Dielectrophoretic-driven convection in the spherical shell experiment AtmoFlow

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Laboratory studies of large-scale flows in planetary atmospheres and interiors are typically conducted in cylindrical or spherical shell systems that preserve the essential physical dynamics. AtmoFlow is one such experiment, consisting of concentric spherical shells that simulate a planetary atmosphere with equatorial heating and polar cooling. Planetary rotation is replicated by shell rotation, while a central electric field generates an artificial gravity-like force. To eliminate buoyancy effects from natural convection, the experiment will be conducted aboard the International Space Station (ISS), planned for 2026. The central force, as investigated by Futterer et al. [1], is induced by a dielectrophoretic effect that decays proportionally to $1/r^5$. The arising thermo-electrohydrodynamic (TEHD) convection was previously studied in the Geophysical Flow Simulator (GeoFlow), which operated aboard the ISS from 2008 to 2017 to model geophysical flow fields. The present study builds upon the work of Futterer et al. [1], focusing on the AtmoFlow project with an aspect ratio of $\eta = 0.7$ to investigate the basic flow.

Dielectrophoretic thermo-electrohydrodynamic (TEHD) convection arises in a non-isothermal fluid subjected to an electric tension, resulting in an intrinsic or externally induced inhomogeneous electric field, such as one generated by curvature. The resulting central force field is governed by Gauss's law, producing convective flow analogous to Rayleigh-Bénard convection. By employing the continuity, momentum, and energy equations, a forcing parameter analogous to the Rayleigh number can be derived by

$$Ra_{\rm E} = \frac{\varepsilon_0 \varepsilon_{\rm r} \gamma_{\rm e} V_0^2}{2\rho_0 v \kappa}, \quad \text{with} \quad \gamma_{\rm e} = e \Delta T$$
 (1)

where ε_0 is the vacuum electric permittivity, ε_r the relative electric permittivity, γ_e the thermoelectric parameter, V_0 the applied electric tension, ρ_0 the reference density, v the kinematic viscosity and κ the thermal diffusivity.

In this study, we extend the parameter space of Futterer et al. [1] to investigate the flow fields within the parameter range of the AtmoFlow experiment, spanning 1×10^5 to 2×10^7 . Three distinct flow regimes have ben identified: (1) steady-state plume structures with quadratic-shaped mode-6 patterns, (2) a transient periodic regime characterized by the emergence of sheet-like structures and mode reduction and (3) an irregular regime identified by structural variations and fluctuations in mode amplitude, forming structural clusters that diminish with increasing $Ra_{\rm E}$.

A quantitative analysis by integral values shows that the Nusselt number, Nu, increases with $Ra_{\rm E}$ and $\gamma_{\rm e}$. However, the kinetic energy, $E_{\rm kin}$, does not scale linearly with the forcing, indicating non-linear energy dissipation. Comparisons with Moore and Weiss [2] reveal similar but not identical Nu trends, while steady-state simulations align well with GeoFlow results when normalized by the critical Rayleigh number. Furthermore, regime transitions exhibit margins comparable to those reported by Futterer et al. [1], with slight differences arising from variations in the definitions of transient and irregular regimes and the inclusion of a temperature-sensitive Gauss's equation in this study to account for thermal feedback. The relationship between Nu and $E_{\rm kin}$ confirms independence from the Prandtl number and aspect ratio, consistent with findings by Futterer et al. [3]. This study offers novel insights into the dynamics of TEHD convection in spherical shells, its heat transport mechanisms, and their application to the experimental study of geophysical flow phenomena in small-scale laboratory experiments.

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