



Simulating Planned LICIACube Imagery of DART Impact Ejecta based on Ejecta Dynamics Simulation Output

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This document has been reviewed and determined not to contain
export controlled technical data.

Introduction

- DART to impact surface of Dimorphos, secondary of binary NEA 65803 Didymos
 - at 6.12 – 6.77 km/s ← depends on launch date, per trajectory release D210308-SOC
 - on Sept. 25 - Oct. 1, 2022 ← ditto
 - in “retrograde” sense = nearly opposite Dimorphos’ orbital velocity
- Expected to liberate large quantity of ejecta!
- LICIACube will take images with its 2 cameras during fast flyby, to reveal short-timescale ejecta plume evolution, etc.
- **Here we attempt to simulate how ejecta presents in these planned images**

Note: For what follows we started with older launch date assumption and trajectory, for which:

$$V_{rel} = 7.205 \text{ km/s}$$

(adjusted) impact epoch = 02-OCT-2022 21:24:40 TDB

DART mass at impact, $m = 607.9 \text{ kg}$

Binary Dynamics Simulation

F2BP propagation of underlying motion of binary components

- Setup consistent with DRA at the time (v2.26)
- Propagated using GUBAS* pre- & post-impact
(* Davis & Scheeres, 2020; <https://github.com/alex-b-davis/gubas>)
- Impulsive state change at impact epoch consistent with $\beta=2$ and impact geometry:

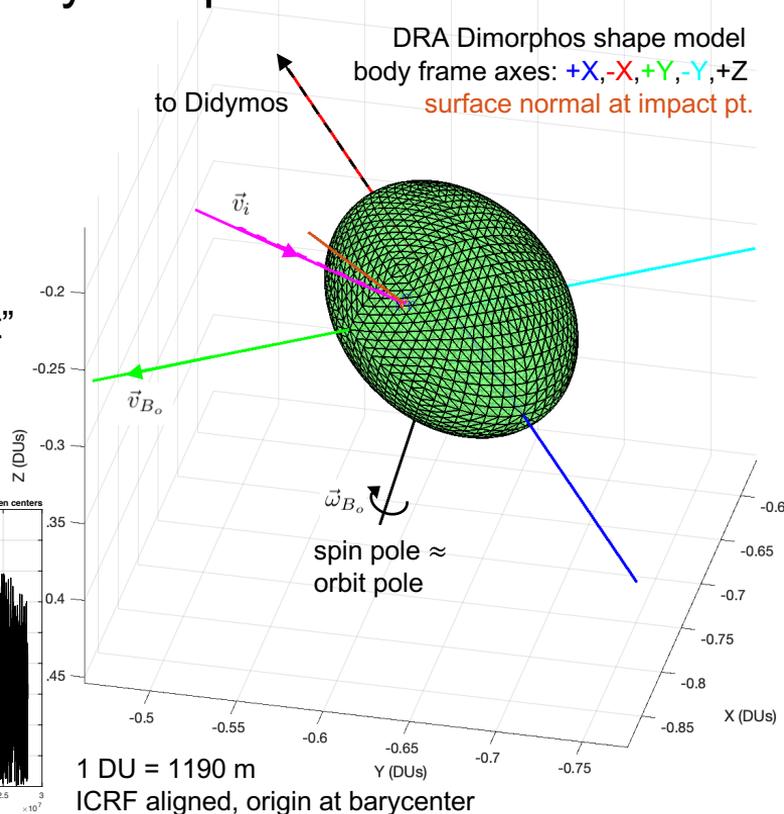
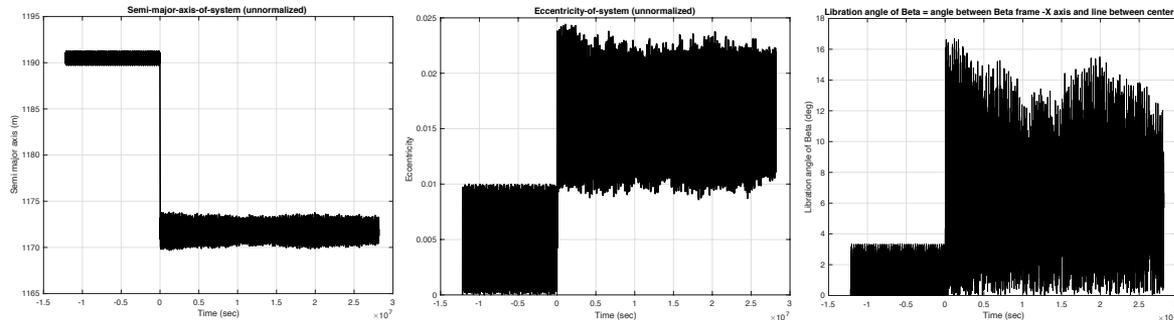
Impact point offset 22.5 m from thru COF, toward “local west”

$$\vec{v}_B = \frac{\beta m_i \vec{v}_i + m_B \vec{v}_{B_o}}{m_i + m_B}$$

$$\vec{\omega}_B = [A_B (I_B + \Delta I_B) A_B^T]^{-1} (\beta m_i (\vec{\rho}_i \times \vec{v}_i) + A_B I_B A_B^T \vec{\omega}_{B_o})$$

$$\vec{v}_B \approx \frac{\beta m_i}{m_B} \vec{v}_i + \vec{v}_{B_o}$$

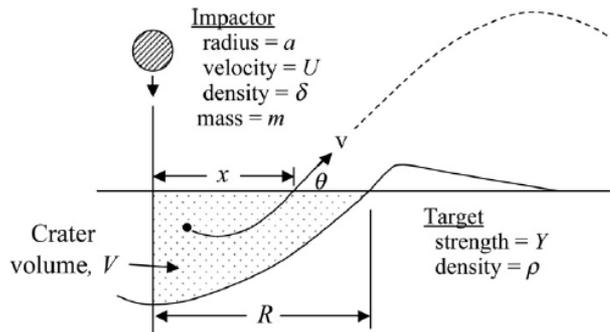
$$\vec{\omega}_B \approx [A_B I_B A_B^T]^{-1} (\beta m_i (\vec{\rho}_i \times \vec{v}_i)) + \vec{\omega}_{B_o}$$



Ejecta Initialization

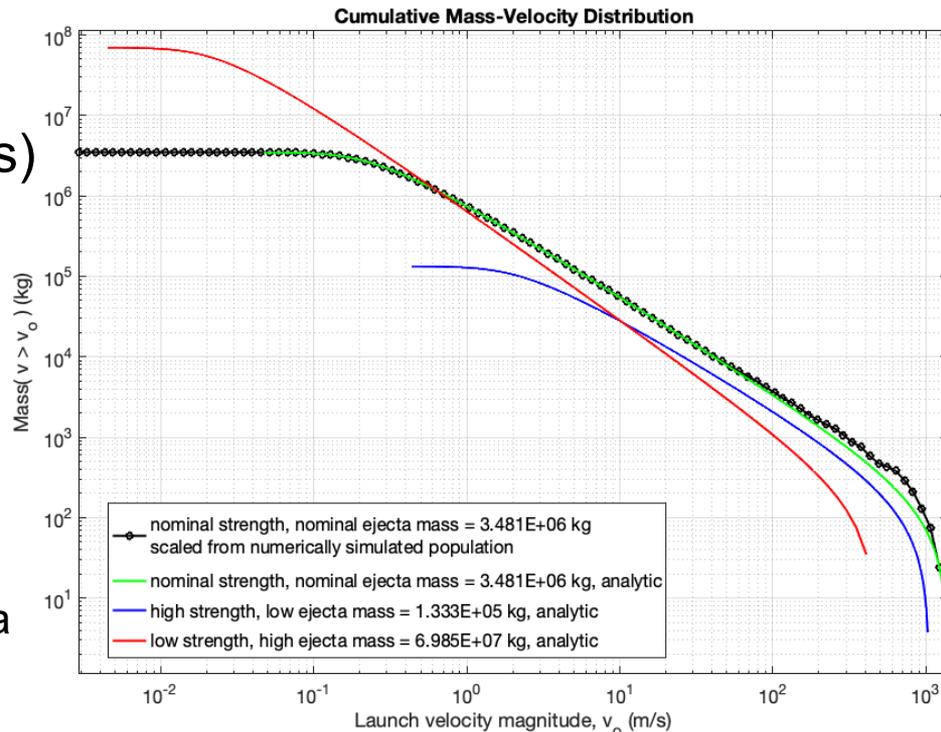
Using Crater Scaling Equations (CSEs)

K.R. Housen, K.A. Holsapple / Icarus 211 (2011) 856-875



- Explored bounding low, high, & nominal ejecta mass cases
- For what follows, we use nominal mass case:

Dimorphos surface bulk density, $\rho = 2111 \text{ kg/m}^3$
 assumed surface cohesive strength, $Y = 100 \text{ Pa}$
 assumed surface porosity = 35%
 Dimorphos centroid to impact point vector in inertial frame (ICRF) = [-63.68 17.39 38.41] m
 surface acceleration, g (net of gravity & centripetal) = $5.927\text{E-}05 \text{ m/s}^2$
 DART equivalent radius, $a = 0.5 \text{ m}$
 DART mass, $m = 607.9 \text{ kg}$
 DART bulk density, $\delta = 1161 \text{ kg/m}^3$
 $V_{rel} = U = 7.205 \text{ km/s}$

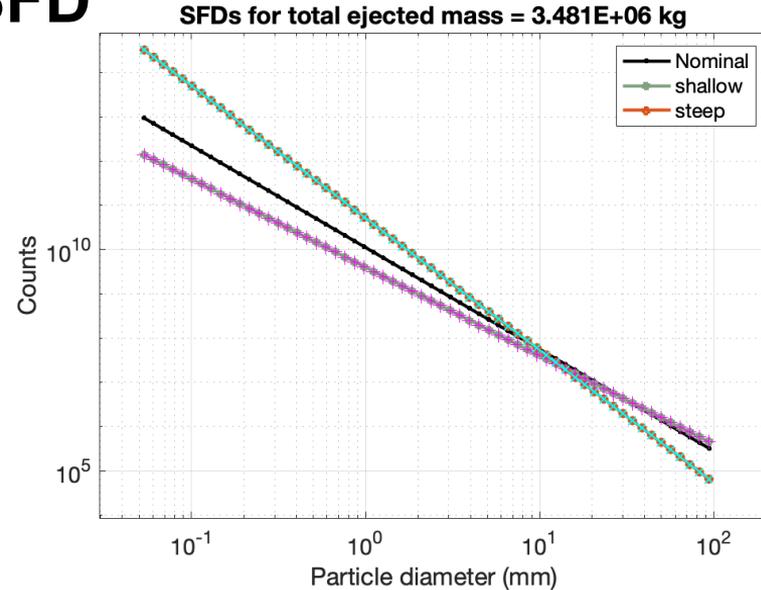


$v = 0.4$
 $\mu = 0.41$
 $C_1 = 0.55$
 $k = 0.3$
 $H_1 = 0.59$
 $H_2 = 0.39$
 $n_1 = 1.2$
 $n_2 = 1.0$
 $p = 0.3$

crater duration = ~ 137.84 seconds
 final crater radius, $R = \sim 17.65 \text{ m}$
 total ejected mass = $3.481\text{E}+06 \text{ kg}$
 max ejecta launch speed (surface relative) = 1.403 km/s
 min ejecta launch speed (surface relative) = $\sim 0 \text{ m/s}$
 ejecta particle grain density = Dimorphos surface bulk density
 (i.e. gross assumption that all porosity is micro-porosity!)

Ejecta Dynamics Propagation & cSFD

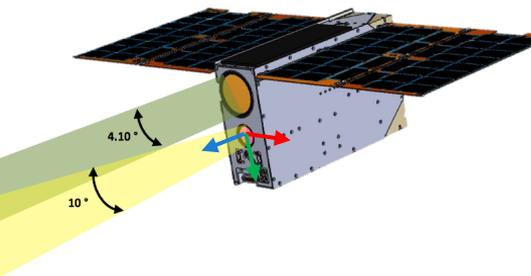
- Initial states translated to barycentric inertial frame
- RF3BP: collision-less, no inter-particle interaction
- Active force models:
 - Gravity of binary components, polyhedral, turned off >1000 km from barycenter
 - SRP (cannonball, inc. body shadowing)
 - Differential solar gravity (“solar tides”)
- Ejecta flagged as “escaped” once crossing plane 200 km anti-sunward (+R in RTN frame) with $E_{2BP} > 0$
- Ejecta size range: 0.05–100 mm, 50 μm – 0.1 m
- Cumulative Size Frequency Distribution (cSFD) power-law slopes:
 - “shallow” = -2.0
 - nominal = -2.3017 ← Itokawa heritage
 - “steep” = -2.99



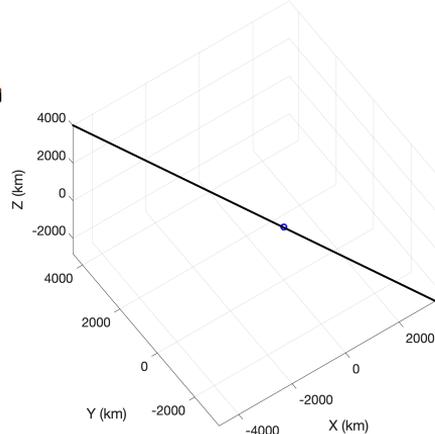
- For what follows, we use nominal slope
- $N \geq 1\text{e}6$ particles propagated to 90 days duration, or to sooner return impact, transfer impact, or escape
- Full population scaled from propagated population using particle multipliers

LICIACube Observation Plan

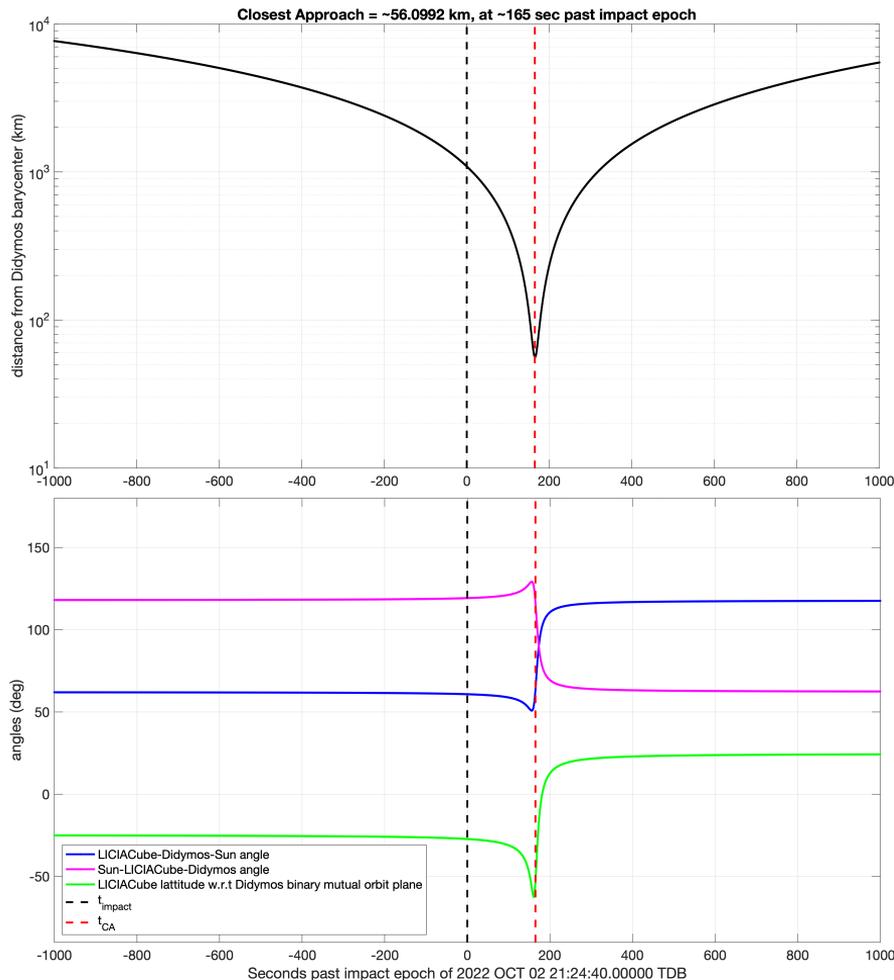
Fly-by geometry & pointing



LICIACube trajectory w.r.t. Didymos barycenter (blue) (ICRF frame)



- PL2's long pixel dimension aligned to red (\hat{y}_{LBF}) unit vector
- Boresight (blue, \hat{z}_{LBF}) assumed maintained pointed from current LICIACube position to centroid of target (Dimorphos)



LICIACube Observation Plan

Camera specs. & image capture times

Property	PL1 = LEIA	PL2 = LUKE
focal length	393 mm	70 mm
f-number	5.2	5
diag. FOV	$\pm 2.05^\circ$	$\pm 5^\circ$
iFOV	25 μ rad	75 μ rad
pixel dims	2048 x 2048	2054 x 1090
color filters	Panchrom. (400 – 900 nm)	RGB
pixel bit depth	12 bit	8, 10 bit
integration time range	0.1 ms – seconds	--
integration time granularity	0.1 ms	--
chosen integration times	0.3, 0.6, 0.9 ms	0.3, 0.6, 0.9 ms
Frame rate	up to 7 fps	up to 5 fps
Image size	~6.3 MB/image	~2.8 MB/image (10 bit)

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For impact epoch 02-OCT-2022 21:24:40.0000 TDB
```

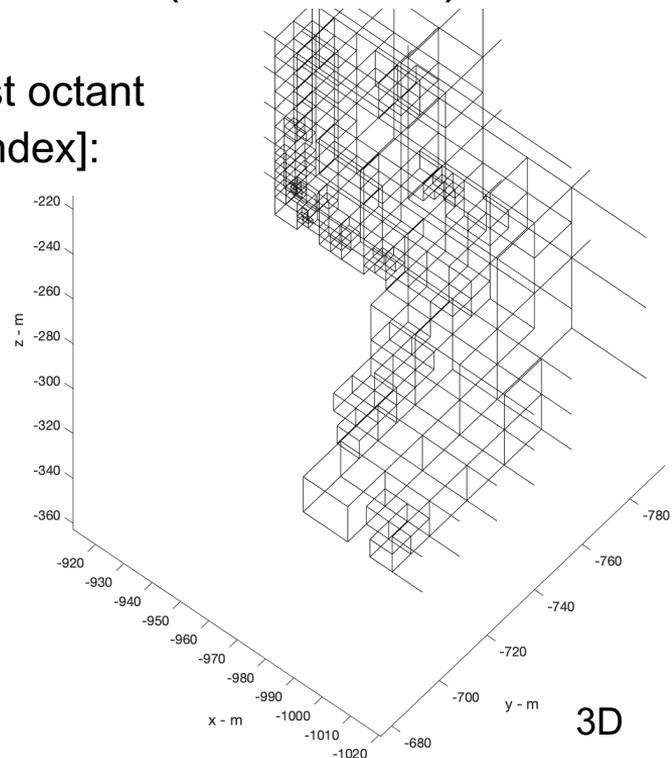
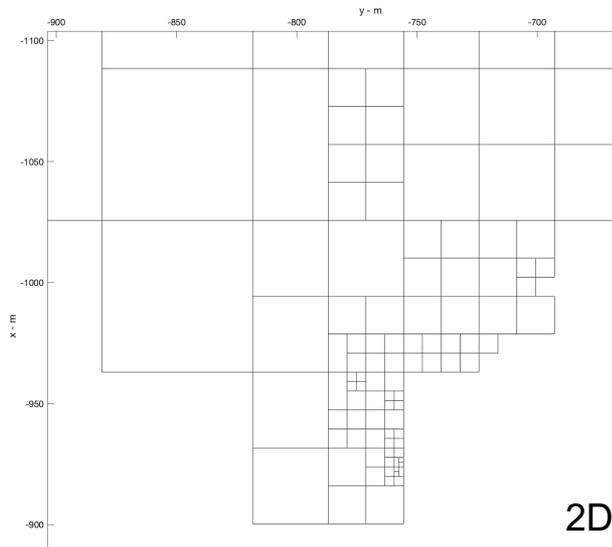
PL1: epoch		w.r.t. impact (s)	w.r.t. C.A. (s)
02-OCT-2022	21:23:54.4100 TDB	-45.5900	-211.0000
02-OCT-2022	21:24:03.4100 TDB	-36.5900	-202.0000
02-OCT-2022	21:24:12.4100 TDB	-27.5900	-193.0000
02-OCT-2022	21:24:21.4100 TDB	-18.5900	-184.0000
02-OCT-2022	21:24:30.4100 TDB	-9.5900	-175.0000
02-OCT-2022	21:24:39.4100 TDB	-0.5900	-166.0000
02-OCT-2022	21:24:48.4100 TDB	8.4100	-157.0000
02-OCT-2022	21:24:57.4100 TDB	17.4100	-148.0000
02-OCT-2022	21:25:05.4100 TDB	25.4100	-140.0000
02-OCT-2022	21:25:11.4100 TDB	31.4100	-134.0000
02-OCT-2022	21:25:17.4100 TDB	37.4100	-128.0000
02-OCT-2022	21:25:23.4100 TDB	43.4100	-122.0000
02-OCT-2022	21:25:29.4100 TDB	49.4100	-116.0000
02-OCT-2022	21:25:35.4100 TDB	55.4100	-110.0000
02-OCT-2022	21:25:41.4100 TDB	61.4100	-104.0000
02-OCT-2022	21:25:47.4100 TDB	67.4100	-98.0000
02-OCT-2022	21:25:53.4100 TDB	73.4100	-92.0000
02-OCT-2022	21:25:59.4100 TDB	79.4100	-86.0000
02-OCT-2022	21:26:05.4100 TDB	85.4100	-80.0000
02-OCT-2022	21:26:11.4100 TDB	91.4100	-74.0000
02-OCT-2022	21:26:17.4100 TDB	97.4100	-68.0000
02-OCT-2022	21:26:23.4100 TDB	103.4100	-62.0000
02-OCT-2022	21:26:29.4100 TDB	109.4100	-56.0000
02-OCT-2022	21:26:35.4100 TDB	115.4100	-50.0000
02-OCT-2022	21:26:41.4100 TDB	121.4100	-44.0000
02-OCT-2022	21:26:47.4100 TDB	127.4100	-38.0000
02-OCT-2022	21:26:53.4100 TDB	133.4100	-32.0000
02-OCT-2022	21:26:56.4100 TDB	136.4100	-29.0000
02-OCT-2022	21:27:02.9100 TDB	142.9100	-22.5000
02-OCT-2022	21:27:09.4100 TDB	149.4100	-16.0000
02-OCT-2022	21:27:15.9100 TDB	155.9100	-9.5000
02-OCT-2022	21:27:23.4100 TDB	163.4100	-2.0000
02-OCT-2022	21:27:24.1600 TDB	164.1600	-1.2500
02-OCT-2022	21:27:24.9100 TDB	164.9100	-0.5000
02-OCT-2022	21:27:25.6600 TDB	165.6600	0.2500
02-OCT-2022	21:27:26.1100 TDB	166.1100	0.7000
02-OCT-2022	21:27:27.7300 TDB	167.7300	2.3200
02-OCT-2022	21:27:29.3500 TDB	169.3500	3.9400
02-OCT-2022	21:27:30.9700 TDB	170.9700	5.5600
02-OCT-2022	21:27:32.5900 TDB	172.5900	7.1800
02-OCT-2022	21:27:34.2100 TDB	174.2100	8.8000
02-OCT-2022	21:28:25.4100 TDB	225.4100	60.0000
02-OCT-2022	21:29:25.4100 TDB	285.4100	120.0000
02-OCT-2022	21:30:25.4100 TDB	345.4100	180.0000
02-OCT-2022	21:31:25.4100 TDB	405.4100	240.0000
02-OCT-2022	21:32:25.4100 TDB	465.4100	300.0000
02-OCT-2022	21:33:25.4100 TDB	525.4100	360.0000
02-OCT-2022	21:34:25.4100 TDB	585.4100	420.0000
02-OCT-2022	21:35:25.4100 TDB	645.4100	480.0000
02-OCT-2022	21:36:25.4100 TDB	705.4100	540.0000
02-OCT-2022	21:37:25.4100 TDB	765.4100	600.0000

PL2: epoch		w.r.t. impact (s)	w.r.t. C.A. (s)
02-OCT-2022	21:25:05.4100 TDB	25.4100	-140.0000
02-OCT-2022	21:25:11.4100 TDB	31.4100	-134.0000
02-OCT-2022	21:25:17.4100 TDB	37.4100	-128.0000
02-OCT-2022	21:25:23.4100 TDB	43.4100	-122.0000
02-OCT-2022	21:25:29.4100 TDB	49.4100	-116.0000
02-OCT-2022	21:25:35.4100 TDB	55.4100	-110.0000
02-OCT-2022	21:25:41.4100 TDB	61.4100	-104.0000
02-OCT-2022	21:25:47.4100 TDB	67.4100	-98.0000
02-OCT-2022	21:25:53.4100 TDB	73.4100	-92.0000

Interface to Image Simulation

3D spatial number density in gridded interface data files (“cube files”)

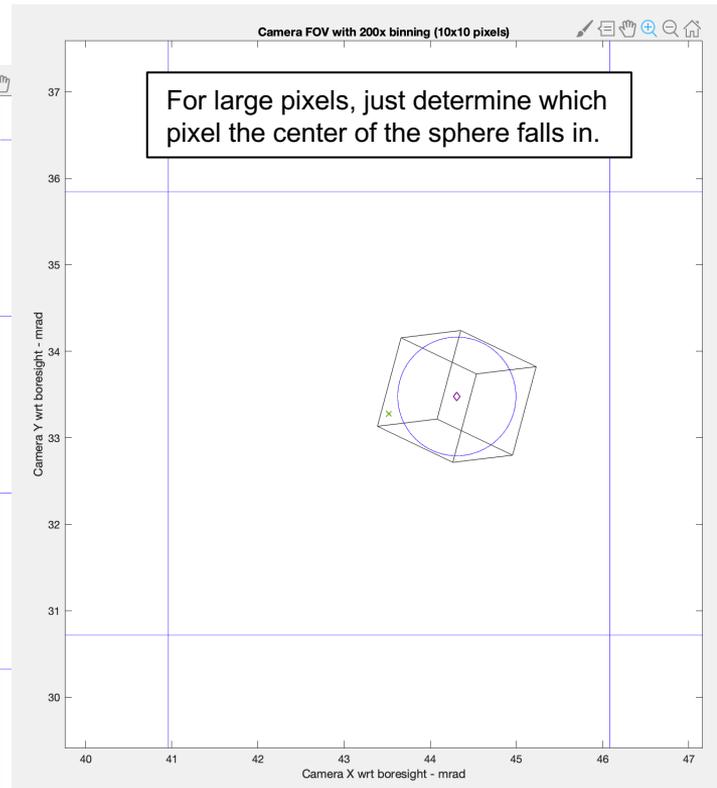
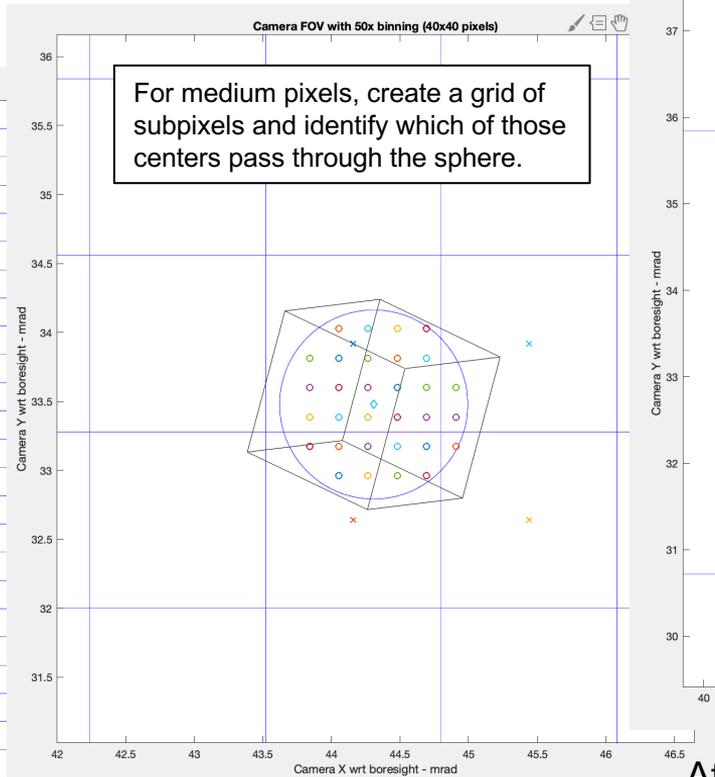
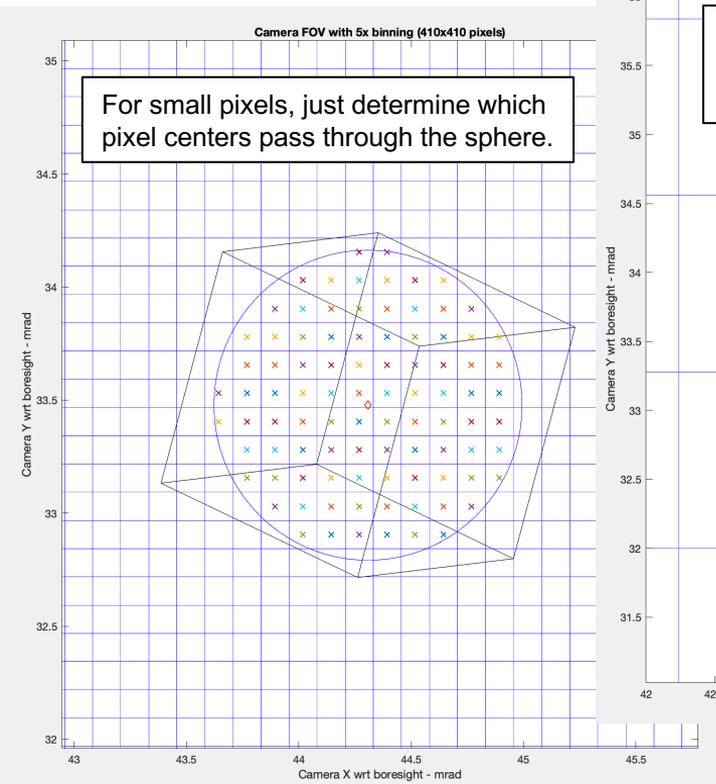
- One such file at each unique image capture time
- Define grid of base cubes with origin s.t. all ejecta lies in first octant
- For each unique [particle size bin + non-empty base cube index]:
 - Enter recursive subdivision algorithm that tests sub-cubes formed by bisection along each cardinal axis
 - Don't do another level of subdivision if that would result in # of actually simulated particles in non-empty sub-cubes $< n$
 - # of actually simulated particles in non-empty sub-cubes scaled to full pop. using particle multipliers, divided by volume



Replace cube with equivalent volume sphere when figuring which pixels it falls on.

Mapping Cubes onto Pixels

Three cases:



After “filling” the pixels by looping over all cubes, you can create the image.

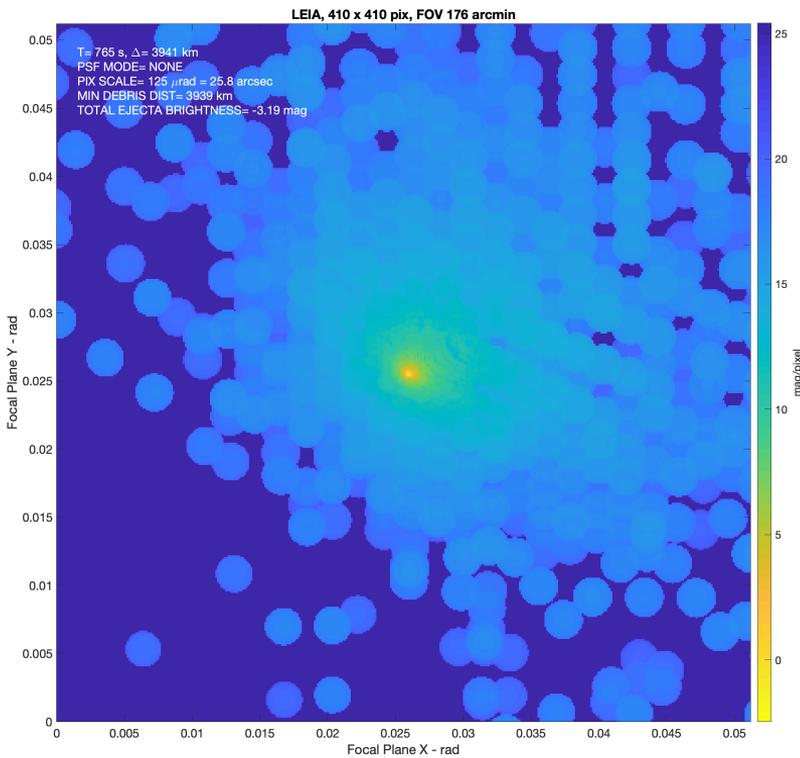
Photometric Model

Converting # of particles per pixel, for each size bin, to brightness:

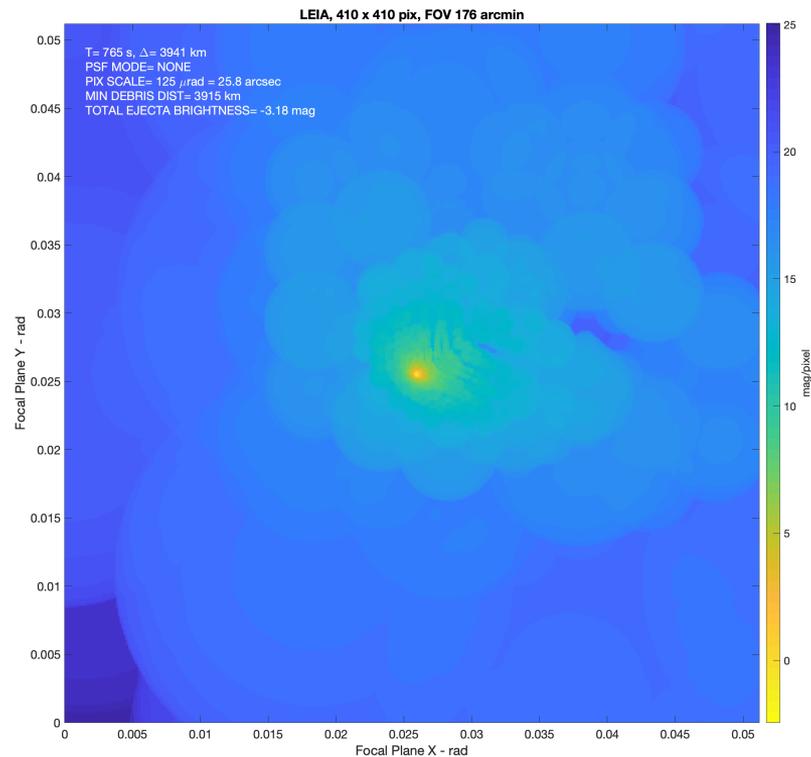
- Recall our particle size range is 0.05 – 100 mm
- Follow Hergenrother's model from Bennu particles (1 – 50 mm)
- $V = H + S \phi + 5 \log_{10} (R\Delta)$
 - H: absolute mag
 - S: phase slope = 0.013 mag/deg (Hergenrother, 2020)
 - ϕ : phase angle
 - R: Heliocentric distance (in au)
 - Δ : Observer distance (in au)

Simulated LEIA (PL1) Images

Default >100 base cubes per axis (fine)
 # of levels of subdivision limited to 10



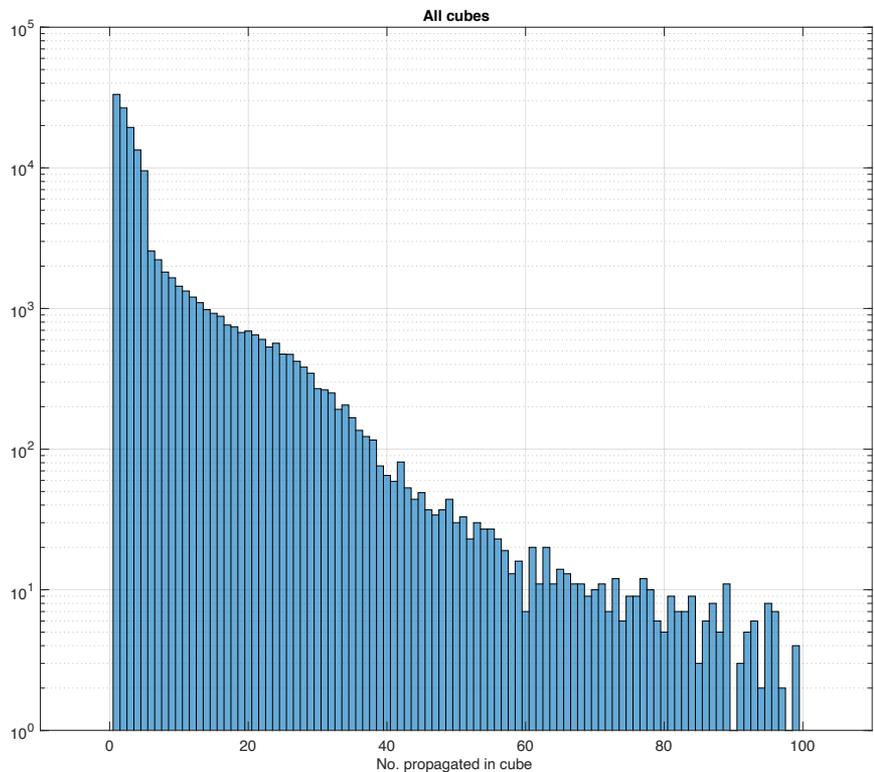
Default >10 base cubes per axis (coarse)
 # of levels of subdivision limited to 13



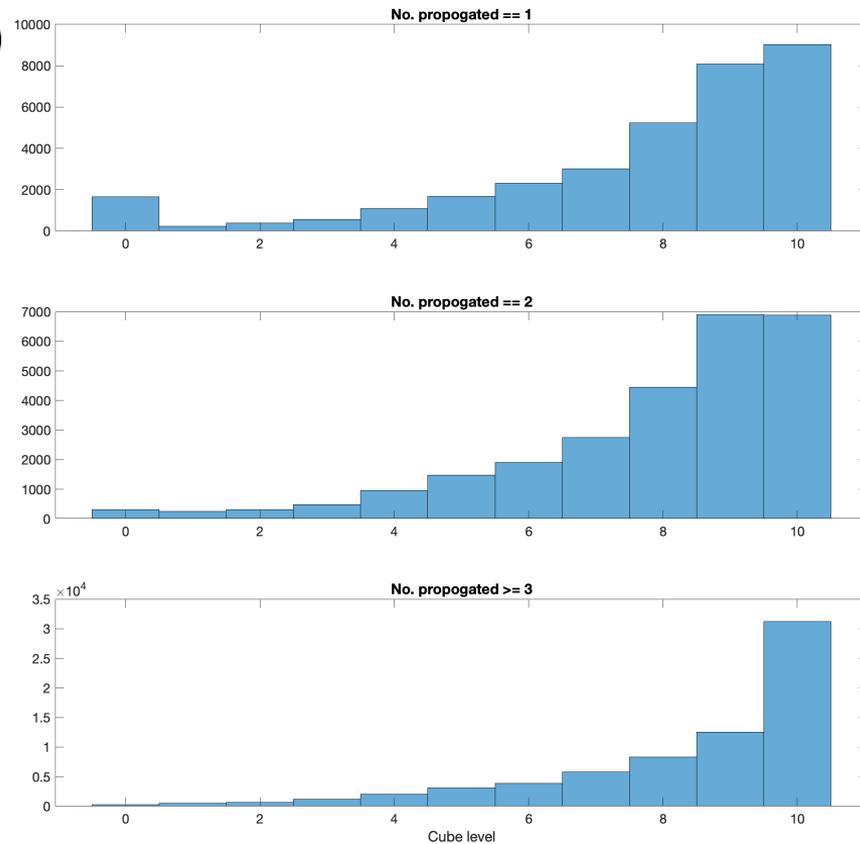
Backup Slides

Results of Subdivision

Default >100 base cubes per axis (fine)

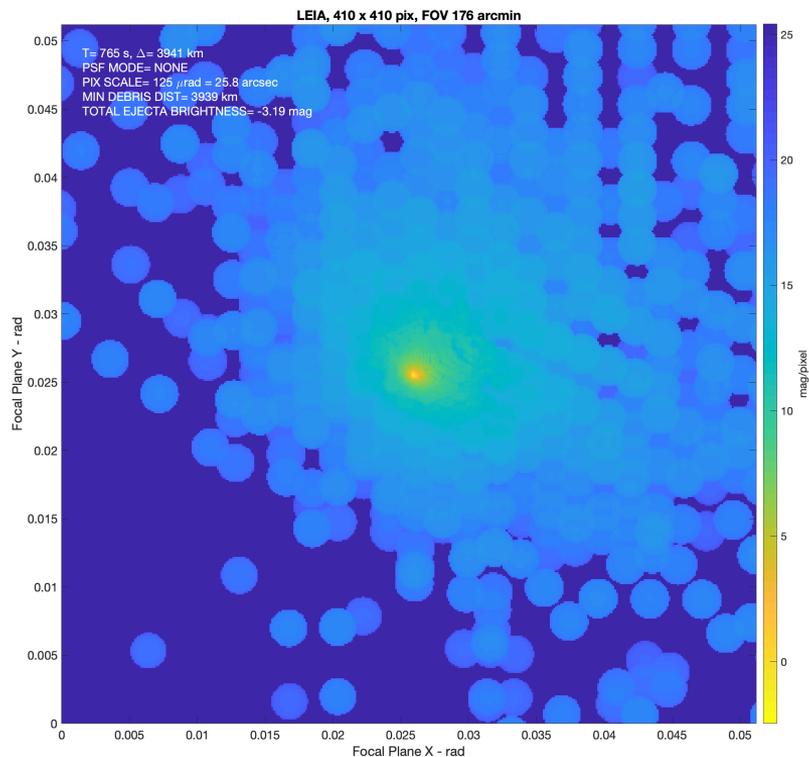


of levels of subdivision limited to 10



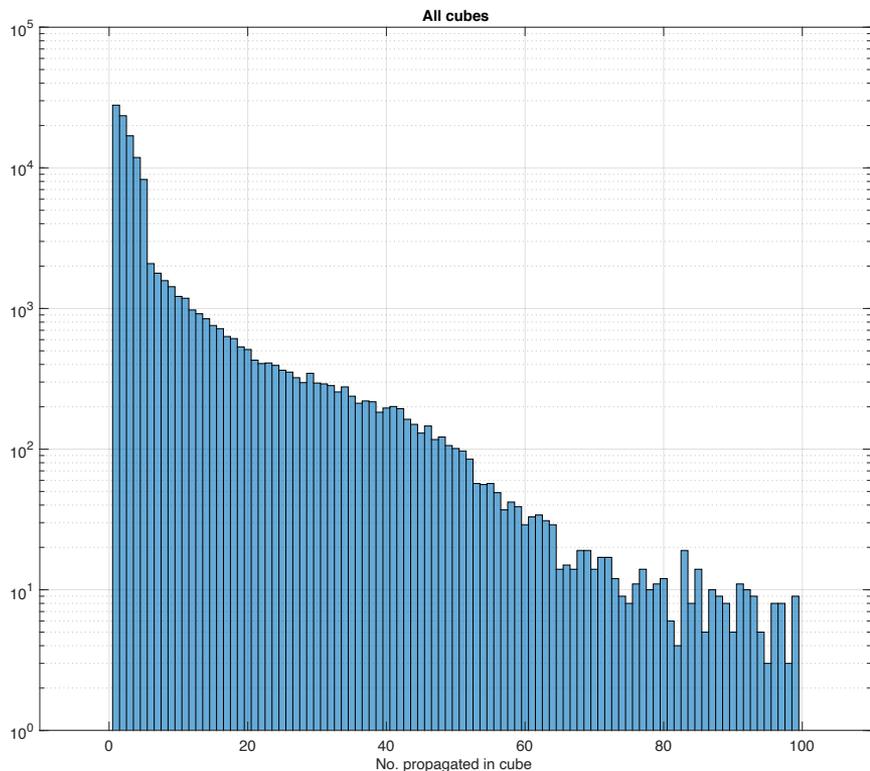
Simulated LEIA (PL1) Image

Default >100 base cubes per axis (fine)

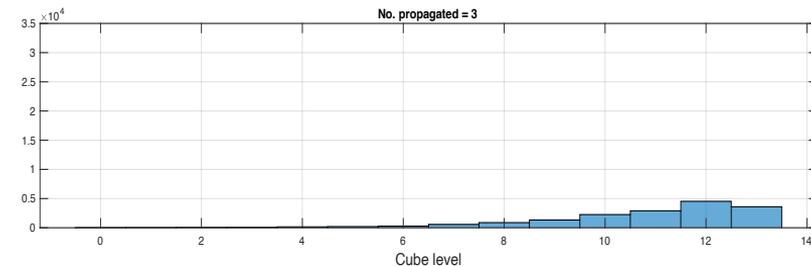
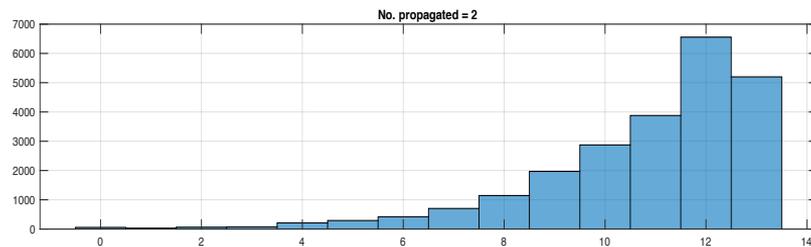
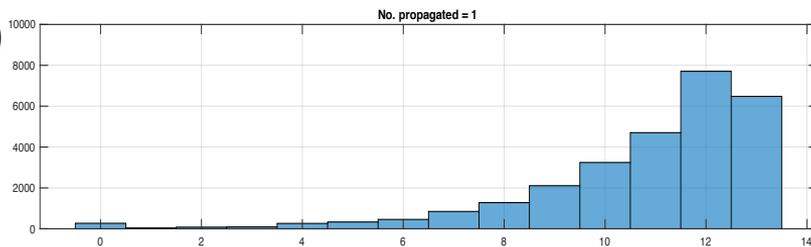


Results of Subdivision

Default >10 base cubes per axis (coarse)

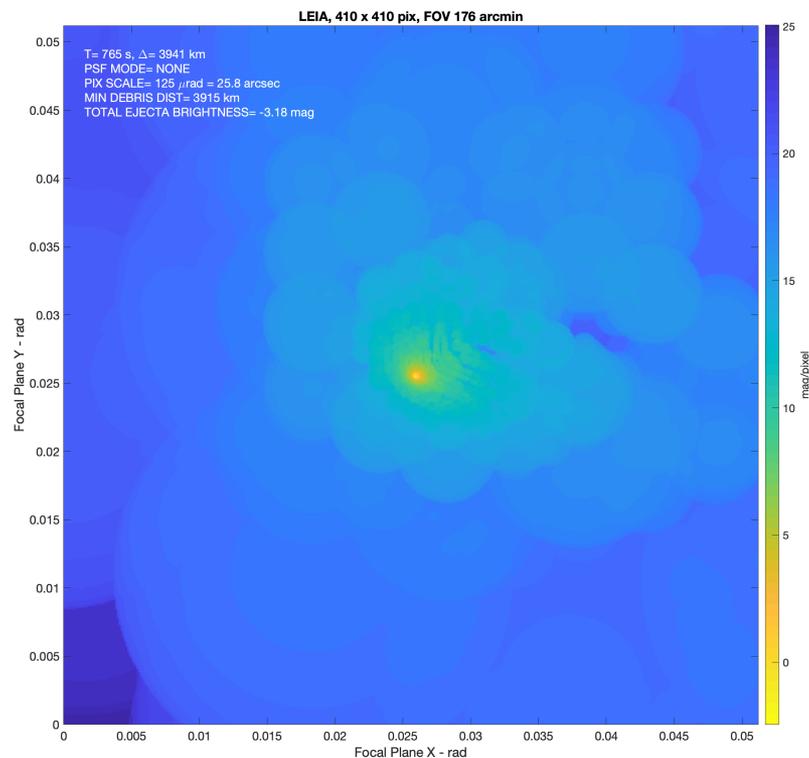


of levels of subdivision limited to 13



Simulated LEIA (PL1) Image

Default >10 base cubes per axis (coarse)





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