PDC 2021 Vienna, Austria

Please submit your abstract at https://atpi.eventsair.com/7th-iaa-planetary-defense-conference-2021/abstractsubmission

You may visit https://iaaspace.org/pdc

(please choose one box to be checked) (you may also add a general comment - see end of the page)

Key International and Political Developments Advancements in NEO Discovery New NEO Characterization Results Deflection & Disruption Modeling and Testing Mission & Campaign Design Impact Consequences Disaster Response The Decision to Act Public Education and Communication

Spectroscopic Observations of (65803) Didymos with VLT/X-Shooter

Dominik A. Kiersz^{a,*}, Simon F. Green^b, Andrew S. Rivkin^c, Alan Fitzsimmons^a, Tom Seccull^d

^a Astrophysics Research Centre, Queen's University Belfast, University Road, Belfast, BT7 1NN, UK
^b School of Physical Sciences, The Open University, Walton Hall, Milton Keynes, MK7 6AA, UK
^c Applied Physics Laboratory, Johns Hopkins University, 11101 Johns Hopkins Rd, Laurel, MD 20723, USA
^d Gemini Observatory/NSF's NOIRLab, 670 N. A'ohoku Place, Hilo, HI 96720, USA

Keywords: Asteroids, near-Earth asteroids, Spectroscopy, ESA Hera

We present spectroscopic observations of the binary near-Earth asteroid (65803) Didymos supporting the upcoming NASA DART and ESA Hera missions, which form the first space-based planetary defence collaboration. Reconnaissance efforts are ongoing, attempting to determine the target's physical properties. (65803) Didymos comprises (780 ± 30)m Didymos and (150 ± 30)m Dimorphos [1]. The first optical spectroscopic observations suggested that Didymos is an Xk-type in Bus-DeMeo taxonomy [2], although, subsequent observations indicated an S-type classification [3] with prominent features from olivines and pyroxenes around $\lambda \approx 1 \ \mu m$ and $\lambda \approx 2 \ \mu m$ respectively [4]. Recent observations with WHT/ACAM [5] indicated a shallow 1- μm band, suggesting deficiency in silicates. Inconsistencies exist between previously reported reflectance spectra of Didymos, in particular, their optical slopes and 1- μm band-depths.

Observations of Didymos were performed at an excellent seeing of $\approx 0.6^{\circ}$ at $\lambda = 0.5 \,\mu m$ on the night of 6th April 2019¹ using the intermediate resolution X-Shooter [6] at the ESO 8.2m VLT/UT2. Didymos had visible magnitude V ≈ 20.7 . The data reduction methodology builds upon previous software developed to robustly obtain spectra of faint targets [7]. Our processing adds optimal extraction routines [8], to further increase the S/N.

The combined spectrum of Didymos is shown in Figure 1. As anticipated, Didymos shows a $\lambda \approx 1$ μm silicate absorption indicative of silicate asteroids. Our X-Shooter spectrum has an excellent match

^{*}Corresponding author

Email addresses: dkiersz01@qub.ac.uk (Dominik A. Kiersz), simon.green@open.ac.uk (Simon F. Green),

andy.rivkin@jhuapl.edu (Andrew S. Rivkin), a.fitzsimmons@qub.ac.uk (Alan Fitzsimmons), tseccull@gemini.edu (Tom Seccull)

¹Observing campaign ESO 0103.C-0224

with the ACAM data at the optical part of the spectrum. The 1- μ m band may also exhibit splitting associated with cooler olivine surfaces [9], although it would not be expected given the heliocentric distance of Didymos. Preliminary optical slope measurements between 0.5 and 0.7 μ m, were performed and defined as BR_{Slope} . We report a mean value of $BR_{Slope} = (10.9 \pm 1.0)\%$ per 0.1 μ m, higher than (8.5 \pm 0.9)% per 0.1 μ m from the ACAM data but confirming the red optical slope of the spectrum. Based on our spectrum, an S-type classification is apparent as our near infrared data show both 1- μ m and 2- μ m bands in agreement with previous observations [3]. A complete analysis of the 2- μ m band is not deemed possible at this stage due to extremely low signal beyond 1.80 μ m. We will present a taxonomic and mineralogical analysis of the data, including an investigation of any possible longitudinal variations in composition.



Figure 1: The combined VLT reflectance spectrum of Didymos between 0.35 and 1.80 μm , normalised to unity at $\lambda = 0.55 \ \mu m$. Data from the UVB, VIS and NIR arms of X-Shooter have been scaled via linear regression, assuming reflectance curve continuity between arms and linearity at arm border regions. The spectrum is robustly binned to bin sizes of $\Delta \lambda = 0.002 \ \mu m$ and $\Delta \lambda = 0.004 \ \mu m$ in the UVB/VIS and NIR arms respectively. The ACAM data [5] is plotted for reference, shifted +0.1 in reflectance for clarity.

References

- S. Naidu, L. A. M. Benner, M. Brozovic, M. C. Nolan, S. J. Ostro, J. L. Margot, J. D. Giorgini, T. Hirabayashi, D. J. Scheeres, P. Pravec, et al., Radar observations and a physical model of binary near-Earth asteroid 65803 Didymos, target of the DART mission, Icarus 348 (2020) 113777.
- [2] R. P. Binzel, A. S. Rivkin, J. S. Stuart, A. W. Harris, S. J. Bus, T. H. Burbine, Observed spectral properties of near-Earth objects: results for population distribution, source regions, and space weathering processes, Icarus 170 (2004) 259–294.
- [3] J. de León, J. Licandro, M. Serra-Ricart, N. Pinilla-Alonso, H. Campins, Observations, compositional, and physical characterization of near-Earth and Mars-crosser asteroids from a spectroscopic survey, Astronomy and Astrophysics 517 (2010) 23.
- [4] T. L. Dunn, T. H. Burbine, W. F. Bottke Jr, J. P. Clark, Mineralogies and source regions of near-Earth asteroids, Icarus 222 (2013) 273–282.
- [5] A. Fitzsimmons, S. Green, J. de Leon, S. Lowry, P. Pravec, B. Rozitis, C. Snodgrass, Optical spectroscopy of Didymos during the 2019 Apparition, 2019. AIDA International Workshop.
- [6] J. Vernet, H. Dekker, S. D'Odorico, L. Kaper, P. Kjaergaard, F. Hammer, S. Randich, F. Zerbi, P. J. Groot, J. Hjorth, et al., X-shooter, the new wide band intermediate resolution spectrograph at the ESO Very Large Telescope, Astronomy & Astrophysics 536 (2011) 105.
- [7] T. Seccull, W. C. Fraser, T. H. Puzia, M. E. Brown, F. Schönebeck, 2004 EW95: A Phyllosilicate-bearing Carbonaceous Asteroid in the Kuiper belt, The Astrophysical Journal Letters 855 (2018) L26.
- [8] K. Horne, An optimal extraction algorithm for CCD spectroscopy., Publications of the Astronomical Society of the Pacific 98 (1986) 609.
- [9] L. Moroz, U. Schade, R. Wäsch, Reflectance spectra of olivine-orthopyroxene-bearing assemblages at decreased temperatures: Implications for remote sensing of asteroids, Icarus 147 (2000) 79–93.