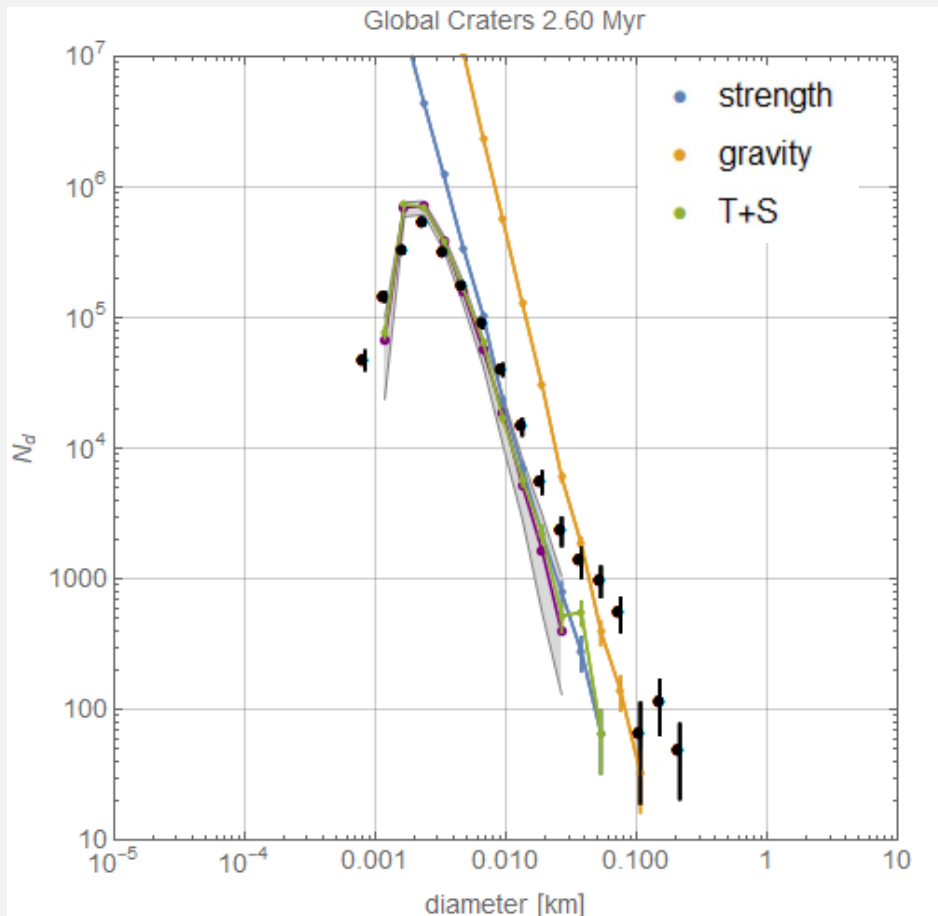
Impact armoring

- Tatsumi and Sugita (2018) [TS2018] conducted a series of experiments that elucidated an “impact armoring” behavior
- Occurs when the impactor size is comparable to the average grain size of the target surface
- They updated standard crater-scaling relationships to include this armoring effect
- We implemented simulations to apply TS2018 scaling to Bennu



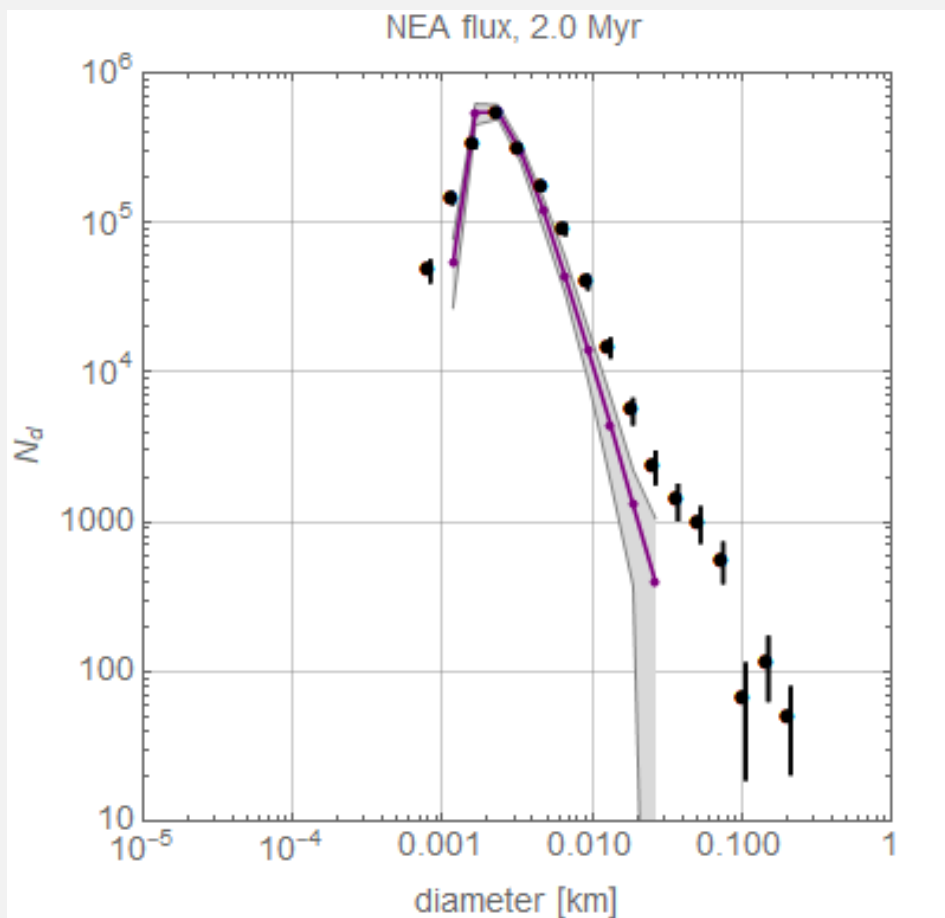
Tatsumi and Sugita (2018), Figure 17

Model results compared with the observations



- **Black** data = Bennu observations
- **Purple** data = median of 100 simulations for 2.6 Myr NEA flux, using TS2018 scaling
- **Gray** band = 99% range of modeled outcomes
- **Orange** = gravity scaling
- **Blue** = strength scaling
- **Green** = a single run of TS2018 scaling
- TS2018 scaling matches the “fish hook” of the differential SFD

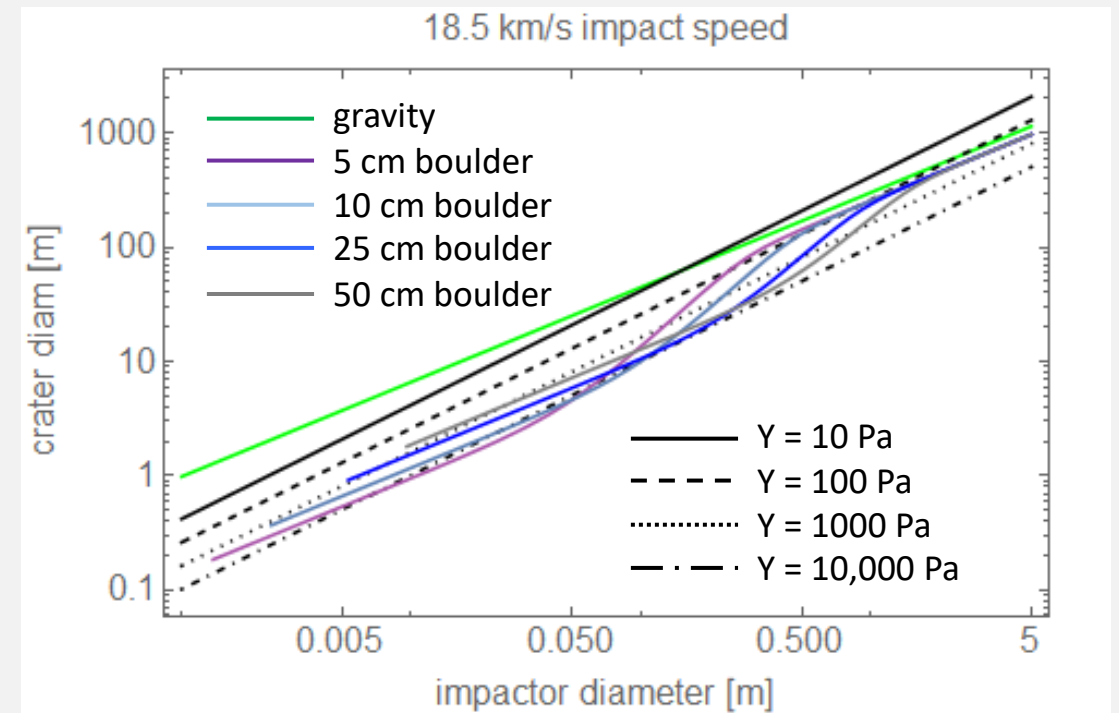
Another look, comparison with TS2018 only



- **Black** data = Bennu observations
- **Purple** data = median of 100 simulations for 2.6 Myr NEA flux, using TS2018 scaling
- **Gray** band = 99% range of modeled outcomes

Armoring is like a strength value in crater-scaling relationships for smaller impacts

- Plot is crater size vs. impactor size for strength, gravity, and TS2018 scaling
 - Single **green** line for gravity
 - **Black** lines are different strength values
 - Other colors are TS2018 for different target boulder sizes
- TS2018 results span a range of 10^3 Pa in strength for small impactor sizes and boulder sizes



Consequences for planetary defense

- On a rubble-pile asteroid the same projectile will have different outcomes depending on the size of the target boulder
 - The impact energy may be transmitted to the bulk object efficiently, or
 - The impact energy may be dissipated largely by disrupting a boulder
- An important consideration for the DART mission: is it possible to determine the size of the boulder(s) that reside at the impact point?
- Any impact-deflection mission should consider the outcome variability introduced by the size of the target boulder

