





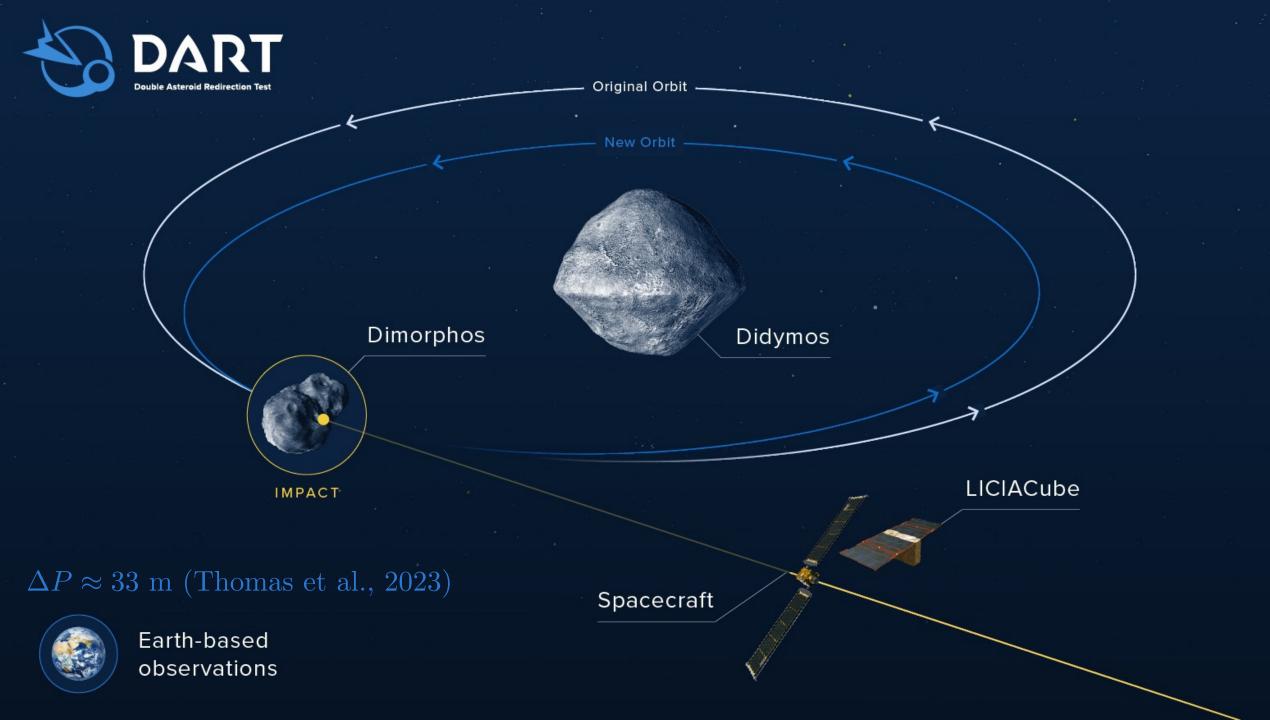
Determination of Dimorphos's change in velocity resulting from the DART kinetic impact

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Cheng, A.F. et al., *Nature*, doi: 10.1038/s41586-023-05878-z

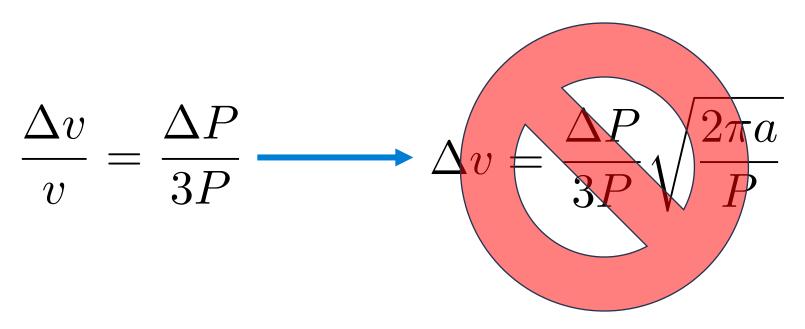


$$\frac{\Delta v}{v} = \frac{\Delta P}{3P}$$

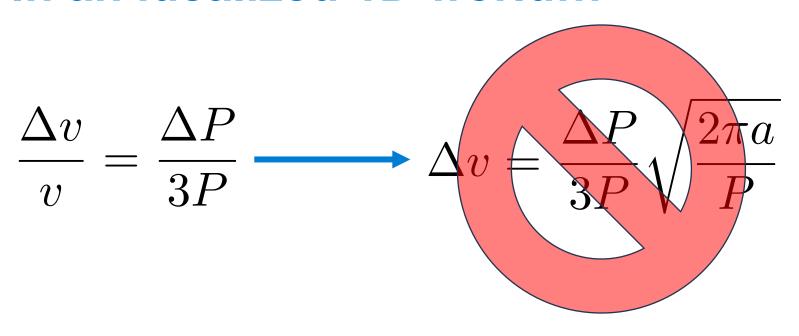


$$\frac{\Delta v}{v} = \frac{\Delta P}{3P} \longrightarrow \Delta v = \frac{\Delta P}{3P} \sqrt{\frac{2\pi a}{P}}$$





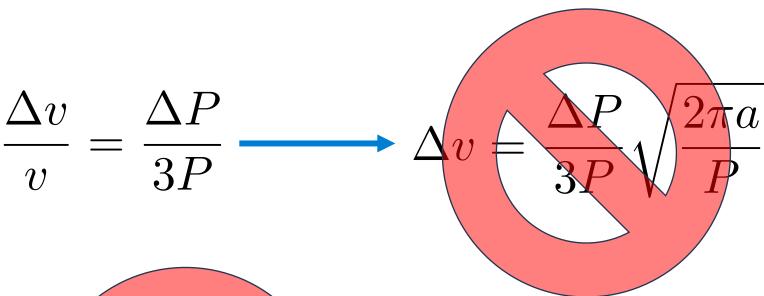
In reality, the dynamics are non-Keplerian. Δv must be computed numerically from ΔP .



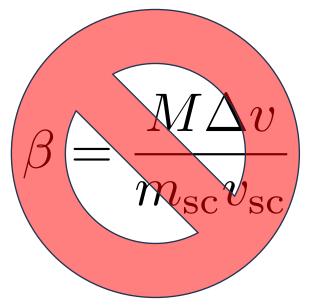
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$$\beta = \frac{M\Delta v}{m_{\rm sc}v_{\rm sc}}$$





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Additionally, the calculation of β must account for 3D impact geometry. (See next talk)

What does ΔP measure?

The change in orbit period is only sensitive to Δv_T , the "along-track" change in orbital velocity

Due to non-Keplerian dynamics & spin-orbit coupling, we use a Full Two-body Problem (F2BP) code to determine Δv_T from ΔP :

- General Use Binary Asteroid Simulator (GUBAS, Davis & Scheeres 2020)
- Assume dynamically-relaxed pre-impact state (circular orbit & tidally-locked)
- Neglect mass loss, time-dependent momentum transfer, etc.
- Rigid body dynamics evolved under 2nd-order mutual gravity
- Didymos and Dimorphos treated as uniform density ellipsoids



Determining Δv_T : Monte Carlo approach coupled to F2BP simulations

Need to account for uncertainties in pre- and postimpact state of Didymos system

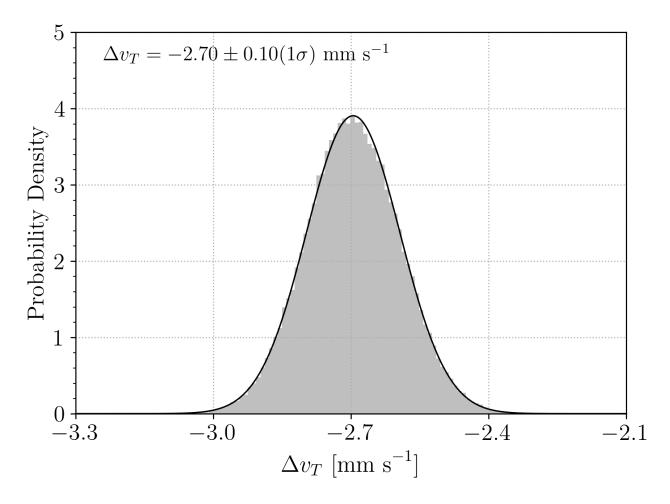
- 1. Sample pre-impact system with uncertainties
- Numerically determine the required mass of primary required to achieve sampled preimpact state
- 3. Numerically determine required Δv_T to achieve sampled post-impact orbit period
- 4. Repeat ~100,000 times

Parameter	Value (1 σ)	Note
Didymos extents (x,y,z) [h]	$851 \pm 15,$ $849 \pm 15,$ 620 ± 15	Sampled uniformly
Dimorphos extents (x,y,z) [m]	$177 \pm 2,$ $174 \pm 4,$ 116 ± 2	Sampled uniformly
Dimorphos density [kg/m³]	2400 ± 300	Sampled uniformly to 3σ
Pre-impact semimajor axis [m]	1206 ± 35	Gaussian
Pre-impact orbit period [h]	11.92148 ± 0.000044	Gaussian
Post-impact orbit period [h]	11.372 ± 0.0055	Gaussian

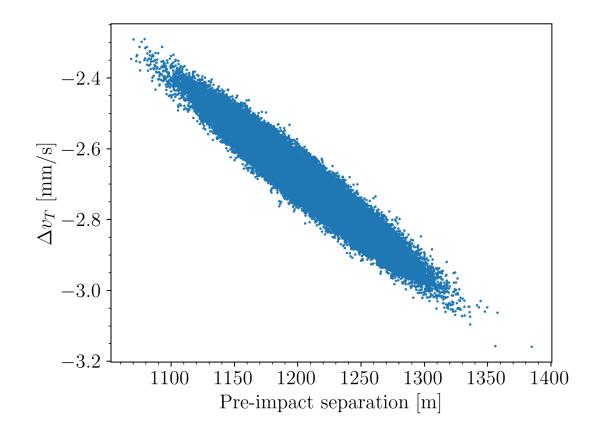
Determination of Δv_T

$$\Delta v_T = -2.7 \pm 0.1(1\sigma) \text{ mm s}^{-1}$$

Sub-mm/s measurement of Dimorphos's change in velocity!

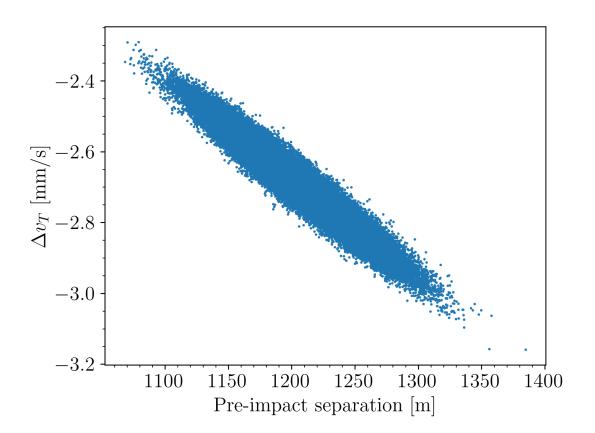


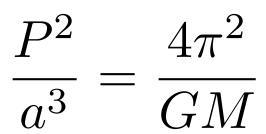
$$a_{\rm pre} = 1206 \pm 35(1\sigma) \text{ m}$$



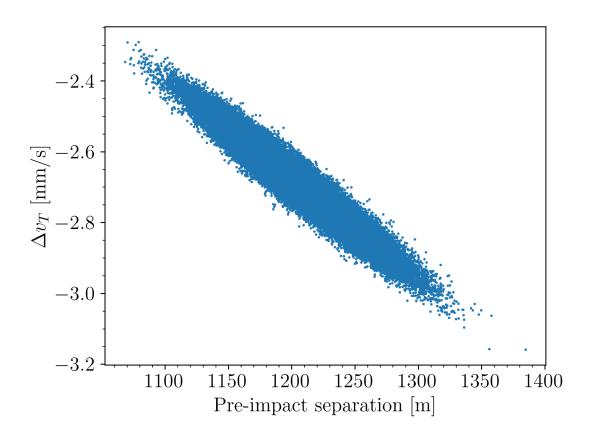


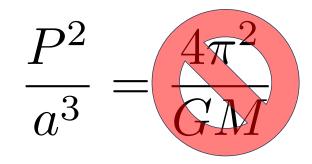
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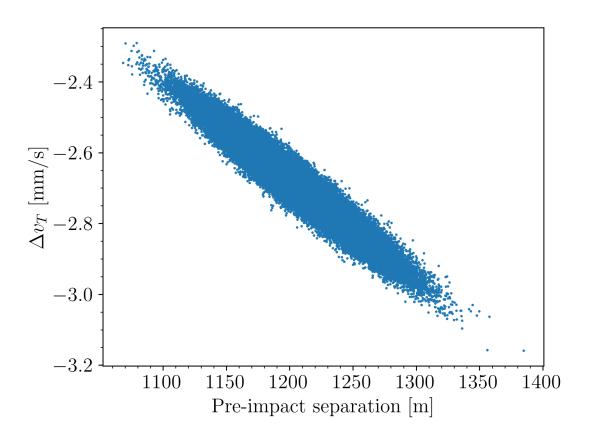
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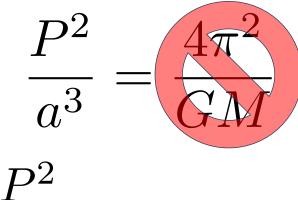






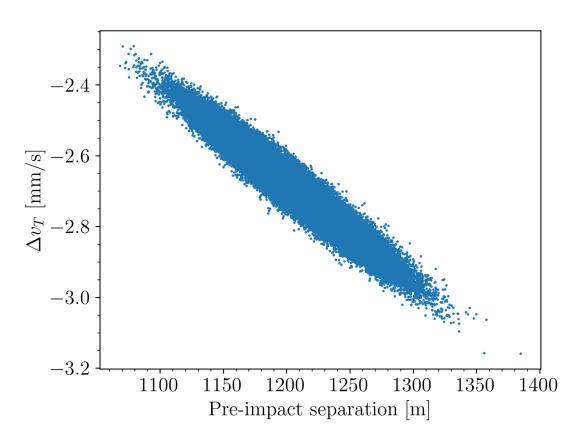
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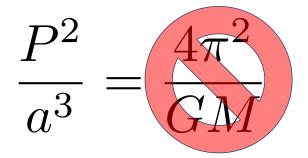




$$\frac{P^2}{a^3} = \text{constant}$$

$$a_{\rm pre} = 1206 \pm 35(1\sigma) \text{ m}$$

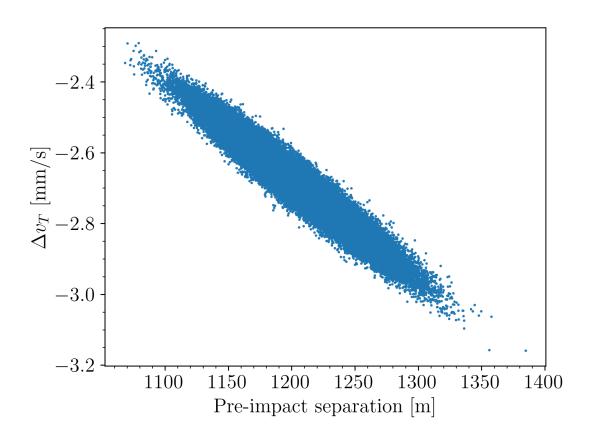




$$\frac{P^2}{a^3} = \text{constant}$$

$$\frac{P_{\text{pre}}^2}{a_{\text{pre}}^3} = \frac{P_{\text{post}}^2}{a_{\text{post}}^3}$$

$$a_{\rm pre} = 1206 \pm 35(1\sigma) \text{ m}$$



$$\frac{P^2}{a^3} = \frac{4\pi^2}{GM}$$

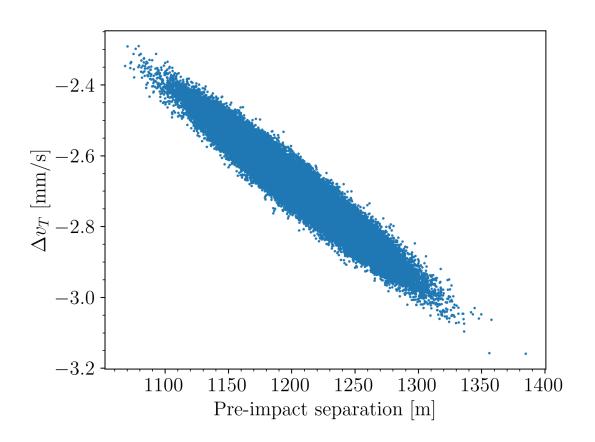
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$$a_{\rm pre} = \left(\frac{P_{\rm pre}}{P_{\rm post}}\right)^{2/3} a_{\rm post}$$

Largest source of uncertainty: pre-impact semimajor axis

$$a_{\rm pre} = 1206 \pm 35(1\sigma) \text{ m}$$



$$\frac{P^2}{a^3} = \frac{4\pi^2}{GM}$$

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$$\frac{P_{\text{pre}}^2}{a_{\text{pre}}^3} = \frac{P_{\text{post}}^2}{a_{\text{post}}^3}$$

Well known



Hera!

$$a_{\rm pre} = \left(\frac{P_{\rm pre}}{P_{\rm post}}\right)^{2/3} a_{\rm post}$$

Conclusions

- DART significantly altered the binary orbit, reducing the orbital velocity by ~2.7 mm/s and the orbit period by ~33 min
- The measurement of Δv_T (and therefore β) can be substantially improved upon Hera's arrival in late 2026

